

# Harmonisation of microwave humidity sounder time series for climate applications

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### Motivation: microwave sensor differences

Channel 183.31 ±1 GHz



- Large offsets
- Some differences change with time

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# Inter-satellite re-calibration

Motivation:

- long term Climate Data Records (CDR's) need to combine observations from several sensors
- **but** there are unexpected **sensor to sensor differences** even after careful calibration Methods:
- <u>bias</u> and trend <u>correction</u> is applied after calibration on >Level-1b data
- <u>homogenisation</u> does not take differences in Spectral Response Function (SRF) into account
- <u>harmonisation</u> (inter-satellite re-calibration)
  - based on measurement equations
  - given SRF differences are taken into account
  - uses fundamental measurements, i.e. counts in level-1a data
  - considers all uncertainties, **metrological** based approach
  - uses simultaneous overpasses or **match-ups**

# Match-ups



Sensors on satellites looking

- at the same **area** on earth
- at the same time
- from the same distance
- with the same viewing angle
- with the same photon **sensitivity** shall measure the same radiance

Here we assume

- largely overlapping area
- small time difference
- small viewing angle differences
- differences in Spectral Response Functions (SRF)

Due to the spatial and temporal variability of the radiances these intervalls add a match-up uncertainty



### Harmonisation

Solution:

- From simultaneous overpasses of satellites match-up data sets are derived.
- Additional 'warm' match-ups of targets varying slowly based on geostationary satellites are used.
- Uncertainty of each fundamental datum is assessed.
- Expected differences in radiance due to different SRF are calculated. Uncertainty of expected differences is assessed.
- Uncertainty in radiance difference due to match-up is assessed. Zero mean difference is assumed.
- A Bayesian based cost function (J) is defined to quantify the sensor to sensor misfit.
- The cost function is minimized by changing parameters of the sensors measurement equation.
- Parameters (x) can be polynomial coefficients and/or calibration coefficients determined pre-launch or in-orbit.
- Certain other parameters (sensor state variables q) of the measurement equation have uncertainties to be considered
- All derivative code generated by Automatic Differentiation (AD)
- Marginalised Error In Variables (MEIV) method:
  - does not optimize sensor state variables (q)
  - substabtially reduced memory and computing time resources
  - takes all uncertainties into account (including correlation)

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### Harmonisation Framework



### Harmonisation Framework MEIV

C<sub>i</sub> is **huge**, handling and inverting a huge matrix is expensive ==> using matrix free algorithms

$$J_{i,j} = \frac{1}{2} (R_i - R_j - K_{ij})^t (C_i + C_j + C_{kij})^{-1} (R_i - R_j - K_{ij})$$

$$(C_i + C_j + C_{kij})^{-1} (R_i - R_j - K_{ij}) \text{ by solving } (C_i + C_j + C_{kij}) x = (R_i - R_j - K_{ij}) \text{ using conjugate gradient solver}$$

$$C_i x = (\frac{\partial S_i}{\partial q_i}) W C_q W^t (\frac{\partial S_i}{\partial q_i})^t x$$

$$(\frac{\partial S_i}{\partial q_i})^t x \text{ matrix free by reverse/adjoint mode Automatic Differentiation}$$

$$\frac{\partial S_i}{\partial q_i} v \text{ matrix free by forward/tangent mode AD}$$

minimization of **J** by limited memory quasi Newton BFGS algorithm using:

 $\frac{\partial J}{\partial x}$  gradient by <u>reverse/adjoint</u> mode AD

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### measurement equations METOP-A,B (MHS)

#### Antenna temperature by Planck law

 $R_{OBCT} = Planck \left( T_{OBCT} \right)$  $R_{DSV} = Planck \left( T_{DSV} \right)$ 

#### Two-point calibration

 $a = \frac{R_{OBCT} - R_{DSV}}{C_{OBCT} - C_{DSV}}$ 

$$R_{ME} = R_{OBCT} + a * (C_E - C_{OBCT}) + u * a^2 * (C_E - C_{DSV}) * (C_E - C_{OBCT})$$

Antenna Pattern Correction (APC)

$$R = \frac{1}{g_E} * \left[ R_{ME} - (1 - g_E) * R_{DSV} \right]$$

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Brightness temperature by inverse Planck law

 $T = Planck^{-1}(R)$ 

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#### R : radiance

Т

- : brightness temperature
- ν : frequence

#### sensor state variables:

- C<sub>DSV</sub> : Deep Space View counts
- C<sub>E</sub> : Earth counts
- COBCT : On Board Calibration Target counts
- u : non-linearity factor
- $g_E$  : portion from Earth (APC)



Planck law

$$R = \frac{(2hc^2 * v^3)}{(e^{((hc/k_b * v)/T)} - 1)}$$

$$= \frac{hc/k_b * v}{\log\left(\frac{2hc^2v^3}{R} + 1\right)}$$

### measurement equations Fengyun-3A (MWHS)

#### Antenna temperature by Planck law

 $R_{OBCT} = Planck(T_{OBCT})$  $R_{DSV} = Planck(T_{DSV})$ 

#### Two-point calibration

$$a = \frac{R_{OBCT} - R_{DSV}}{C_{OBCT} - C_{DSV}}$$
$$R_{ME} = R_{OBCT} + a * (C_E - C_{OBCT})$$

Brightness temperature by inverse Planck law

$$T_{ME} = Planck^{-1}(R_{ME})$$

Non linear correction

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$$T_{nl} = u_1 + u_2 * T_{ME} + u_3 * T_{ME}^2$$

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#### sensor state variables:

- C<sub>DSV</sub> : Deep Space View counts
- C<sub>E</sub> : Earth counts
- COBCT : On Board Calibration Target counts
- $R_{\mbox{\tiny OBCT}}\,$  : On Board Calibration Target radiance
- u<sub>i</sub> : coefficients of APC
- APC<sub>1</sub> : pixel dependent Antenna Pattern Correction



Antenna Pattern Correction

 $T = APC_1 + APC_2 * T_{nl}$ 

### measurement equations Fengyun-3B/C (MWHS/MWHS2)

#### Antenna temperature by Planck law

 $R_{OBCT} = Planck \left( T_{OBCT} \right)$  $R_{DSV} = Planck \left( T_{DSV} \right)$ 

#### Two-point calibration

 $a = \frac{R_{OBCT} - R_{DSV}}{C_{OBCT} - C_{DSV}}$ 

$$R_{ME} = R_{OBCT} + a * (C_E - C_{OBCT}) + u * a^2 * (C_E - C_{DSV}) * (C_E - C_{OBCT})$$

Brightness temperature by inverse Planck law

$$T_{ME} = Planck^{-1}(R_{ME})$$

Antenna Pattern Correction

$$T = APC_1 + APC_2 * T_{ME}$$

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#### sensor state variables:

- C<sub>DSV</sub> : Deep Space View counts
- C<sub>E</sub> : Earth counts
- COBCT : On Board Calibration Target counts
- ROBCT : On Board Calibration Target radiance
- u : non-linearity factor
- APC1 : pixel dependent Antenna Pattern Correction



3 varied parameters

### measurement equations SNPP (ATMS)

#### Cold and warm target temperatures

$$T_{C} = A + B \left( T_{COSMIC} + dT_{C} \right)$$
$$T_{W} = A + B \left( T_{OBCT} + dT_{W} \right)$$

#### Antenna temperature

$$T_{lin} = T_{W} + (C_{E} - C_{OBCT}) * \frac{T_{W} - T_{C}}{C_{OBCT} - C_{DSV}}$$

Non-linear correction

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$$T_{nonlin} = T_{lin} + u * \left( 1 - 4 \left( \frac{T_{lin} - T_E}{T_W - T_C} - 0.5 \right)^2 \right)$$

Antenna Pattern Correction

$$T = T_{nonlin} - APC$$

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#### sensor state variables:

- C<sub>DSV</sub> : Deep Space View counts
- C<sub>E</sub> : Earth counts
- COBCT : On Board Calibration Target counts
- u : non-linearity factor
- APC : pixel dependent Antenna Pattern Correction



### Harmonisation setup for microwave

### Sensors: MetOp-A, MetOp-B, ATMS, Fengyun-3A, -3B, -3C

- Cold match-ups:
  - difference in space: 10 km
  - difference in time: **5 min**
  - difference in viewing angle: ratio of path length < 0.05
- Intermediate warm match-ups, LEO-GEO:
  - MVIRI homogene 3x3pixel scene, std.dev. < 1 K
  - MVIRI ΔT < 1 K
  - difference in space: 4 km
  - difference in time: 5 min
  - difference in viewing angle: ratio of path length < 0.05
- Warm match-ups, LEO-LEO:
  - difference in space: 10 km
  - difference in time: **6 hours**
  - difference in viewing angle: ratio of path length < 0.05
- 1.5 million match-ups by 9 sensor pairs
- Channel 183.31 ±1 GHz, also available are ±3 GHz, and ±7 GHz
- 15 parameters to optimise: 5 sensors a 3 parameters
- one sensor serves as **reference**, i.e. only radiances with uncertainties are used (not necessary, can be replaced by additional constraint on parameters)

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### Microwave sensor differences

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Sensor Pair

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### **Residuals after harmonisation**

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- distribution is similar for all pairs
- a problematic sensor would stand out

### **Distribution of residuals**



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Final residuals are allmost Gaussian distributed.

### Posterior uncertainty covariance matrix

 $C_i = \left(\frac{\partial J^2}{\partial x^2}\right)^{-1}$ 

posterior uncertainty covariance is inverse Hessian

 $\frac{\partial J^2}{\partial x^2}$ 

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Full Hessian by vector forward over scalar reverse mode AD



correlation matrix

Please note, correlation is covariance divided by variance

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### Conclusions

- Harmonisation of Microwave sensors based on MEIV successful
  - misfit is reduced substantially
  - posterior statistics are consistent with assumptions
  - full range coverage by including 'warm' match-ups from targets changing slowly
- For good results, estimation and propagation of uncertainties is important
- Automatic Differentiation (AD) with TAF (Transformation of Algorithm in Fortran) greatly helped
  - only function code must be written
  - efficient derivative codes generated 'on the fly'
  - can operate in scalar mode for matrix free uncertainty propagation or vector mode for full Jacobian/Hessian computation
- Optimal parameters are used to produce new FCDR's with uncertainties
- Algorithm published in remote sensing 2019, 11(9), 1002; https://doi.org/10.3390/rs11091002



# Outlook

- Analysis of remaining misfit
  - revisit measurement equation formulation
  - introduce new parameter in measurement equation
- Simultaneous harmonisation of several channels if the misfit is correlated (eg. HIRS)
- MEIV Harmonisation can be extended to use other sources of information
  - simulated radiances
  - known targets

