

# The 3MI Instrument onboard MetOp Second Generation

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<sup>1</sup> ESA/ESTEC, Noordwijk, The Netherlands



<sup>2</sup> Leonardo, Business Space, Firenze, Italy



<sup>3</sup> Airbus Defence and Space, Toulouse, France



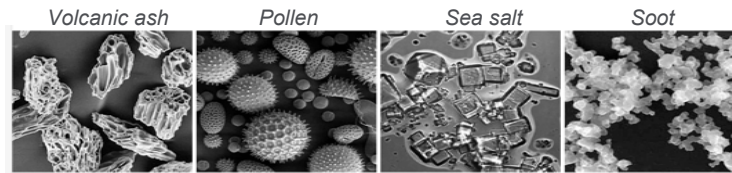
<sup>4</sup> EUMETSAT, Darmstadt, Germany

# Monitor Global Change

## Aerosol

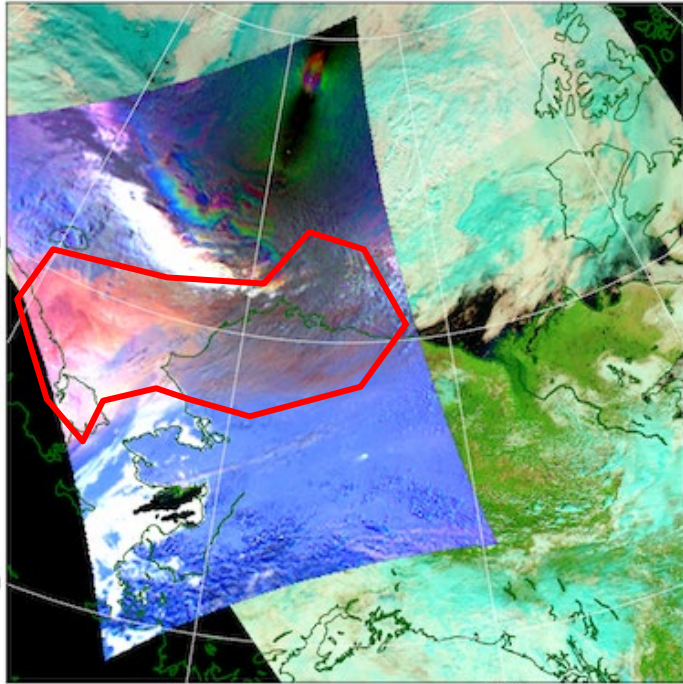


## Aerosol Composition

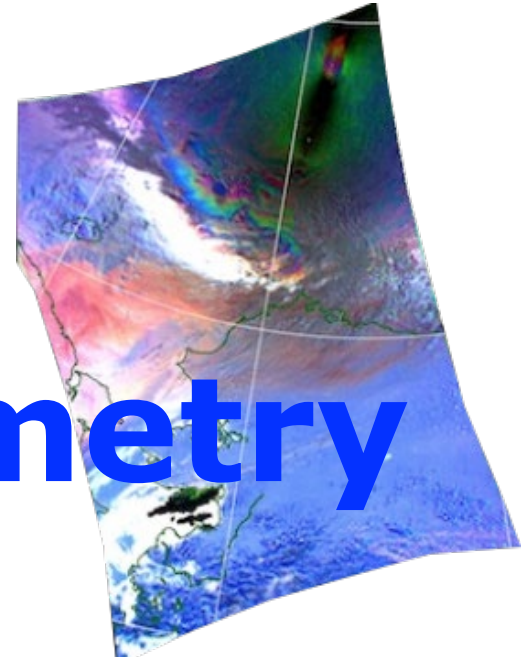


# Monitor Global Change

## Aerosol



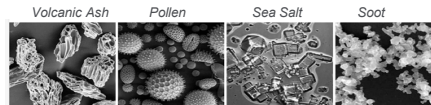
Polarization composite view from POLDER3/PARASOL (CNES mission)  
Superimposed to MODIS/AQUA true color composite



POLDER3/PARASOL (CNES mission)  
(December 2004 - December 2013)

# Polarimetry

## Aerosol Composition



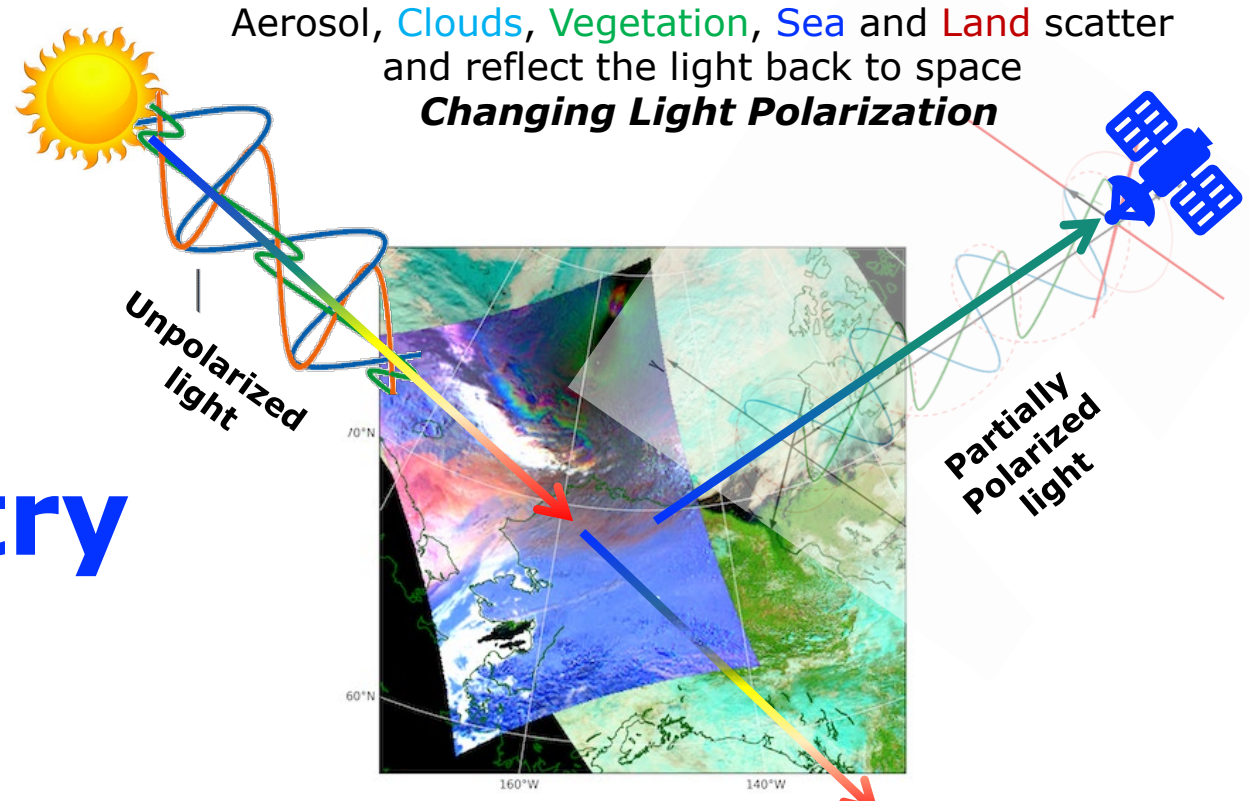
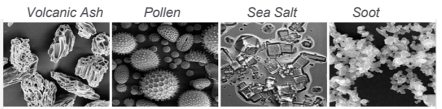
# Polarimetry

## Aerosol



# Polarimetry

## Aerosol Composition



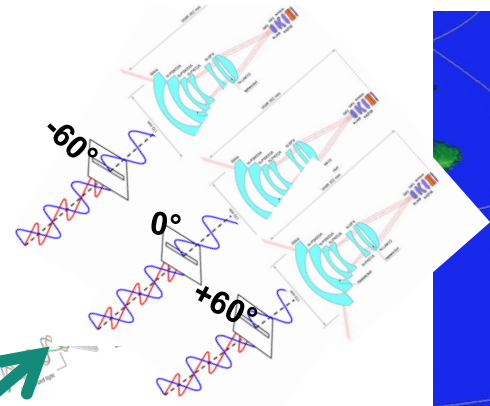
# Polarimetry

$$\begin{pmatrix} I_{k,\ell,p} \\ Q_{k,\ell,p} \\ U_{k,\ell,p} \end{pmatrix} = \frac{1}{G_k} t_k A_k M^{-1} \begin{pmatrix} X_{k,1,\ell,p} \\ X_{k,2,\ell,p} \\ X_{k,3,\ell,p} \end{pmatrix}_{Cor}$$

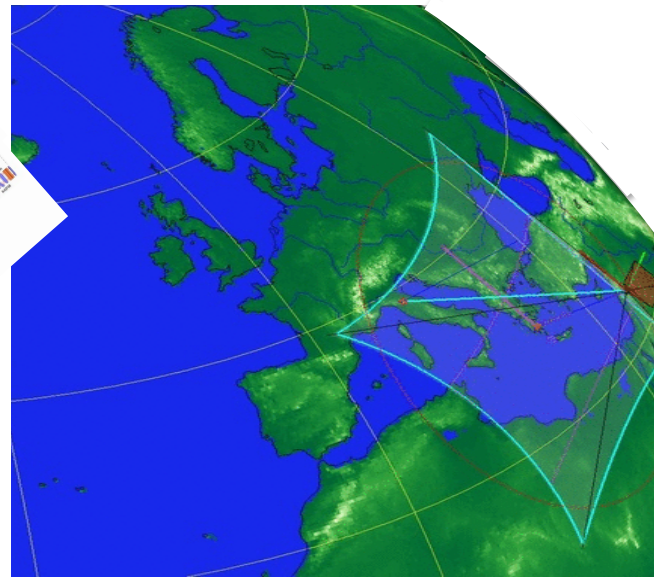
→ Channel k → Pixel LP → Gain G → Integration time t → Absolute Calibration A → X digital counter for:
 

- Channel k
- Polarisation 1 (-60°), 2(0°), 3(+60°)
- Pixel LP

## Multi Polarization

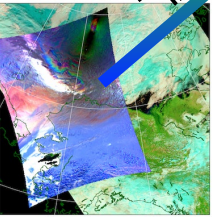


## Multi View



## Multi Channel

	Channel	Central Wavelength [μm]	Channel Spectral Width [μm]	Pol
VNIR	3MI-2	0.410	0.02	Y
	3MI-3	0.443	0.02	Y
	3MI-4	0.490	0.02	Y
	3MI-5	0.555	0.02	Y
	3MI-6	0.670	0.02	Y
	3MI-7	0.763	0.01	N
	3MI-8	0.765	0.04	N
	3MI-9	0.865	0.04	Y
	3MI-9a	0.910	0.02	N
SWIR	3MI-10	1.370	0.04	Y
	3MI-11	1.650	0.04	Y
	3MI-12	2.130	0.04	Y



Partially Polarized light

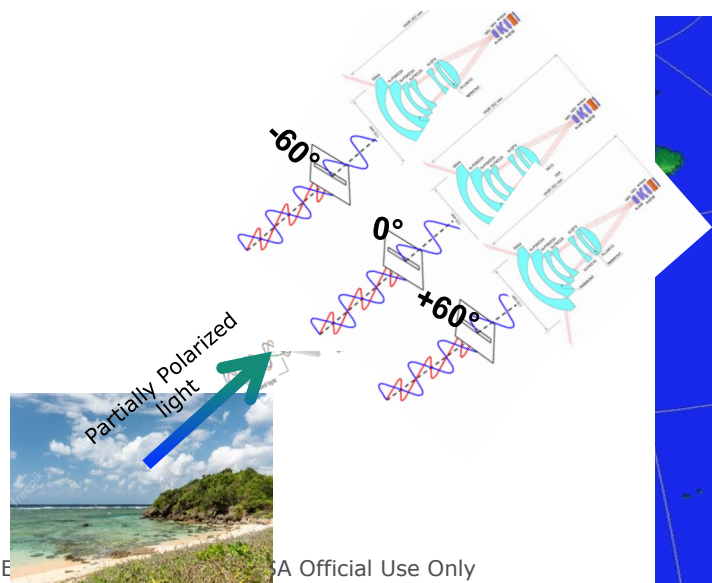
For ESA Official Use Only



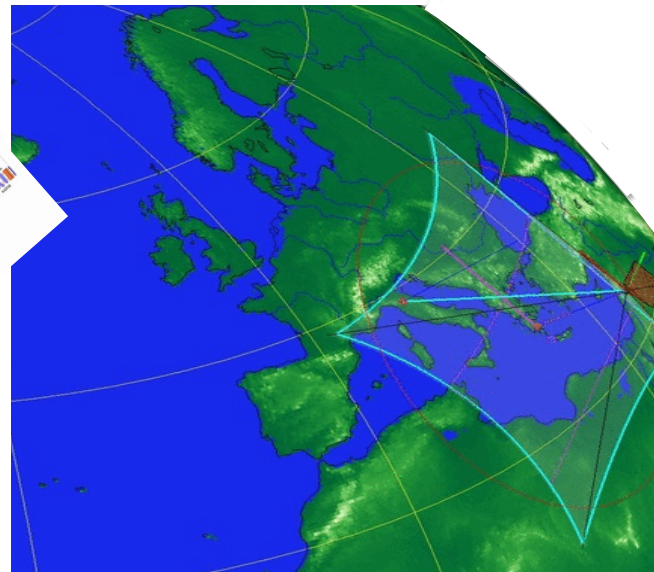
## 3MI dedicated to **Aerosol Characterisation** for:

- Primary Objective: Climate monitoring, Air quality monitoring and forecasting, NWP
- Secondary Objective: Cloud microphysics characterization and Surface BRDF/BPDF

### Multi Polarization

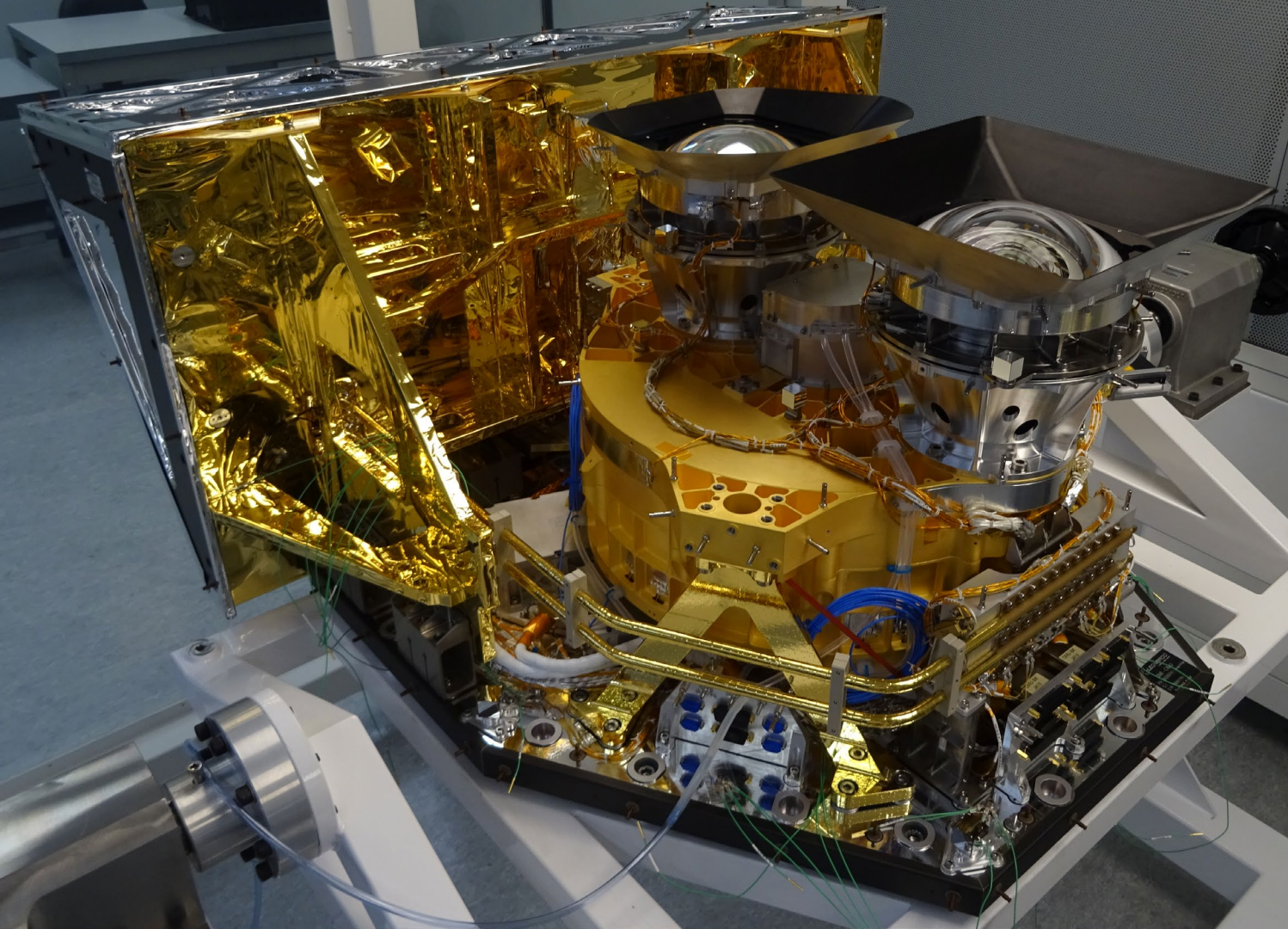


### Multi View

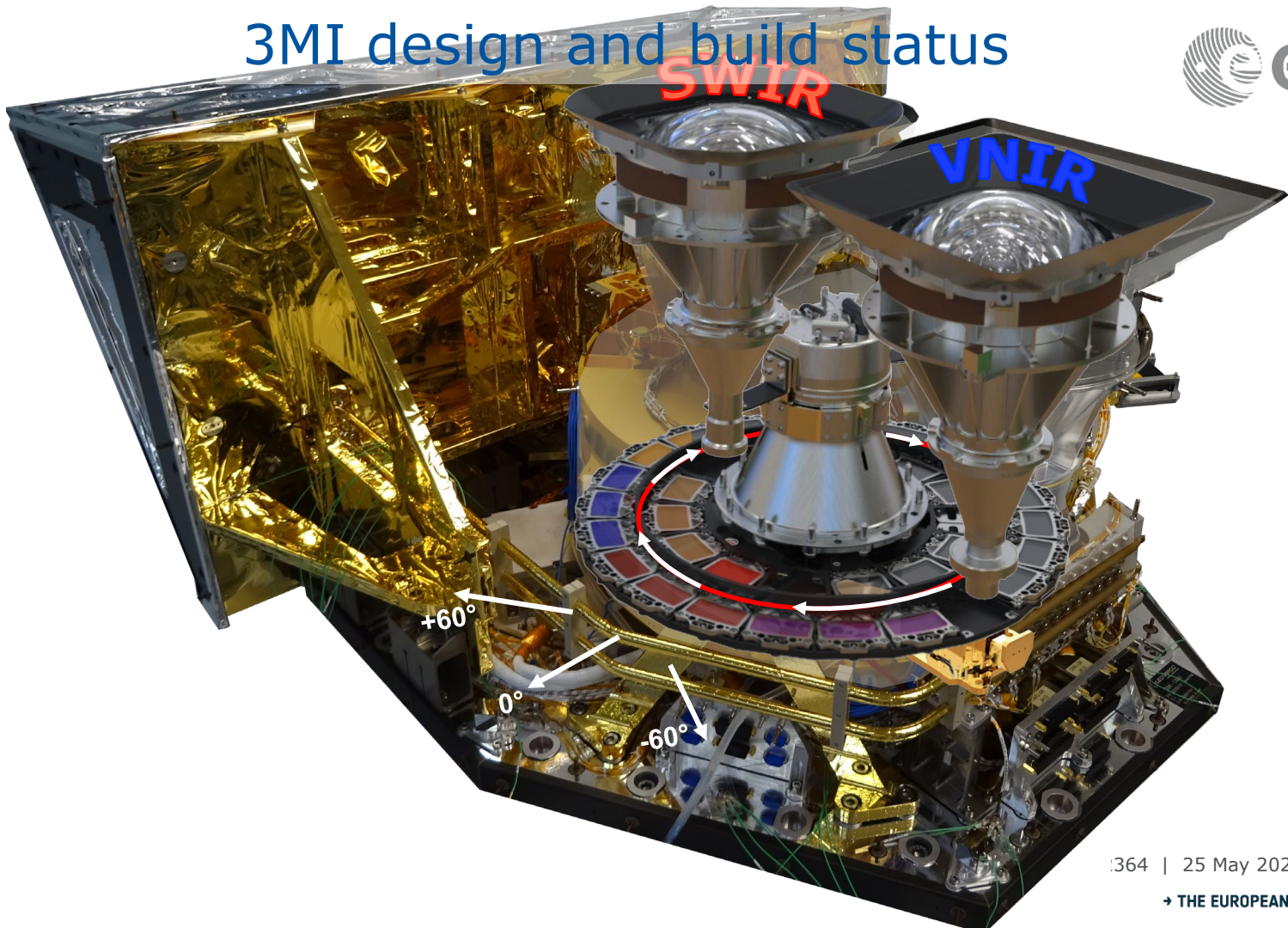


### Multi Channel

	Channel	Central Wavelength [μm]	Channel Spectral Width [μm]	Pol
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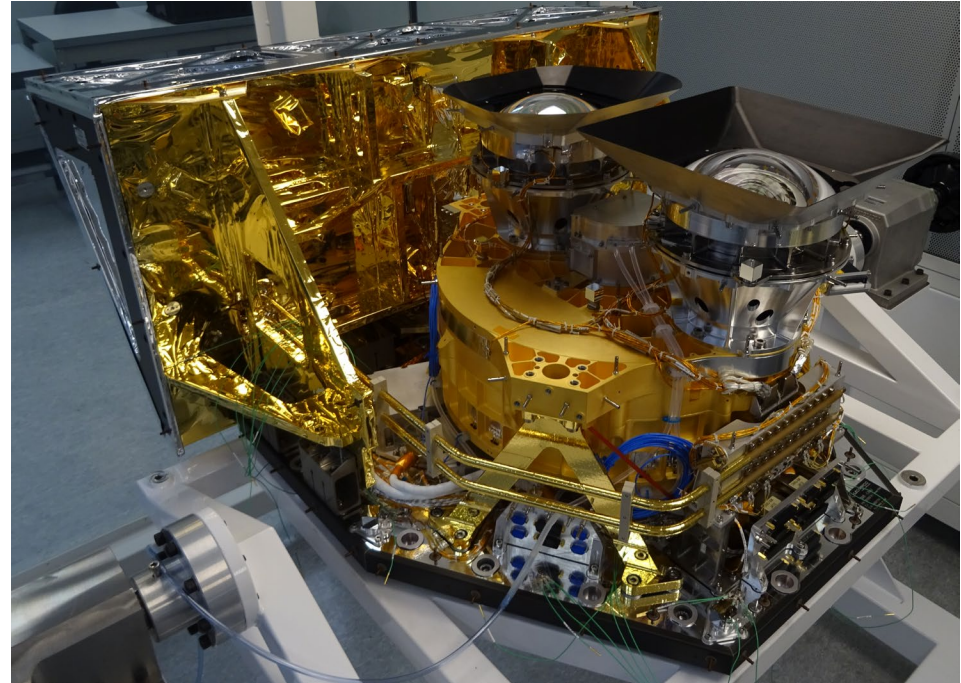
# 3MI design and build status



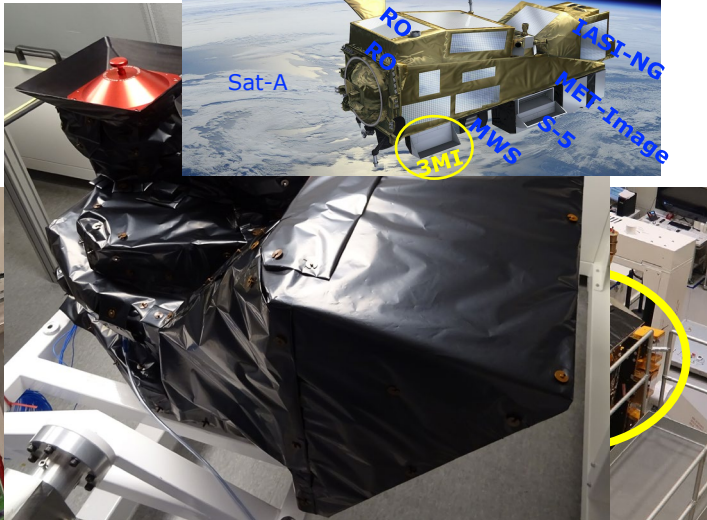
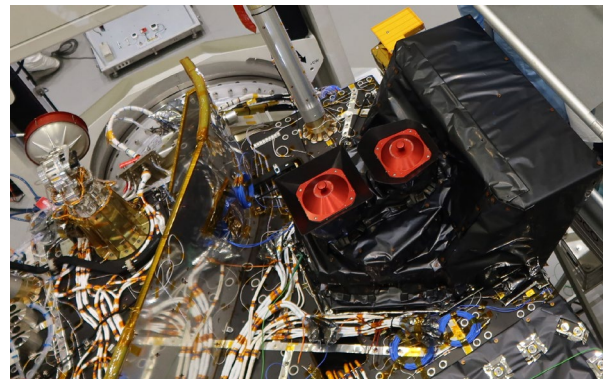
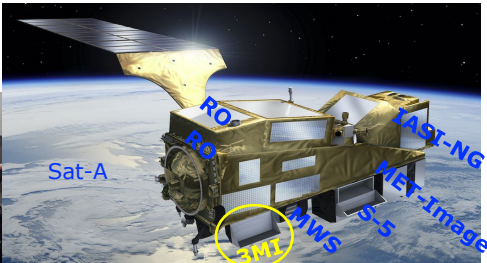




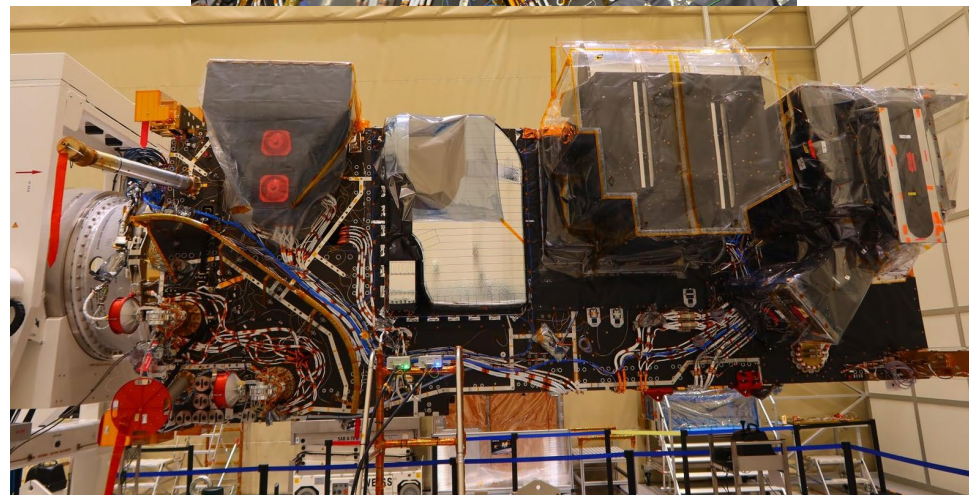
*3MI PFM with black MLI*



*3MI PFM mounted on MGSE*



3MI PFM with black MLI



## 3MI is dedicated to **Aerosol Characterisation** for:

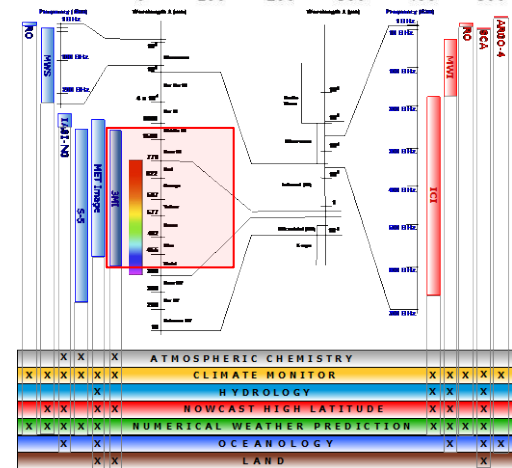
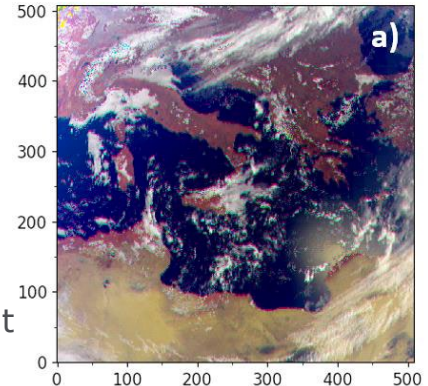
- Primary Objective: Climate monitoring, Air quality monitoring and forecasting, NWP
- Secondary Objective: Cloud microphysics characterization and Surface BRDF/BPDF

## 3MI does *not* have onboard calibration source therefore to reach 3MI's objectives:

- In-flight performances are based on the use of vicarious calibration, therefore on-ground calibration shall be achieved at the maximum possible accuracy
- Design and Performances stability for the 7.5 years of nominal mission is paramount

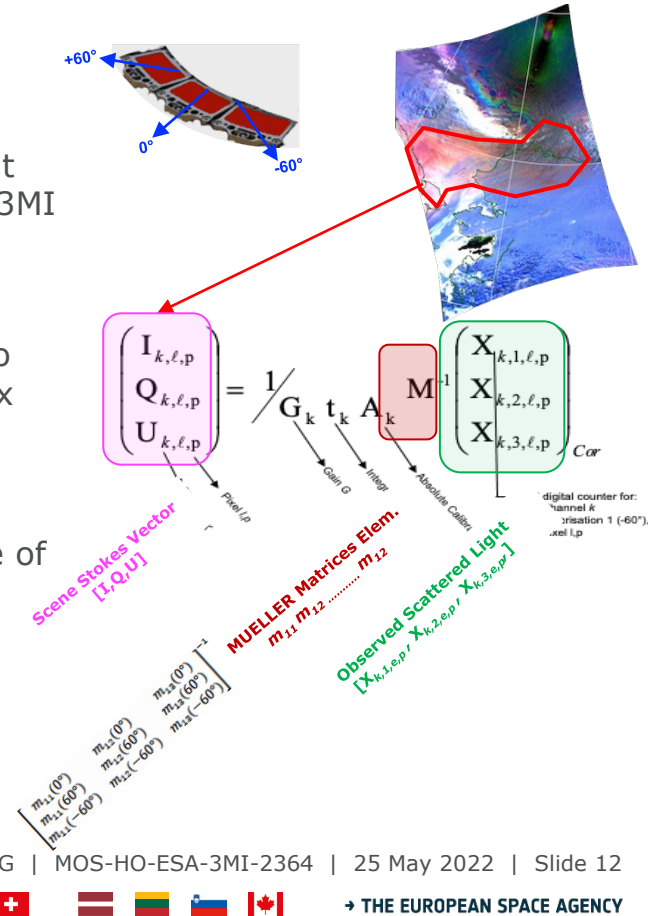
## 3MI Design and On-Ground Calibration shall guarantee:

- *Design* - Instrument short and long term stability:
  - Thermo-Mechanical
  - Radiometric Stability
  - Polarization Sensitivity & Orientation
- *On-Ground Calibration* - Instrument artefact correction:
  - Straylight (main challenge)
  - Linearity
  - Dark current and smearing
  - Flat field
  - Calibration database accuracy



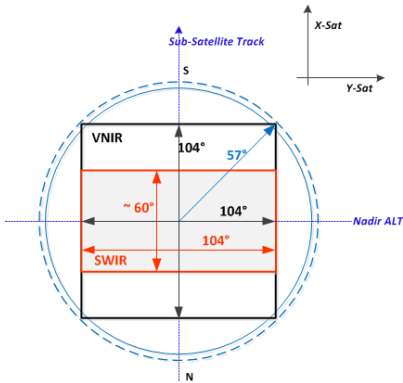
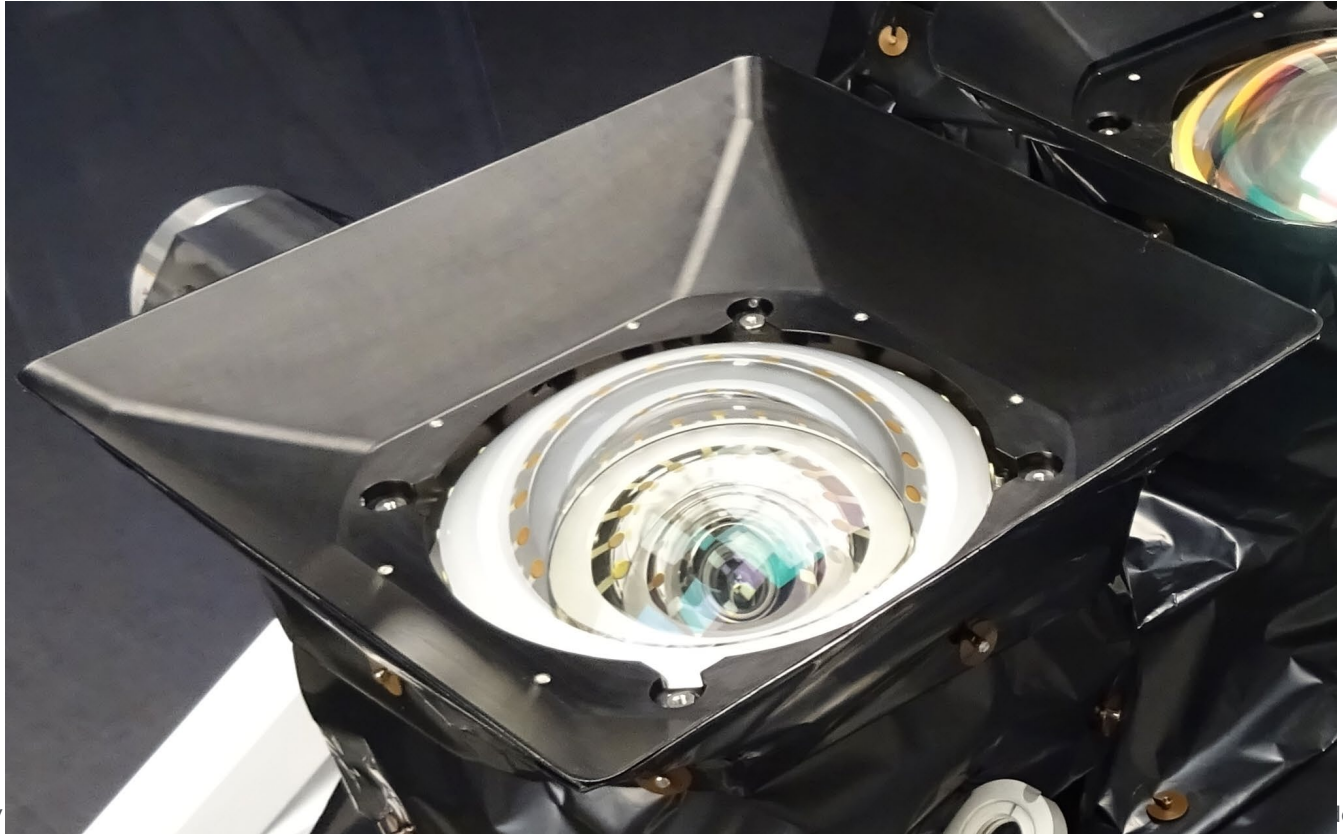
## Polarization Sensitivity and Polarization Orientation

- To calculate **Scene Stokes Vector [I,Q,U]** it is necessary to perform at least three measurements with different polarization angles chosen for 3MI as **0°, -60° and +60°**
- To convert the **Observed Polarized Light [X<sub>k,1,e,p</sub>, X<sub>k,2,e,p</sub>, X<sub>k,3,e,p</sub>]** to the Scene Stokes Vector [I,Q,U] the computation of the MUELLER matrix inversion is needed
- To compute **MUELLER Matrices Element** requires a precise knowledge of absolute and relative polarization angles of each polarizer.





## Straylight Characterization



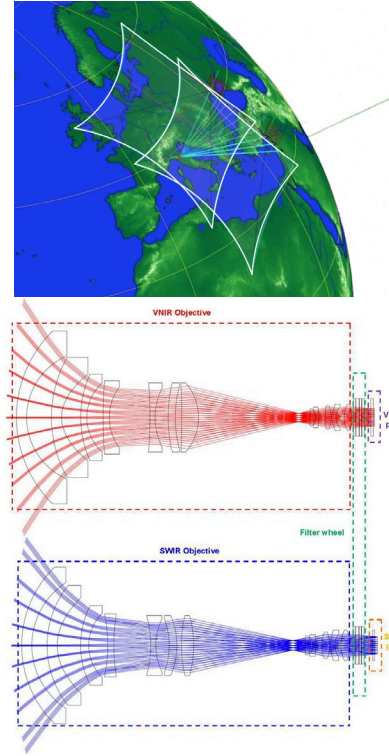
## Straylight Characterization

### Challenges:

- 3MI has a very large FoV **VNIR  $\pm 50.30^\circ$  SWIR  $\pm 31.15^\circ \times \pm 50.85^\circ$**
  - **Complex optical path with 7 lenses (2 aspheric) plus 5 in the focuser**
  - Straylight is corrected using the method of Spatial Point Source Transmittance (SPST) map
  - An SPST map corresponds to the stray-light profile at detector level assuming a punctual field (point source) illumination of the instrument (i.e. collimated beam)
- Full detail on SPST method and correction algorithm can be found in:  
*Stray-light calibration and correction for the MetOp-SG 3MI mission*  
 L. Clermont, C. Michel, E. Mazy, C. Pachot, N. Daddi, C. Mastrandrea, Y. Stockman  
 SPIE 2018 proceeding volume paper 1070406
- Acquire SPST maps to compose a straylight correction database

~~To perfectly calibrate and correct straylight would be needed to calibrate 1 SPST map per each detector's pixels (field) per each of the 13 channels~~

- ~~VNIR  $512 \times 512 = 262.144$~~
  - ~~SWIR  $256 \times 512 = 131.072$~~
- ~~$= 30 \text{ sec} \times \text{SPST} \times 33 \text{ filters slot} = \mathbf{2230 \text{ days}}$~~



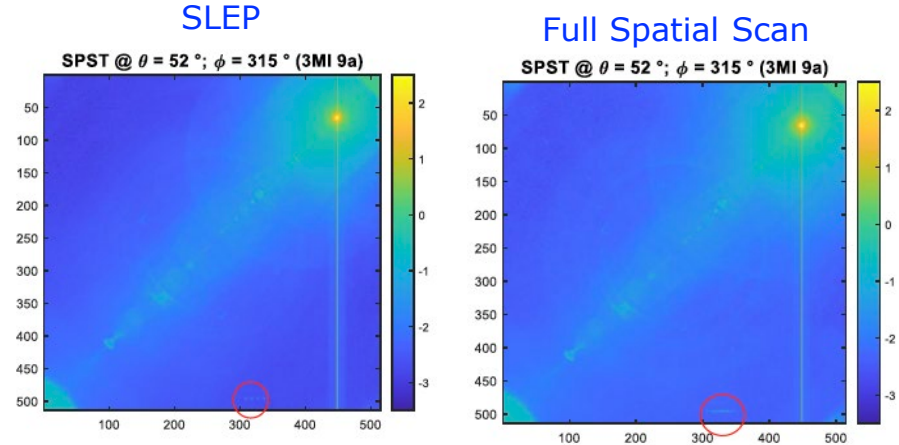
**Impossible task due to measurement time and resulting size of correction database (PByte)**





## Straylight characterization

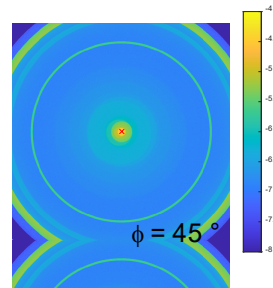
- With a *brutal force* method a scan of the first full lens per each of the Calibration Grid point would be needed
- “**S**tray-**L**ight **E**ntrance **P**upil” (SLEP) method has been developed
- The SLEP corresponds to the minimum area through which rays which are sent will result in straylight at the detector. This pupil is a function of wavelength and field of view
- The SLEP concept has been validated by 3MI EM measurements campaign and confirmed with PFM Calibration preliminary analysis
- No additional SL feature from 3MI have been detected performing measurements with the full spatial scan instead of the SLEP
- Full spatial scan would also introduce more noise as well as more SL from the OGSE



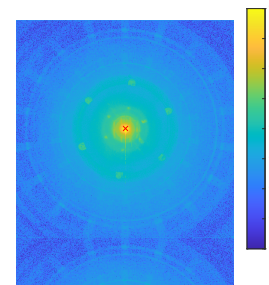
## Straylight characterization

- Measurements at 910nm (3MI-9a channel) has demonstrated accurate measurements of the different SL features
- Similarities with the theoretical maps have been obtained

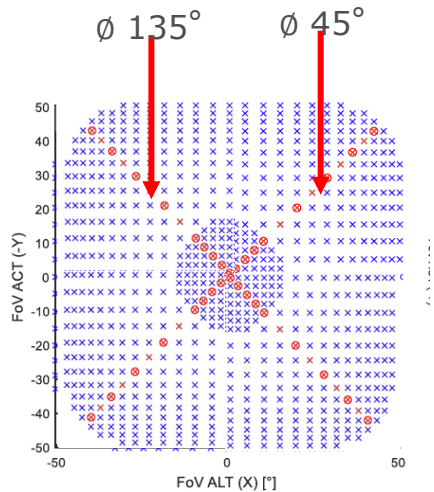
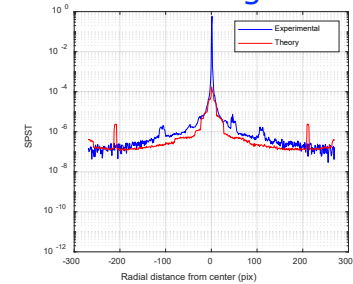
Simulated



Measured

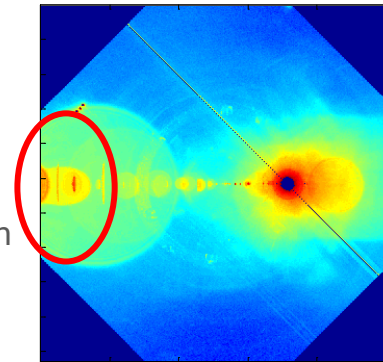


Radial diagram



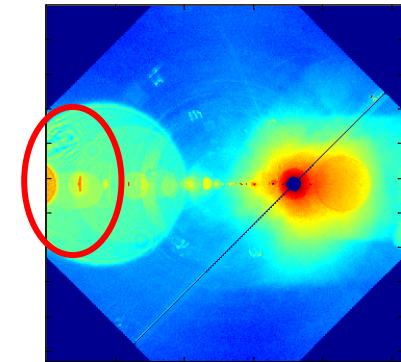
- We tried to exploit revolution symmetry to calibrate *only* 1/4 of FoV
- Polarized ghosts power varies with the azimuth angle

$t = 2.5 \text{ ms} ; \theta = 24^\circ ; \phi = 45^\circ$



Acquisition  $\phi 45^\circ$

$t = 2.5 \text{ ms} ; \theta = 24^\circ ; \phi = 135^\circ$



Acquisition  $\phi 135^\circ$

## Key Innovations:

- 3MI is the first optical polarimeter to fly on an operational satellite
- First time with SWIR channels.
- Filter wheel system with 33 apertures.
- Extensive and novelty method for Straylight characterization

## Benefit:

- 3MI will provide a unique long-term of operational time series for atmosphere characterization for aerosol and clouds

Channel	$\lambda$ Centre (nm)	Bandwidth (nm)	Polarization	Utilization	
VNIR	3MI-2	410	20	-60°, 0°, +60°	Aerosol absorption, ash cloud
	3MI-3	443	20	-60°, 0°, +60°	Aerosol absorption, height indicators
	3MI-4	490	20	-60°, 0°, +60°	Aerosol, surface albedo, cloud reflectance, cloud optical depth
	3MI-5	555	20	-60°, 0°, +60°	Surface albedo
	3MI-6	670	20	-60°, 0°, +60°	Aerosol properties
	3MI-7	763	10	Non-polarize	Cloud and aerosol height
	3MI-8	765	40	Non-polarize	Cloud and aerosol height
	3MI-9	865	40	-60°, 0°, +60°	Vegetation, aerosol, clouds, surf. Features
	3MI-9a	910	20	Non-polarize	Water vapour, atmospheric correction
	3MI-10	1370	40	-60°, 0°, +60°	Cirrus clouds, water vapour imagery
SWIR	3MI-11	1650	40	-60°, 0°, +60°	Ground characterisation for aerosol inversion
	3MI-12	2130	40	-60°, 0°, +60°	Ground characterisation for aerosol inversion, clouds microphysics, vegetation, fire effects

3MI PFM Calibration has lasted for 85 days 24/7 for a total of 2040 hours of data acquisition in vacuum

## Authors

Porciani M. <sup>1</sup>, Celerier N. <sup>3</sup>, Daddi N. <sup>2</sup>, Fuente I. <sup>1</sup>, Granclaude B. <sup>3</sup>, La China F. <sup>2</sup>, Lee C. <sup>4</sup>, Loiselet M. <sup>1</sup>, Mason G. <sup>1</sup>, Pennestri G. <sup>2</sup> Pomilia A. <sup>2</sup>

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<sup>3</sup> Airbus Defence and Space, Toulouse, France

<sup>4</sup> EUMETSAT, Darmstadt, Germany

## Images Credits

- Leonardo, Business Space, Florence (I) – Instrument and optics
- Airbus Defence and Space, Toulouse (F) – Satellite
- RUAG Space, Zurich (CH) – Filter Wheel Subsystem
- Sodern, Limeil-Brévannes (F) – VNIR and SWIR Focal Plan Array and Front End Electronics
- Invent, Braunschweig (D) – CFRP base plate
- AVIOTEC, Torino (I), - MLI
- Centre Spatial de Liège, Liège (B) – Calibration Facility and Straylight

# The 3MI Instrument onboard MetOp Second Generation

