



25/05/2022

NanoMagSat, an optimal New Space low-Earth orbiting nanosatellite constellation to investigate Earth's magnetic field and ionospheric environment

**G. Hulot**<sup>1</sup>, P. Coïsson<sup>1</sup>, J.M. Léger<sup>2</sup>, T. Jager<sup>2</sup>, L.B. Clausen<sup>3</sup>, F. Deconinck<sup>4</sup>, et al.

Université Paris Cité, Institut de Physique du Globe de Paris, CNRS, F-75005 Paris, France
 CEA-Leti, Université Grenoble Alpes, F-38000 Grenoble, France

3. Department of Physics, University of Oslo, Oslo, Norway

CITS

cea

4. Open Cosmos, Didcot, United Kingdom

Session B7.01: Scout - ESA NewSpace Science Missions

IPG

# symposium 2022



# NanoMagSat Science Team additional co-authors

P.Alken<sup>5</sup>, J. Aubert<sup>1</sup>, G. Balasis<sup>6</sup>, C. Beggan<sup>7</sup>, W. Brown<sup>7</sup>, D. Buresova<sup>8</sup>, L. Chauvet<sup>1</sup>, A. Chulliat<sup>5</sup>, G. Consolini<sup>9</sup>, P. De Michelis<sup>10</sup>,
A. De Santis<sup>10</sup>, M.W. Dunlop<sup>11</sup>, C.C. Finlay<sup>12</sup>, A. Grayver<sup>13</sup>, B. Heilig<sup>14</sup>, J. Jorgensen<sup>12</sup>, I. Kolmasova<sup>8</sup>, A. Kuvshinov<sup>13</sup>, V. Lesur<sup>1</sup>,
P. Livermore<sup>15</sup>, J. Mlynarczyk<sup>16</sup>, N. Olsen<sup>12</sup>, M. Rother<sup>17</sup>, O. Santolik<sup>8</sup>, E. Thébault<sup>18</sup>, V. Truhlik<sup>8</sup>, J. Velimsky<sup>19</sup>, P. Vigneron<sup>1</sup>

1. Université Paris Cité, Institut de Physique du Globe de Paris, CNRS, F-75005 Paris, France

5. CIRES, University of Colorado Boulder, USA 6. IAASARS-National Observatory of Athens, Athens, Greece 7. British Geological Survey, The Lyell Centre, Edinburgh, United Kingdom 8. Institute of Atmospheric Physics of the Czech Academy of Sciences, Prague, Czech Republic 9. INAF-Istituto di Astrofisica e Planetologia Spaziali, Rome, Italy 10. Istituto Nazionale di Geofisica e Vulcanologia, Rome, Italy 11. RAL\_Space, STFC, Chilton, Oxfordshire, United Kingdom 12. DTU Space, Technical University of Denmark, Kongens Lyngby Denmark 13. Institute of Geophysics, Department of Earth Sciences, ETH Zurich, Switzerland 14. Eotvos Lorand University, Budapest, Hungary 15. University of Leeds, Leeds, United Kingdom 16. AGH University of Science and Technology, Institute of Electronics, Krakow, Poland 17. Helmholtz Centre Potsdam, GFZ German Research Centre for Geosciences, Potsdam, Germany 18. Laboratoire Magma et Volcans, Université Clermont Auvergne, CNRS, IRD, OPGC, Clermont-Ferrand, France 19. Faculty of Mathematics and Physics, Charles University, Prague, Czech Republic

#### Session B7.01: Scout - ESA NewSpace Science Missions

#### 25/05/2022

# NanoMagSat programmatic status

- Submitted to the first ESA Scout call in August 2019
- One of the four missions selected for a Consolidation Study in November 2019
- Consolidation study leading to a **full report to ESA in August 2020**
- Presentation to ACEO in October 2020, very highly evaluated:

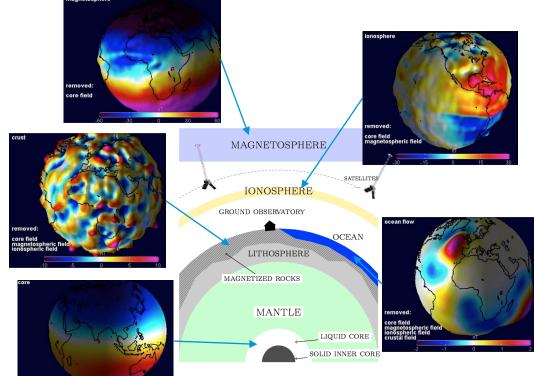
"ACEO confirms the scientific relevance of the NanoMagSat concept"

"ACEO confirms the uniqueness of the NanoMagSat concept"

"unique opportunity for demonstrating long-term magnetic field monitoring, as a continuation of the Swarm mission"

- But not selected on ground of high risks of not meeting costs and schedule criteria
- 18 months ESA funded Risk Retirements Activities kicked off in January 2022, with already good progress, see presentations on technical progress tomorrow (26/05/22):
- -> session B7.03, 14:45 J.M. Léger et al., "NanoMagSat magnetometry payload"
- -> **session B7.03**, **15:55**, **F. Deconink et al.** "NanoMagSat: a cost-effective 3x16U satellite constellation for fast recovery of the Earth magnetic field and the ionospheric environment"
- -> Poster 53, L. Clausen et al. "The multi-needle Langmuir Probe on board NanoMagSat"

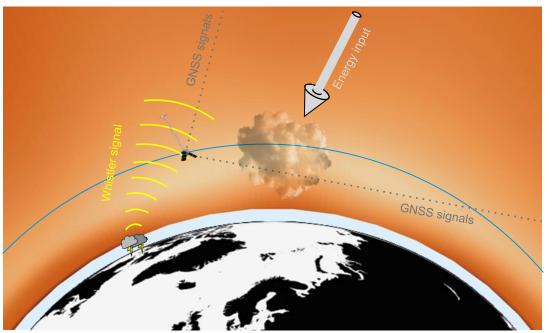
# First family of objectives: Earth's magnetic field



- Targeting recovery of fast planetary changes in core, ionospheric and magnetospheric fields, also improving recovery of crustal and oceanic signals
- To investigate fast core dynamics, solar-terrestrial interactions, crust and deep Earth properties and possible signatures of climate change

### Taking advantage of an **innovative constellation**

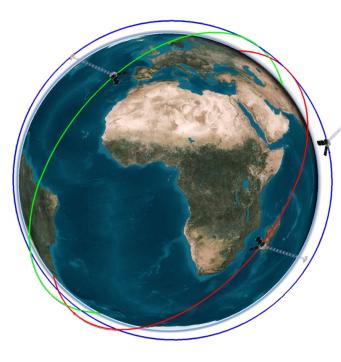
# Second family of objectives: Ionospheric environment



- Targeting ionospheric plasma density dynamics
- Combining standard GNSS techniques
- With new ways of sensing the ionosphere below the satellites and monitoring in situ meter to km scale dynamics and energy input
- To investigate Space weather phenomena that affect radio and GNSS signals, and improve science and operational ionospheric models

### Taking advantage of an **innovative compact payload**

# NanoMagSat mission concept



#### Innovative orbital configuration

A 3-year minimum lifetime constellation composed of 3x16U Cubesats at 575 km initial altitude

- 1 satellite at 60° inclination
- 1 satellite at 60° inclination offset by 90°-RAAN
- 1 satellite in near-polar orbit

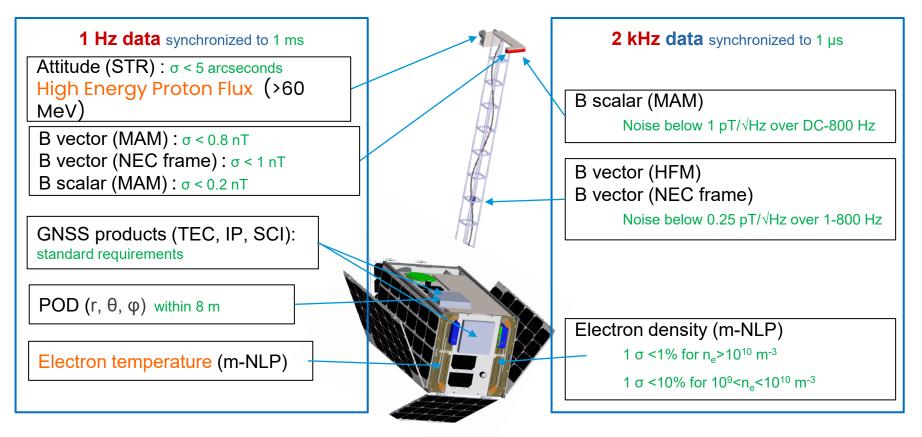
#### State of the art compact payloads

- A Miniaturized Absolute Magnetometer (MAM) with a set of two Star Cameras (STR)
- A High Frequency Magnetometer (HFM)
- A multi-Needle Langmuir Probe (m-NLP)
- 2 dual-frequency GNSS

Initiating a **low-cost scalable collaborative constellation solution for very long-term observations** (extending to space the Intermagnet network of magnetic observatories)

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# NanoMagSat magnetic & ionospheric environment data



# Identified NanoMagSat high level Science targets

#### Constellation

#### Primary objectives:

- Earth's Main Field (Core Field CF Product)
  - $\rightarrow$  MHD processes in the core, improving predictability
- Ionospheric Field (Solar Quiet Field SQF Product)
  - → Investigating thermospheric ionospheric and magnetospheric ionospheric interactions
- Magnetospheric Field (Fast Magnetic Field FMF Product)
  - → Investigating Solar-Terrestrial interactions, recovering solid Earth electrical conductivity properties

#### Secondary objectives:

- Improving mapping of Lithospheric Field (LF Product)
  - $\rightarrow$  With applications to tectonics, geology and natural resources
- Oceanic Lunar Tide Field signals (OLTF Product)
  - $\rightarrow$  Recovering solid Earth electrical conductivity properties, and monitoring global heat content of oceans
- Global Ocean Circulation variability (Under consideration)
  - $\rightarrow$  Monitoring dynamics of large scale ocean currents

#### Single satellite

#### Primary objectives:

- ELF magnetic and Whistler Signals (WS Product)
  - $\rightarrow$  Large-scale ionospheric plasma variability
- Ionospheric meter-scale plasma density structures (MSPDS Product)
  - $\rightarrow$  Small-scale ionospheric plasma variability

#### Secondary objectives:

- Monitoring ionospheric high latitude and equatorial local currents (FAC and EEF Products)
  - → Investigating thermospheric ionospheric and magnetospheric ionospheric interactions
- Monitoring ionospheric Total Electron Content (TEC Product)
- Producing occultation lonospheric Profiles (IP Product)
- Characterising ionospheric scintillations along GNSS radio links (SCI product)
  - $\rightarrow$  lonospheric plasma variability
- High Energy proton flux (Under consideration)
  - $\rightarrow$  Radiation environment monitoring

# NanoMagSat Geographic/Local Time coverage

 $0 \leq LT \leq 3$ 

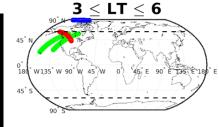
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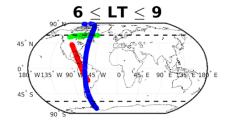
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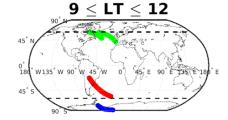


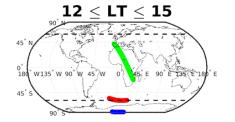




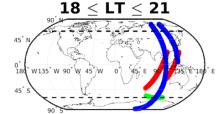


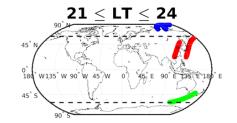






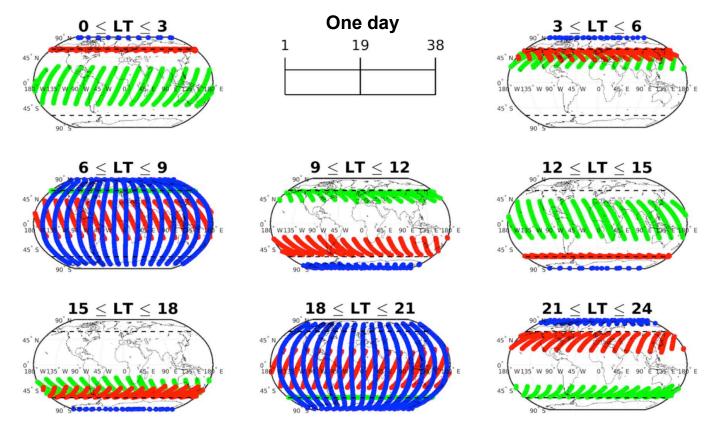
15 ≤ LT ≤ 18 45° N 0° 180(W135° W 90° W 45° W 0) 45° E 90° E 135° E 180° E 45° S



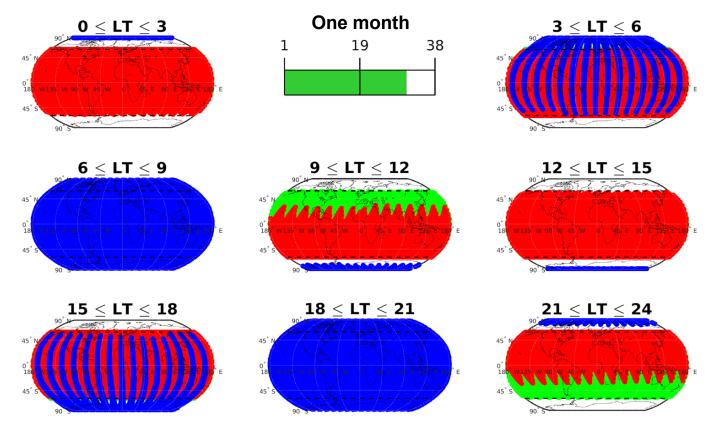


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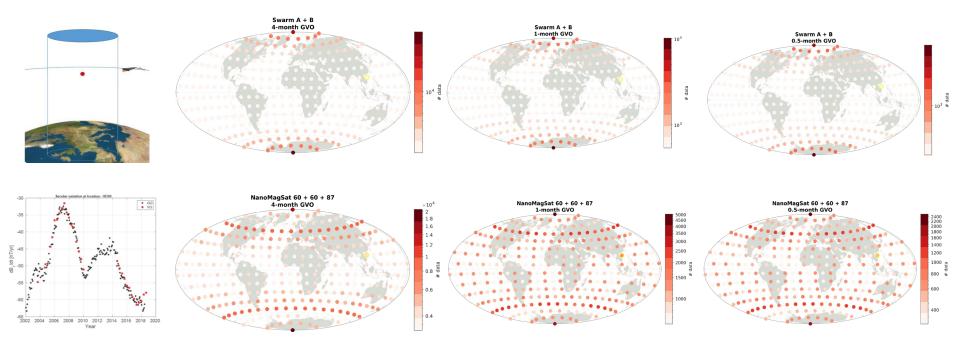
# NanoMagSat Geographic/Local Time coverage



# NanoMagSat Geographic/Local Time coverage

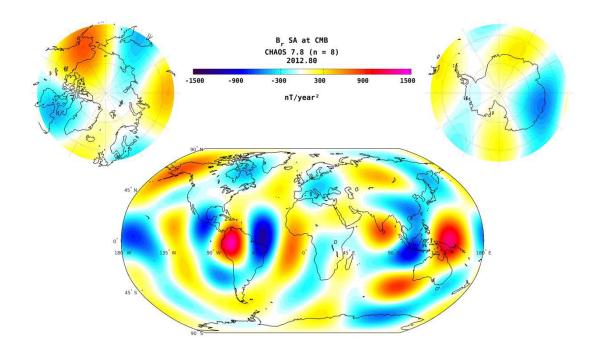


### Will lead to much improved series of Geomagnetic Virtual Observatory (GVO) data for Core Field investigations



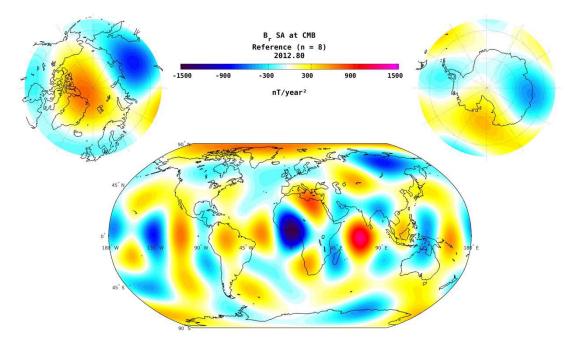
High-quality full LT GVO could then be created on a 3-weekly basis, substantially increasing the current time resolution of 17 weeks based on Swarm. (see **Poster 550 on Thursday 26/05/2022 by Brown et al.,** *Geomagnetic Virtual Observatories for investigating sub-annual core field variation*).

### **Core Field Secular Acceleration currently recovered by Swarm**



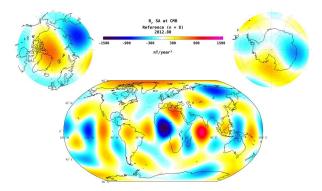
Radial component of the second time derivative of the core field at the core surface up to SH degree 8 between 2012.8 and 2015.8 (3 years), as modelled by the CHAOS 7.8 model (Finlay et al., EPS, 2020)

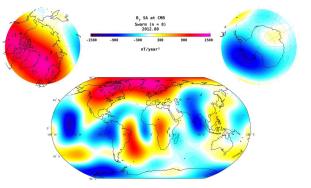
#### Additional faster Secular Acceleration dynamics predicted by dynamo simulations

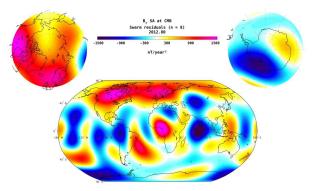


Radial component of additional faster signal predicted in the second time derivative of the core field at the core surface up to SH degree 8, simulated between 2012.8 and 2015.8 (3 years), using dynamo runs by J. Aubert, IPGP (based on Aubert, GJI, 2015; Aubert et al., JFM 2017; Aubert and Gillet, GJI 2021).

### This signal can hardly be recovered by Swarm

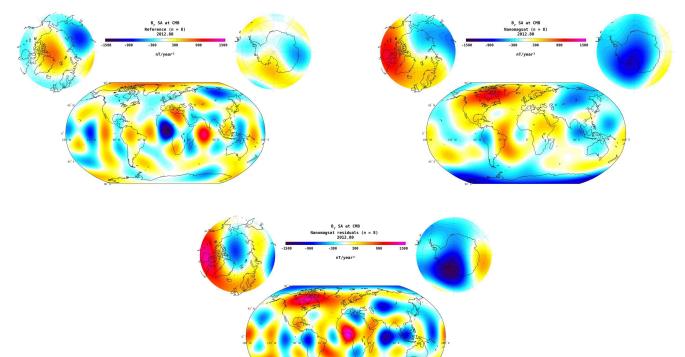






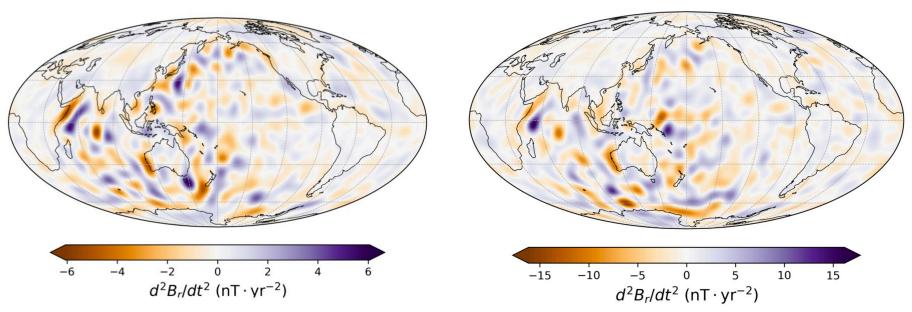
Radial component of the additional faster signal predicted in the second time derivative of the core field at the core surface up to SH degree 8 (upper left), recovered by Swarm (upper right) and residuals (bottom). Model recovery based on methodology of Alken et al., EPS, 2020.

### NanoMagSat can !



Radial component of the additional faster signal predicted in the second time derivative of the core field at the core surface up to SH degree 8 (upper left), recovered by NanoMagSat (upper right) and residuals (bottom). Model recovery based on methodology of Alken et al., EPS, 2020.

### NanoMagSat could also help recover magnetic signals produced by global ocean circulation

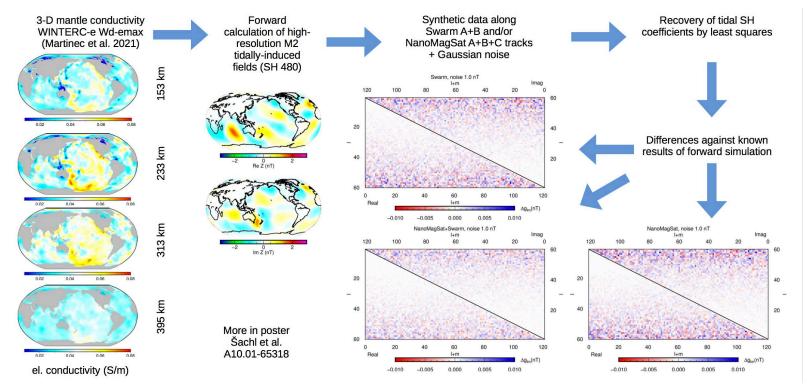


Simulated Secular Acceleration between degrees 9-30 at Earth's surface

Recovered Secular Acceleration between degrees 9-30 at Earth's surface

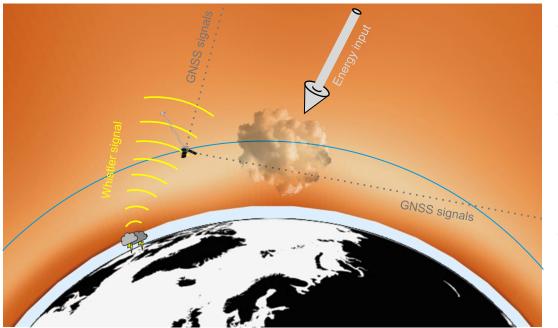
Average Secular Acceleration of the field over 6 months is likely dominated by that produced by Ocean water circulation (and not by the core dynamo) beyond SH degree 8, which Swarm can potentially recover, even better so with NanoMagSat (see Poster 548 on Thursday 26/05/22, by Finlay et al., "Signatures of the global ocean circulation in geomagnetic secular variation and acceleration")

### NanoMagSat will also recover magnetic signals produced by M2 tides



NanoMagSat does as well as Swarm (as shown on Poster 475 on Thursday 26/05/22, by Sachl et al., "Validation and interpretation of the 3-D upper-mantle electrical conductivity reconstructed from the satellite-observed tidal magnetic fields")

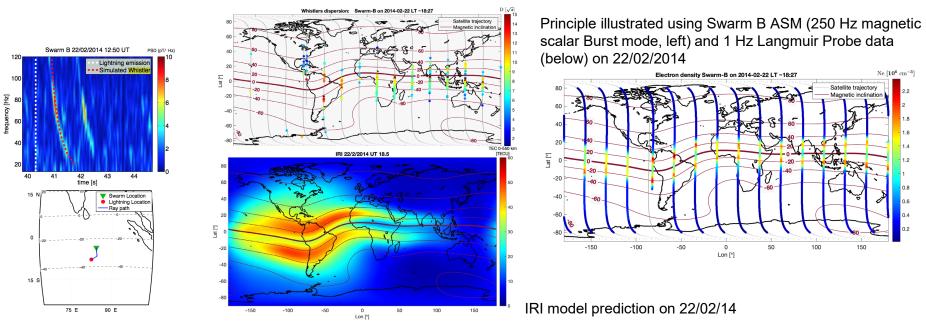
## What about Ionospheric environment ?



- Targeting ionospheric plasma density dynamics
- Combining standard GNSS techniques
- With new ways of sensing the ionosphere below the satellites and monitoring in situ meter to km scale dynamics and energy input
- To investigate Space weather phenomena that affect radio and GNSS signals, and improve science and operational ionospheric models

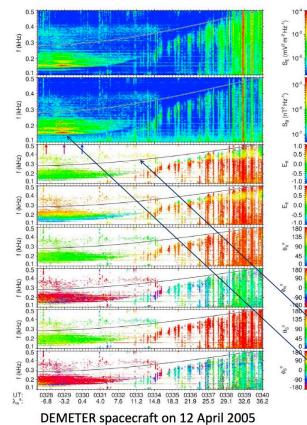
### Taking advantage of an innovative compact payload

### Monitoring the ionosphere below NanoMagSat will be possible



- Take advantage of ELF Whistlers produced by lightning, detected by the MAM (scalar) and HFM (vector) magnetometers (2 kHz)
- These propagate through the ionosphere with a dispersion that allows sensing the ionospheric plasma below the satellites
- Complementing local m-NLP measurements
- Providing information in regions not covered by ionosonde or GNSS techniques
- See talk on Thursday 26/05/22 (Session A6.02, 14:15), P. Coïsson et al., "Whistler in ELF detected from LEO: lightning detection and ionospheric monitoring using Swarm satellites and the future NanoMagSat mission"

# Investigating ELF waves coming from above will also be possible (as shown by Demeter)



# Propagation of equatorial noise to low altitudes

Power-spectral densities of electric field components (not available on NanoMagSat).

Sum of the power-spectral densities of three orthogonal magnetic field components (available on NanoMagSat).

Ellipticity of the magnetic field polarization with a sign corresponding to the sense of polarization (available on NanoMagSat).

Ellipticity of the electric field polarization (not available on NanoMagSat).

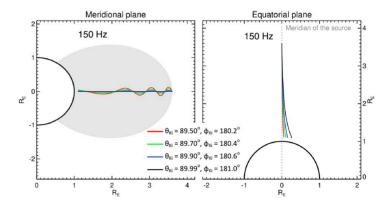
Angle between the wave vector and the background magnetic field (available on NanoMagSat with 180° ambiguity).

Azimuth of the wave vector with respect to the outward direction, positive eastward (available on NanoMagSat with 180° ambiguity).

Poynting vector direction (not available on NanoMagSat).

$$f_{L=0} = f_{H^+} \frac{1 + P_0(R_{OH} - 1)}{R_{OH}}$$

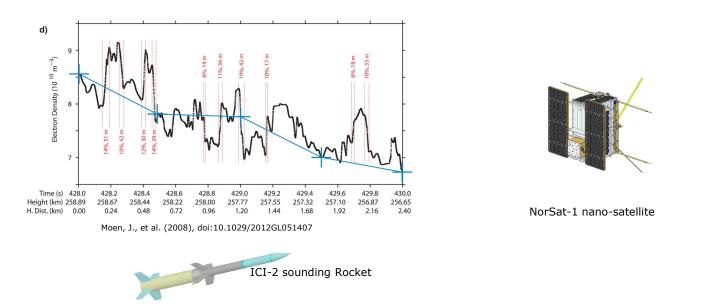
Santolik et al., Geophys. Res. Lett., 2016



See Poster 527, Thursday 26/05/22, by O. Santolik et al., "Coupling of electromagnetic waves between the magnetosphere and the topside ionosphere: new proposed science targets for the NanoMagSat mission".

#### ESA Living Planet Symposium, Bonn, Germany

### As well as investigating in-situ small scale plasma variability



- First joint measurements of electron density and magnetic signals at high frequencies (2 kHz) to access meter to km scale plasma variability and associated electrical currents using the multi-Needle Langmuir Probe (m-NLP), only flown so far on sounding rockets and the NorSat-1 nano-satellite, and the High Frequency Magnetometer (HFM), which NorSat-1 fails to have
  - See poster 53 on Thursday 26/05/22, by L. Clausen "The multi-needle Langmuir Probe on board NanoMagSat"

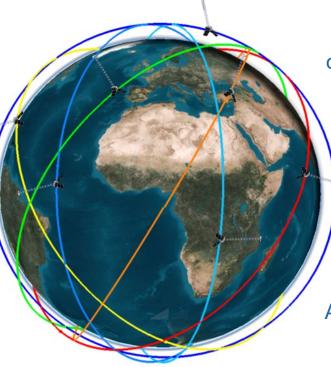
### More examples of the benefit of NanoMagSat for ionospheric studies at this meeting

- Talk this afternoon at 15:40 in session A6.01 by B. Heilig "LEO observations of conjugated magnetospheric-ionospheric phenomena for improved plasmapause monitoring"
- Poster 539 on Thursday 26/05/22 by M. Dunlop et al., "Coordination of ground based and in orbit multipoint measurements: comparison of magnetospheric and ground currents"

# NanoMagSat

A mission for a wide scientific community and 50+ institutional users, strongly aligned with ESA EO Science Strategy

A nanosatellite constellation to investigate Earth's magnetic field and ionospheric environment in a unique way.



A mission ensuring continuity in the monitoring of Earth's magnetic field and ionospheric environment (long-term Science).

A stepping stone for a space network of observatories with international scalability.

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