

Simulations of low-low
satellite-to-satellite tracking
constellations for the ESA/NASA
MAGIC satellite gravity mission

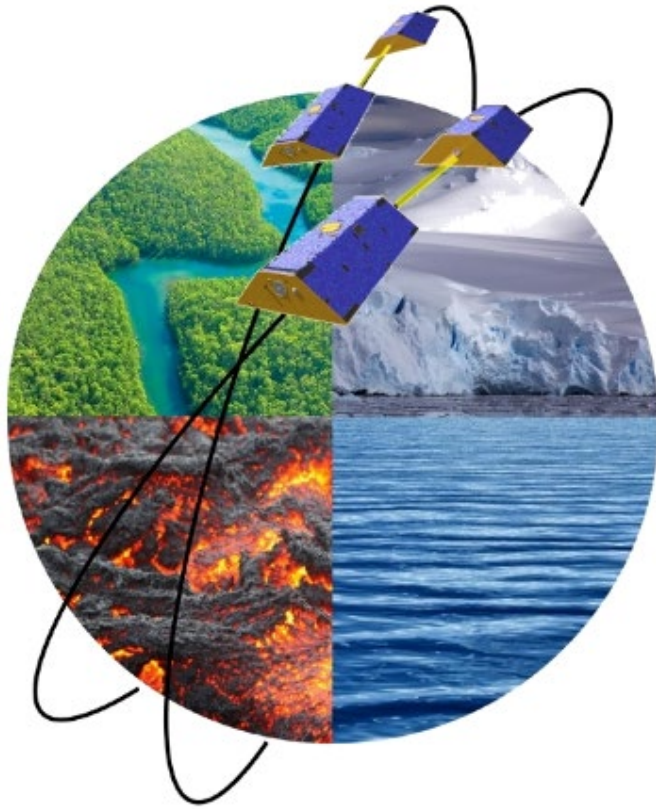
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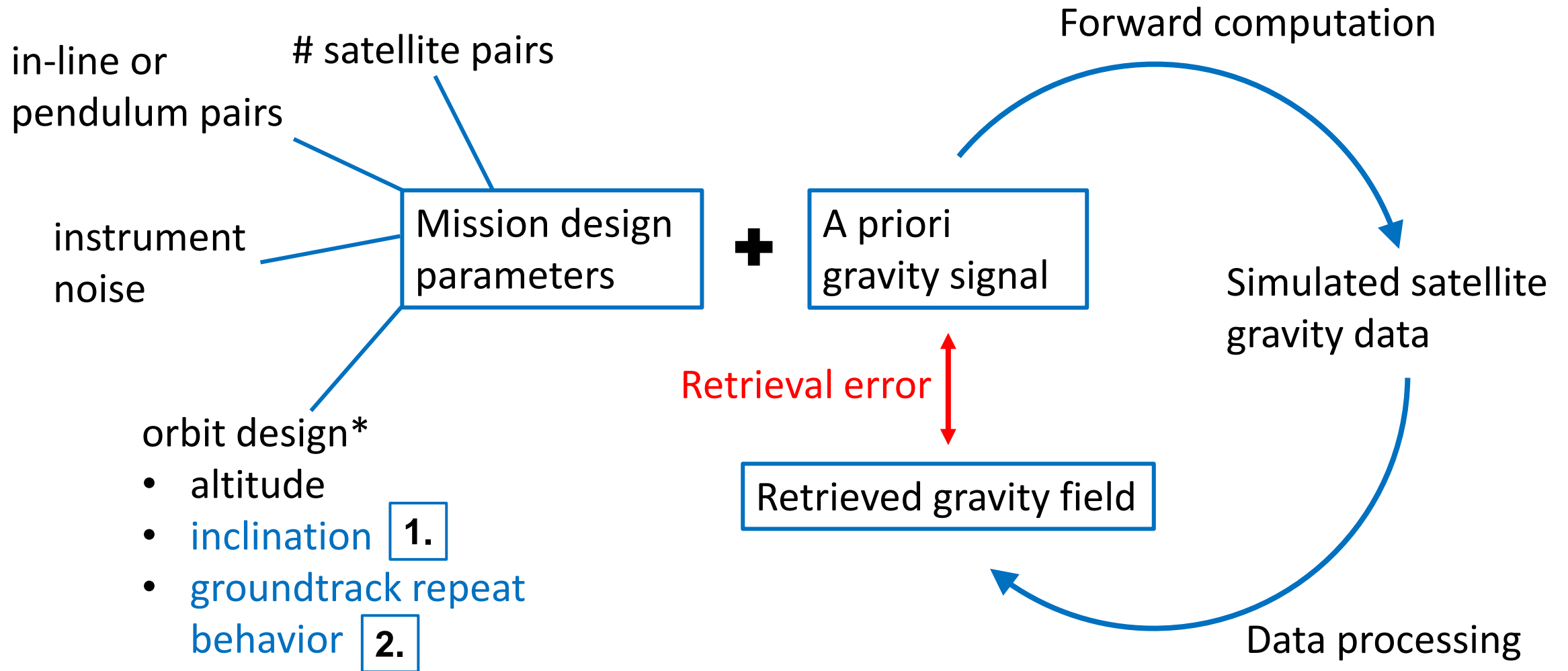


Mass-change And Geosciences International Constellation (**MAGIC**)



Aim of the ESA Science Support Study for MAGIC:

- Evaluating the **performance** of various satellite gravity **mission designs**



*Massotti et al. (2021)

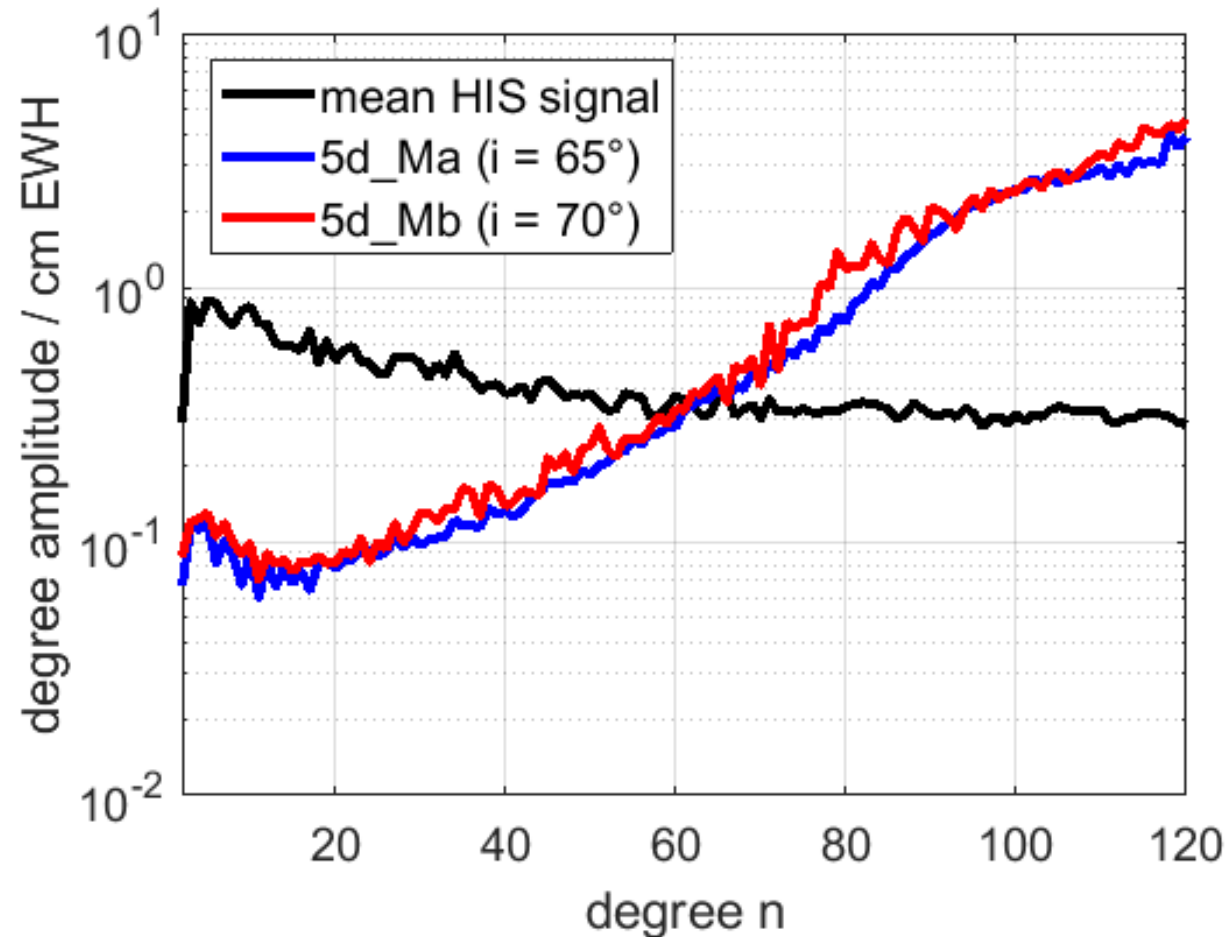
1. Impact of the inclination of the second pair



What is the **impact of the inclination of the second satellite pair** on the retrieval errors of a double-pair mission?

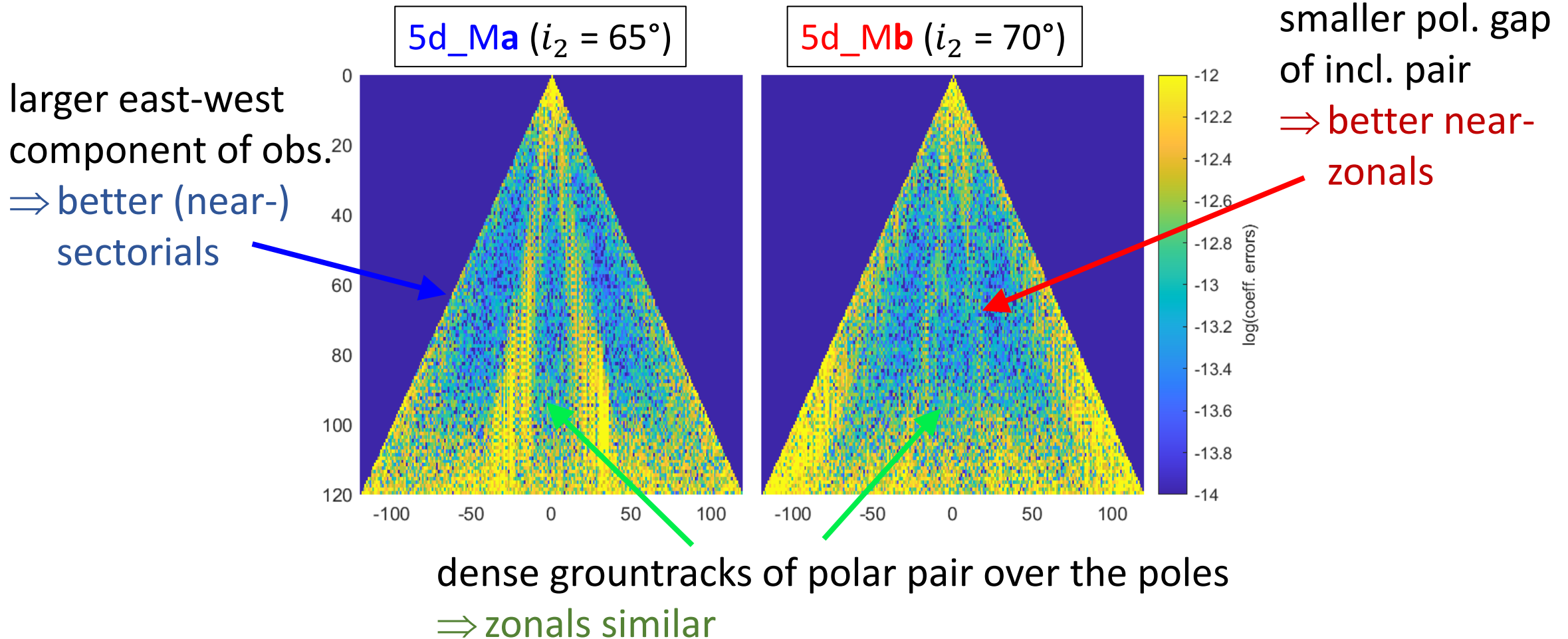
Two **double-pair scenarios** with similar altitude
(31-day simulation incl. full signal + noise):

Global performance:



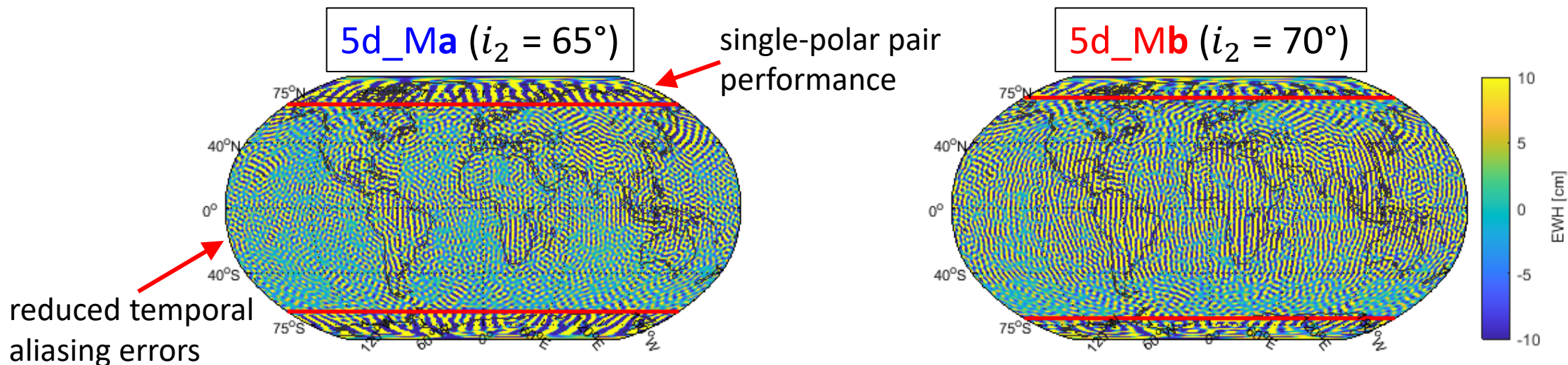
1. Impact of the inclination of the second pair – Unfiltered results

Two **double-pair scenarios** with similar altitude
(31-day simulation incl. full signal + noise):

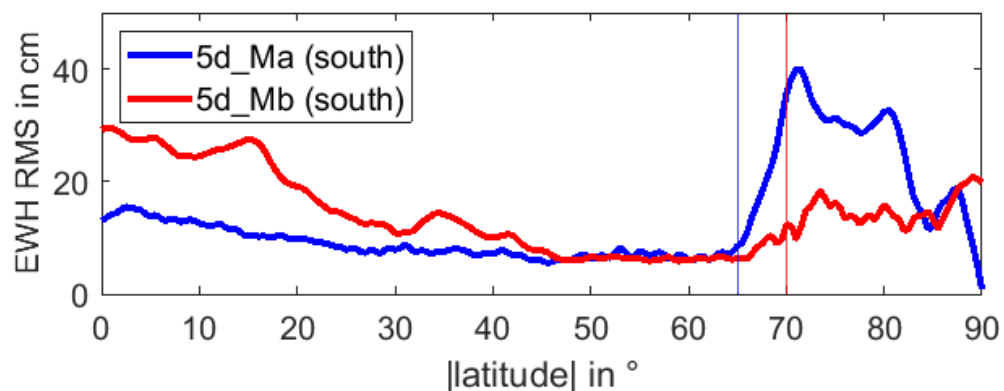


1. Impact of the inclination of the second pair – Unfiltered results

EWH grid differences to true HIS signal:



RMS of errors along parallels:

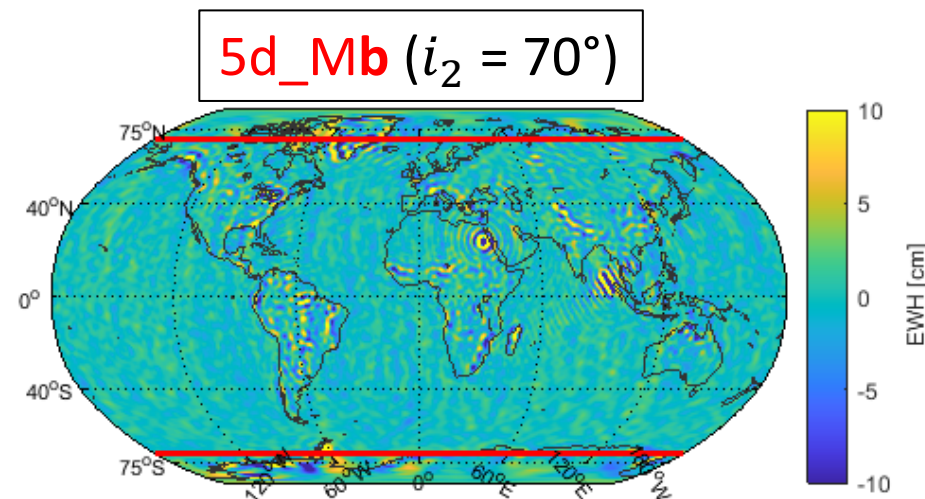
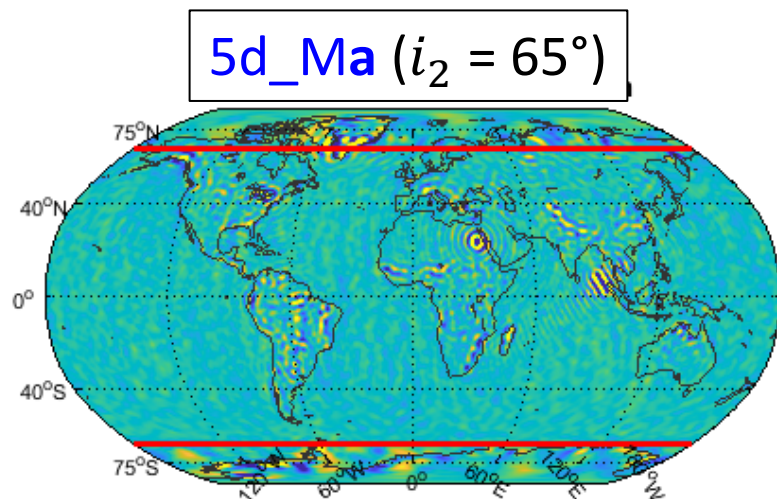


Area-weighted mean of rms errors in cm:

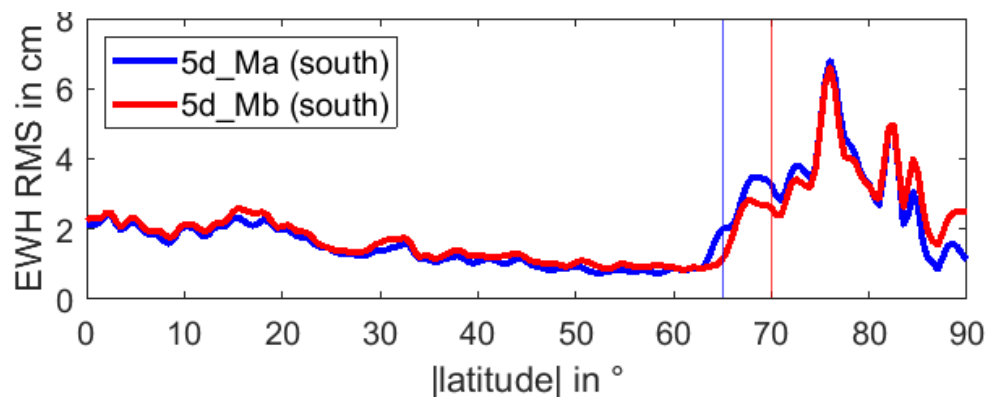
	southern hemisphere			global
	all φ	$ \varphi < 65^\circ$	$ \varphi > 70^\circ$	all φ
5d_Ma (65°)	12.56	9.76	30.94	14.81
5d_Mb (70°)	17.79	18.25	14.44	18.71
	65° case 29% better	65° case 47% better	70° case 53% better	65° case 21% better

1. Impact of the inclination of the second pair – VADER-filtered results

After VADER filtering, $x_\alpha = (N + \alpha M)^{-1} N x = W_\alpha x \dots$



RMS of errors along parallels:



Area-weighted mean of rms errors in cm:

	southern hemisphere			global
	all φ	$ \varphi < 65^\circ$	$ \varphi > 70^\circ$	all φ
5d_Ma (65°)	1.89	1.59	4.02	2.61
5d_Mb (70°)	1.94	1.72	3.88	2.64
	65° case 3% better	65° case 8% better	70° case 3% better	65° case 1% better

1. Impact of the inclination of the second pair – Conclusions

Impact of a **smaller inclination i_2** of the second pair:



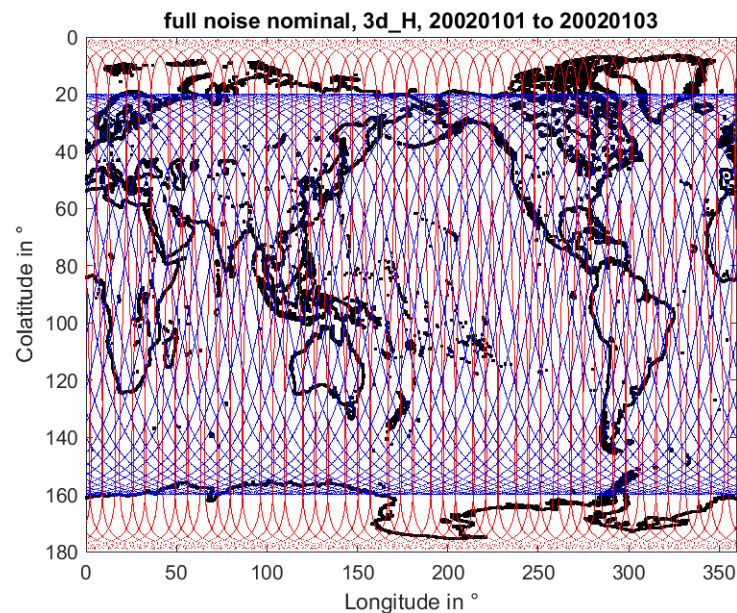
Unfiltered solutions:

- larger east-west component of observations
⇒ **reduction of sectorial noise** for $|\varphi| < i_2$
- larger polar gaps of inclined pair
⇒ stronger **noise over poles**
- ⇒ larger unobserved region near the poles as stand-alone mission

VADER-filtered solutions:

- impact of i_2 smaller
- but: need weaker filter if i_2 is smaller

2. Orbit design aspects for 3-day gravity retrieval



How do the **orbits** of polar and inclined pair need to be designed for **short-term (3-day) gravity retrieval**?

2. Orbit design aspects for 3-day gravity retrieval

3d_H orbits: both polar and inclined pair have a

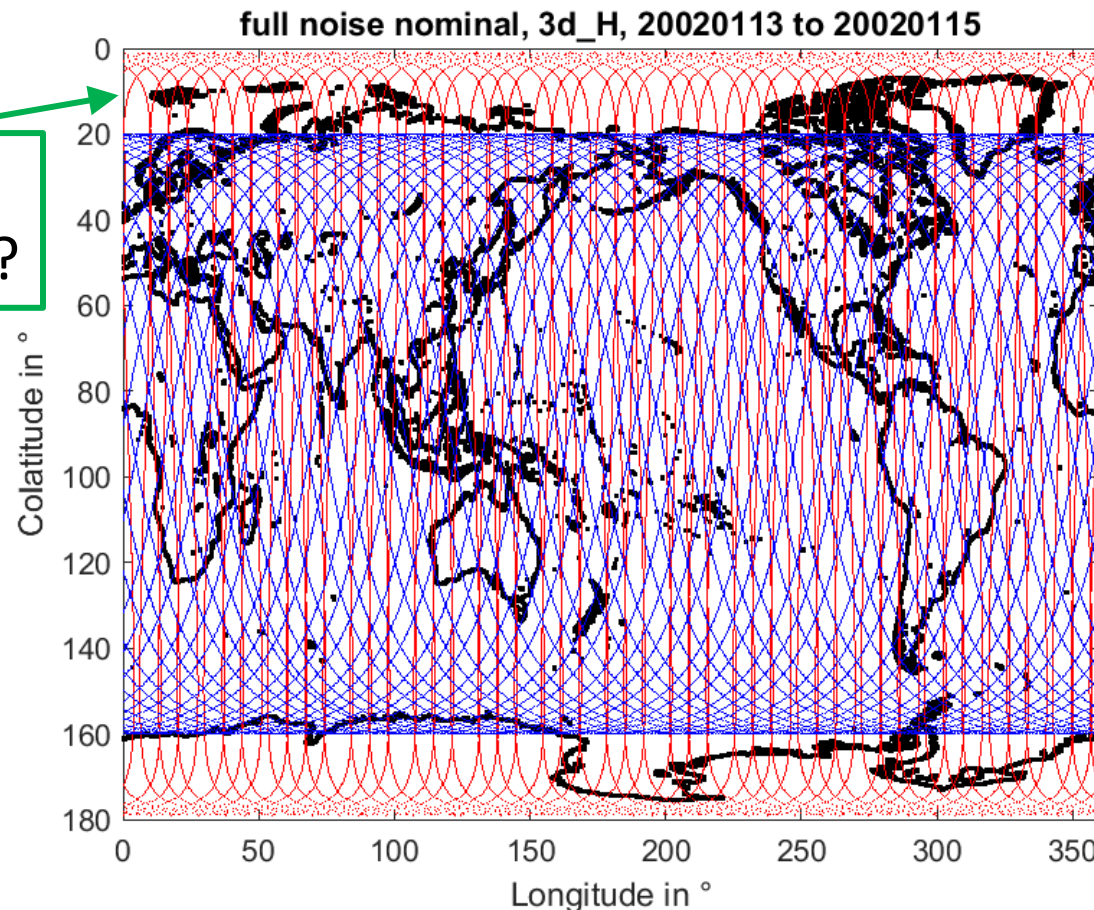
- **3-day subcycle**
- with a common longitudinal **drift rate**

$-3^\circ/3 \text{ days}$



Aspect 1:

Impact of 3-day subcycle of **polar pair**?

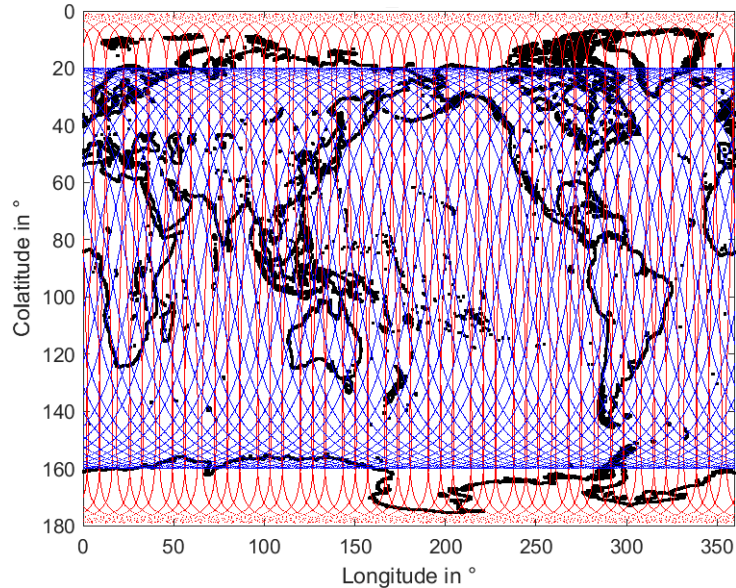


Aspect 2:

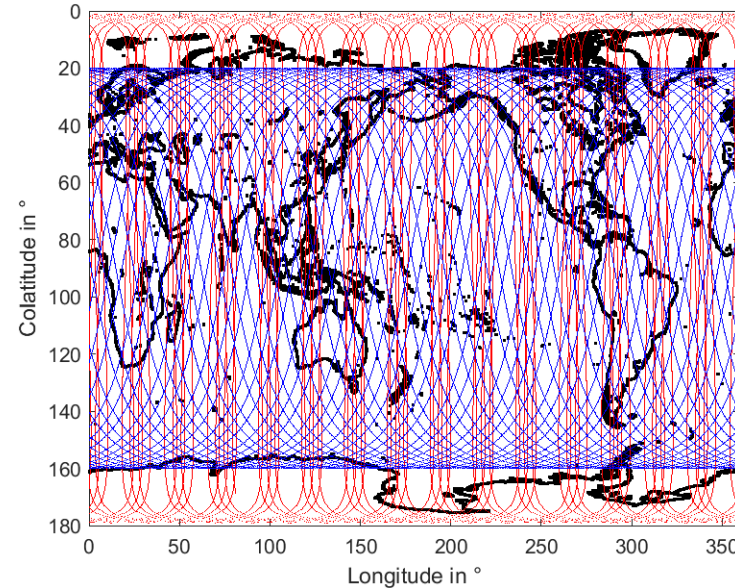
Impact of 3-day subcycle of **inclined pair**?

Subcycles of satellite orbits - test scenarios

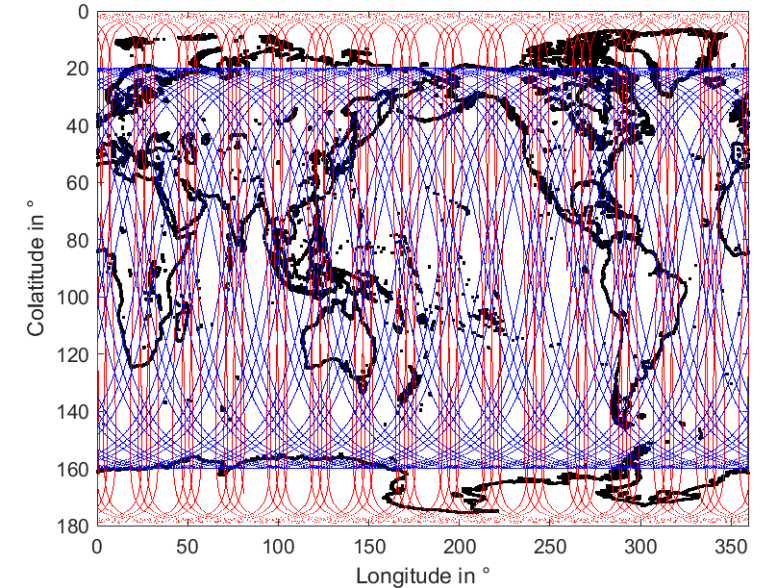
3d_H



U3d5d_H



U5d_H



PP: 3-day subcycle

IP: 3-day subcycle

Aspect 1:

Impact of 3-day subcycle of polar pair?

PP: 5-day subcycle

IP: 3-day subcycle

Aspect 2:

Impact of 3-day subcycle of inclined pair?

PP: 5-day subcycle

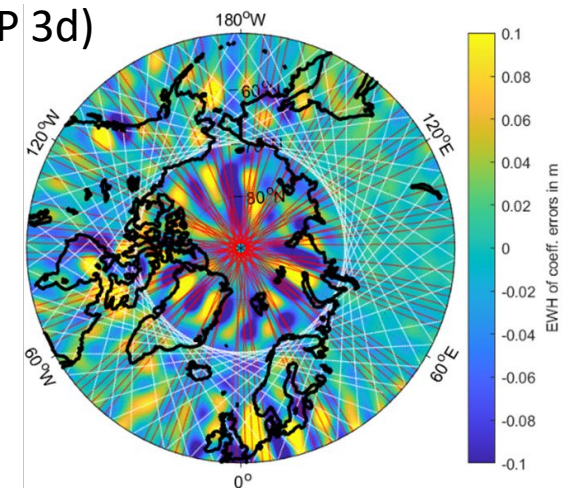
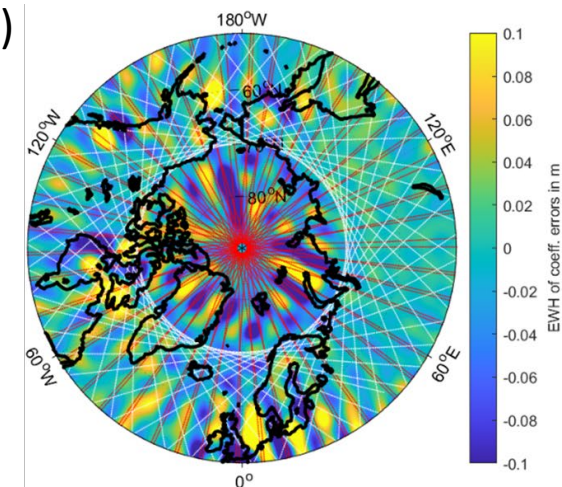
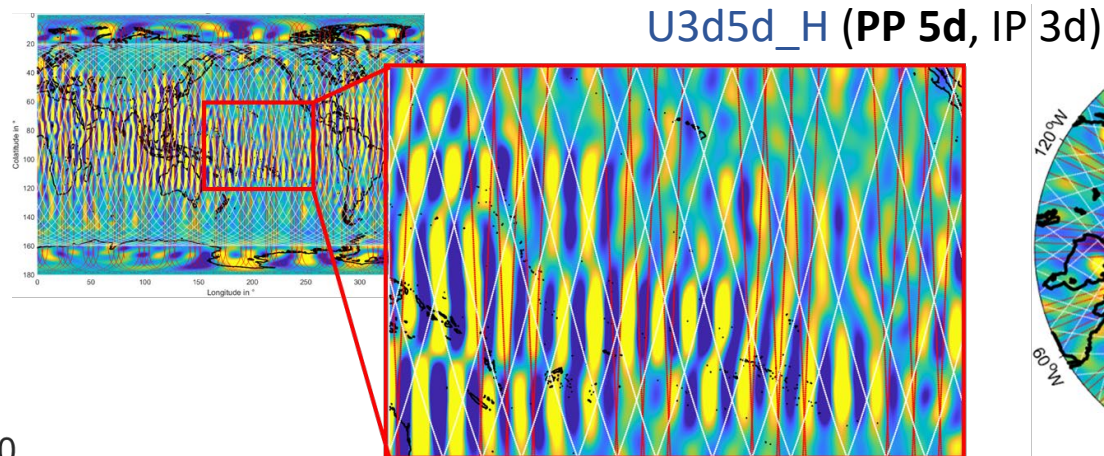
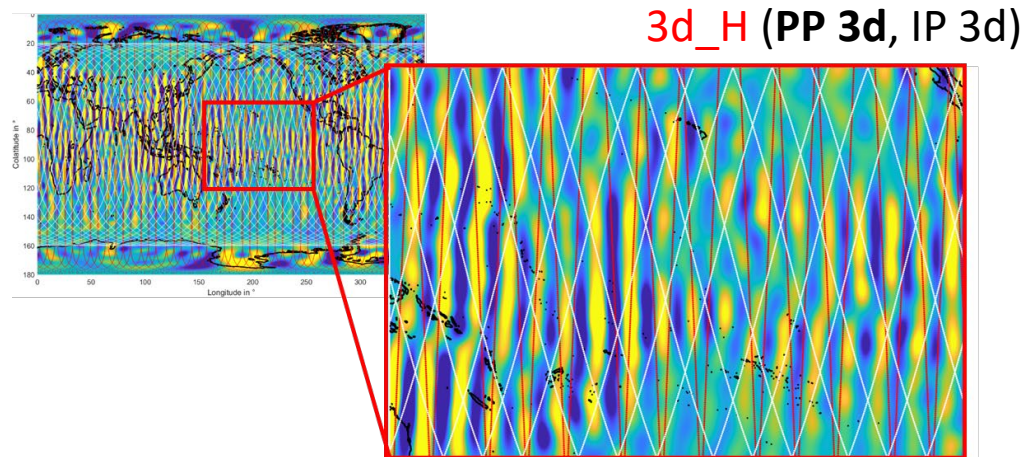
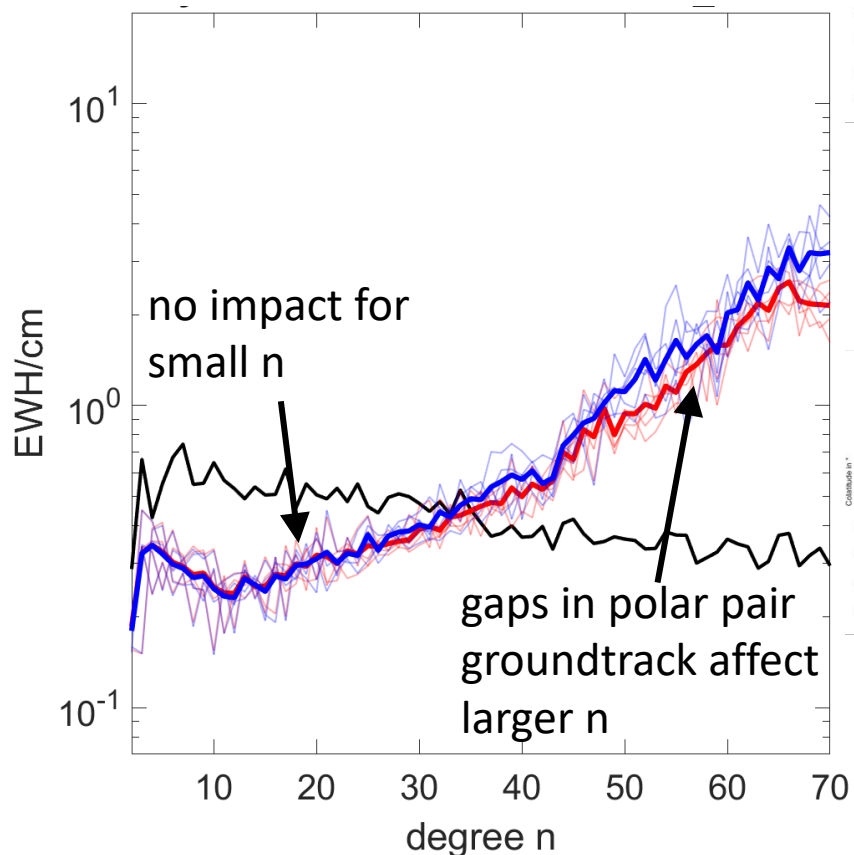
IP: 5-day subcycle

2. Orbit design aspects for 3-day gravity retrieval

Aspect 1: Impact of 3-day subcycle of polar pair

3d_H (PP 3d, IP 3d)

U3d5d_H (PP 5d, IP 3d)



low lat. => more stripes in groundtrack gaps

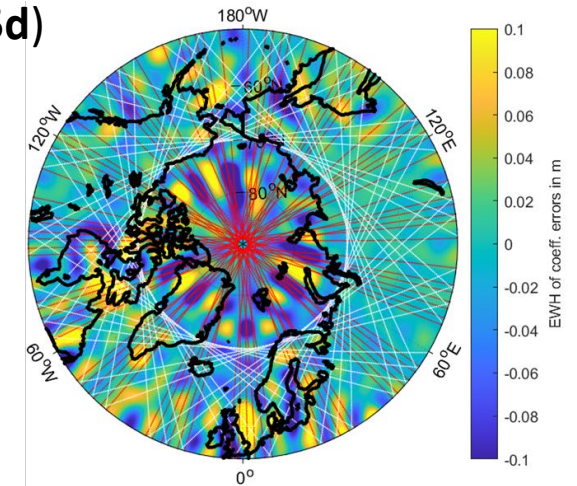
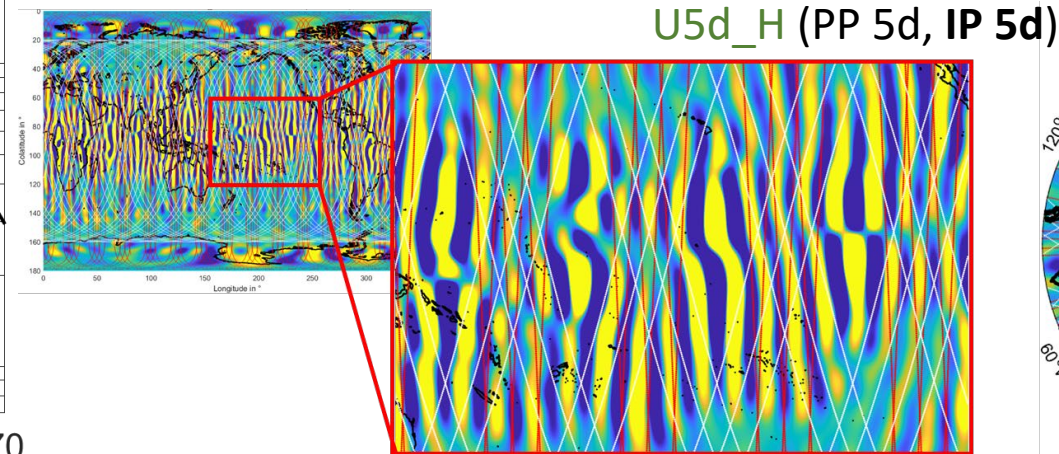
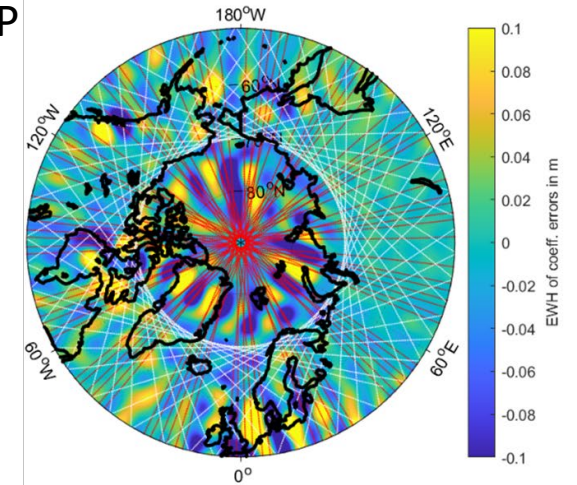
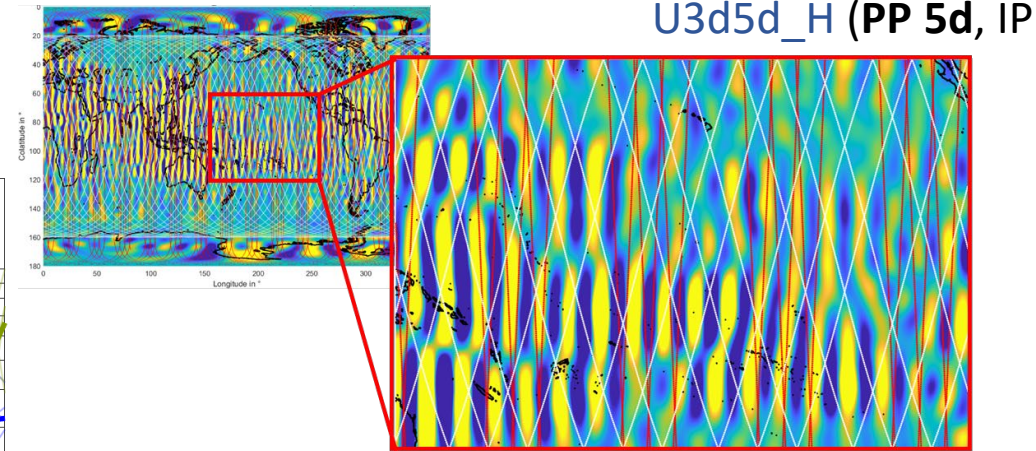
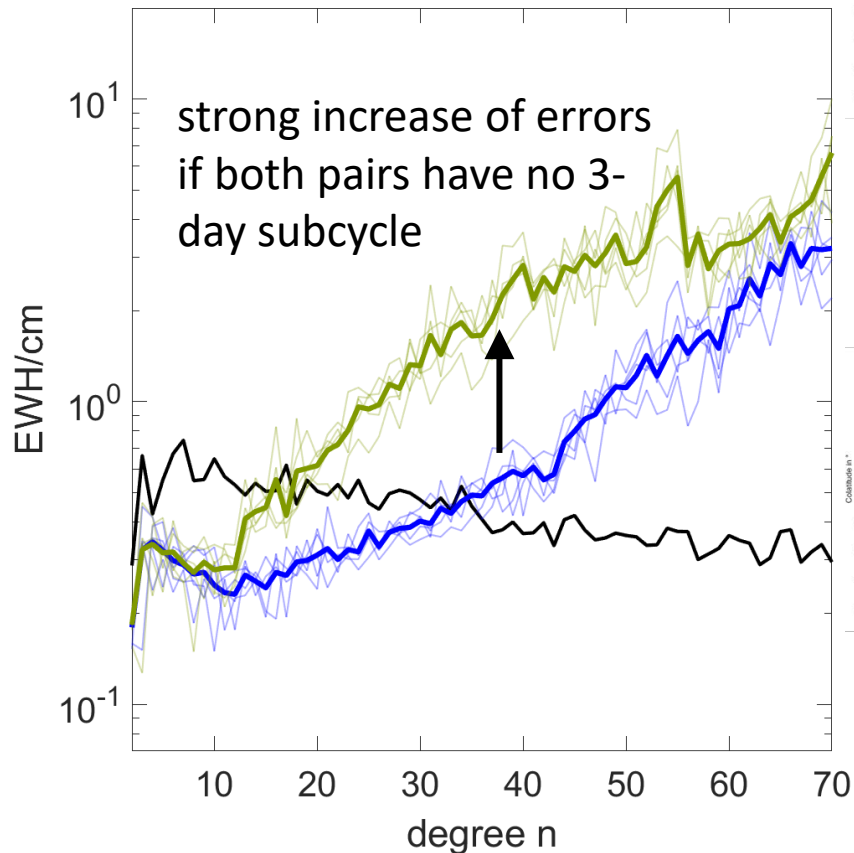
poles: smaller impact

2. Orbit design aspects for 3-day gravity retrieval

Aspect 2: Impact of 3-day subcycle of **inclined pair**

U3d5d_H (PP 5d, IP 3d)

U5d_H (PP 5d, IP 5d)



strong noise correlated with groundtrack gaps

poles: no impact

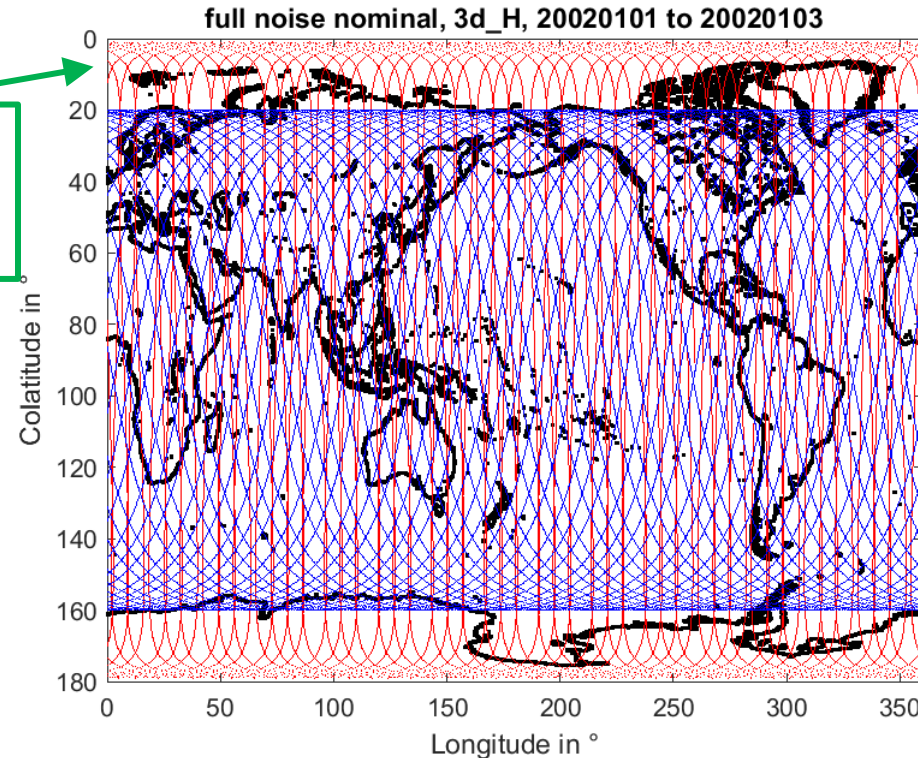
2. Orbit design aspects for 3-day gravity retrieval – Conclusions

Impact of orbit subcycles on short-term gravity retrieval:

Aspect 1:

Impact of 3-day subcycle of **polar pair**?

- missing 3-day subcycle of polar pair acceptable, but resolution of **small-scale features at low latitudes** is affected



Aspect 2:

Impact of 3-day subcycle of **inclined pair**?

- if polar pair orbit is uncontrolled, **orbit control for the inclined pair** to maintain the groundtrack subcycles is **crucial** for the homogeneous quality of **short-term solutions**

⇒ Match between **subcycle length** and **retrieval period** affects short-term solutions!

1. Impact of the inclination of the second pair

Smaller inclination i_2 of the second pair

⇒ reduction of sectorial noise for $|\varphi| < i_2$

⇒ increase of noise over poles

Trade-off between improvement at low latitudes and increasing polar gap

2. Orbit design aspects for 3-day gravity retrieval

Non-homogeneous ground track pattern...

... of the polar pair

⇒ increased striping for low latitudes

... of the inclined pair

⇒ severe degradation over whole spectral range

For short-term solutions, match between orbit subcycles and retrieval period should be considered

- Orbit specifications:

Massotti, L.; Siemes, C.; March, G.; Haagmans, R.; Silvestrin, P. Next Generation Gravity Mission Elements of the Mass Change and Geoscience International Constellation: From Orbit Selection to Instrument and Mission Design. *Remote Sens.* 2021, 13, 3935. <https://doi.org/10.3390/rs13193935>

- MAGIC Mission Requirements Document:

Next Generation Gravity Mission as a Mass change And Geosciences International Constellation (MAGIC). A joint ESA/NASA double-pair mission based on NASA's MCDO and ESA's NGGM studies. Mission Requirements Document. ESA, Earth and Mission Science Division.

Appendix

Table 3. Candidate orbit sets for inclined and polar pairs recommended for further investigation. The ID shows the number of subcycle days for which the set is optimized as a first step and additional information about the altitudes: mid (M) and high (H).

ID	Sats 1 (IP)		Sats 2 (PP)		$h_{l,1}$ (-)	$h_{l,2}$ (-)	$\Delta\lambda_{shift,1}$ (deg)	$\Delta\lambda_{shift,2}$ (deg)	Subcycle (days)
	Alt. (km)	Incl. (deg)	Alt. (km)	Incl. (deg)					
3d_M	409	70	440	89	1.368	1.383	2.308	2.384	2, 3, 8, 11, 30
3d_H	432	70	463	89	1.451	1.449	-3.076	-3.067	3, 7, 31
5d_Ma	396	65	434	89	1.397	1.383	-1.499	-1.458	2, 3, 5, 13, 18, 31
5d_Mb	397	70	425	87	1.168	1.167	0.736	0.733	2, 5, 27, 32
5d_H	465	75	488	89	1.185	1.190	0.762	0.781	4, 5, 29
7d_M	389	70	417	87	1.238	1.253	0.743	0.786	2, 7, 30
7d_H	432	70	463	89	1.218	1.226	0.672	0.692	3, 7, 31

Massotti et al. (2021)

Post-processing of the retrieved fields

- VADER filtering (time-variable decorrelation, Horvath et al (2018)):

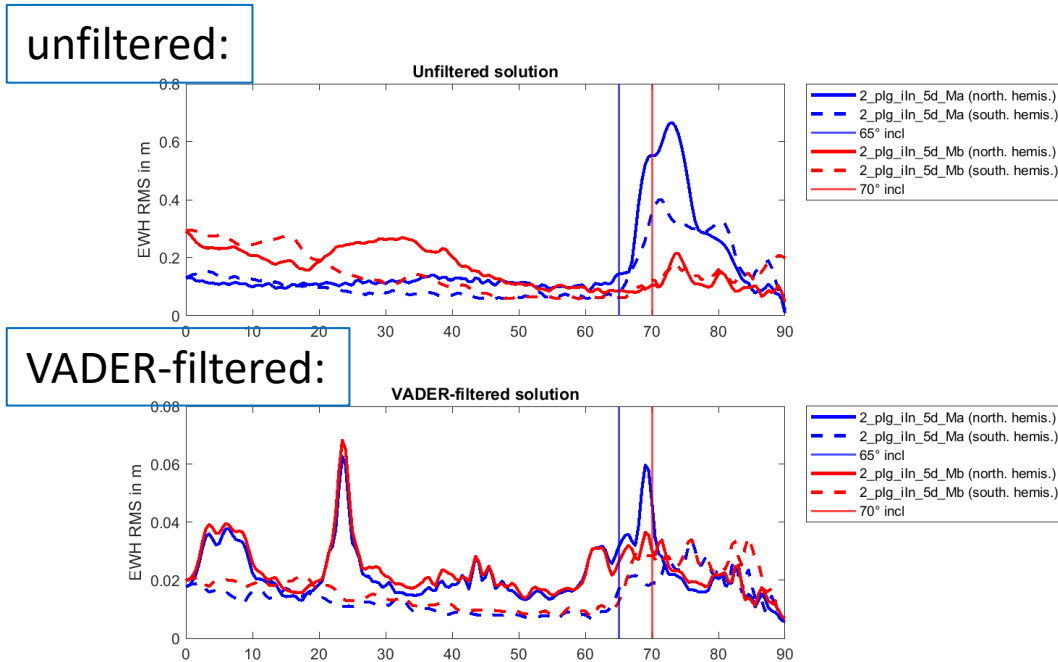
$$x_{\alpha} = (N + \alpha M)^{-1} N x = W_{\alpha} x$$

- x_{α} : vector of filtered SH coefficients
- x : vector of unfiltered SH coefficients
- N : NEQ-Matrix ($A^T P A$) (= inverse of error VCV)
- M : inverse of signal variance matrix (computed from monthly HIS signal)
- α : scaling factor (determined such that RMS of global EWH errors is minimized)
- W_{α} : filter matrix

1. Impact of the inclination of the second pair – VADER-filtered results

Impact of VADER filtering:

EWH RMS values along parallels:

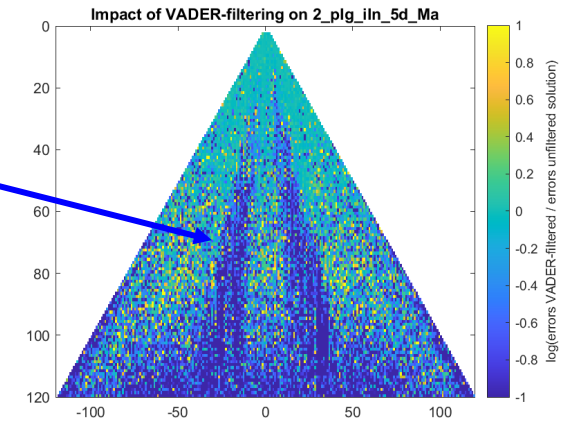


- errors reduced and **less latitude-dependent**
- difference 5d_Ma vs. 5d_Mb reduced
- impact of filtering larger in **areas of larger errors**

Errors(filtered sol.) / errors(unfiltered sol.):

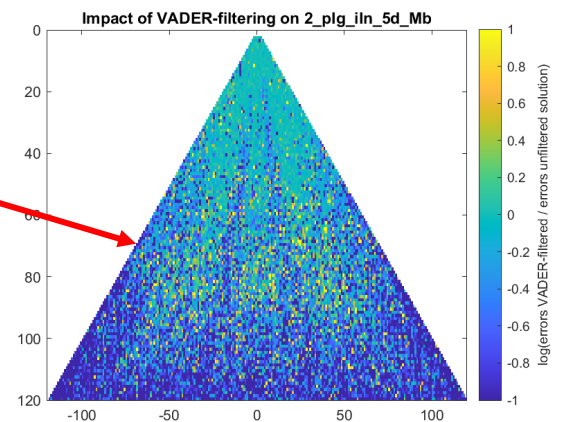
5d_Ma ($i_2 = 65^\circ$)

- improvement of near-zonal polar gap coefficients



5d_Mb ($i_2 = 70^\circ$)

- improvement of sectorials



- improvement largest for coefficients of larger errors

2. Orbit design aspects for 3-day gravity retrieval

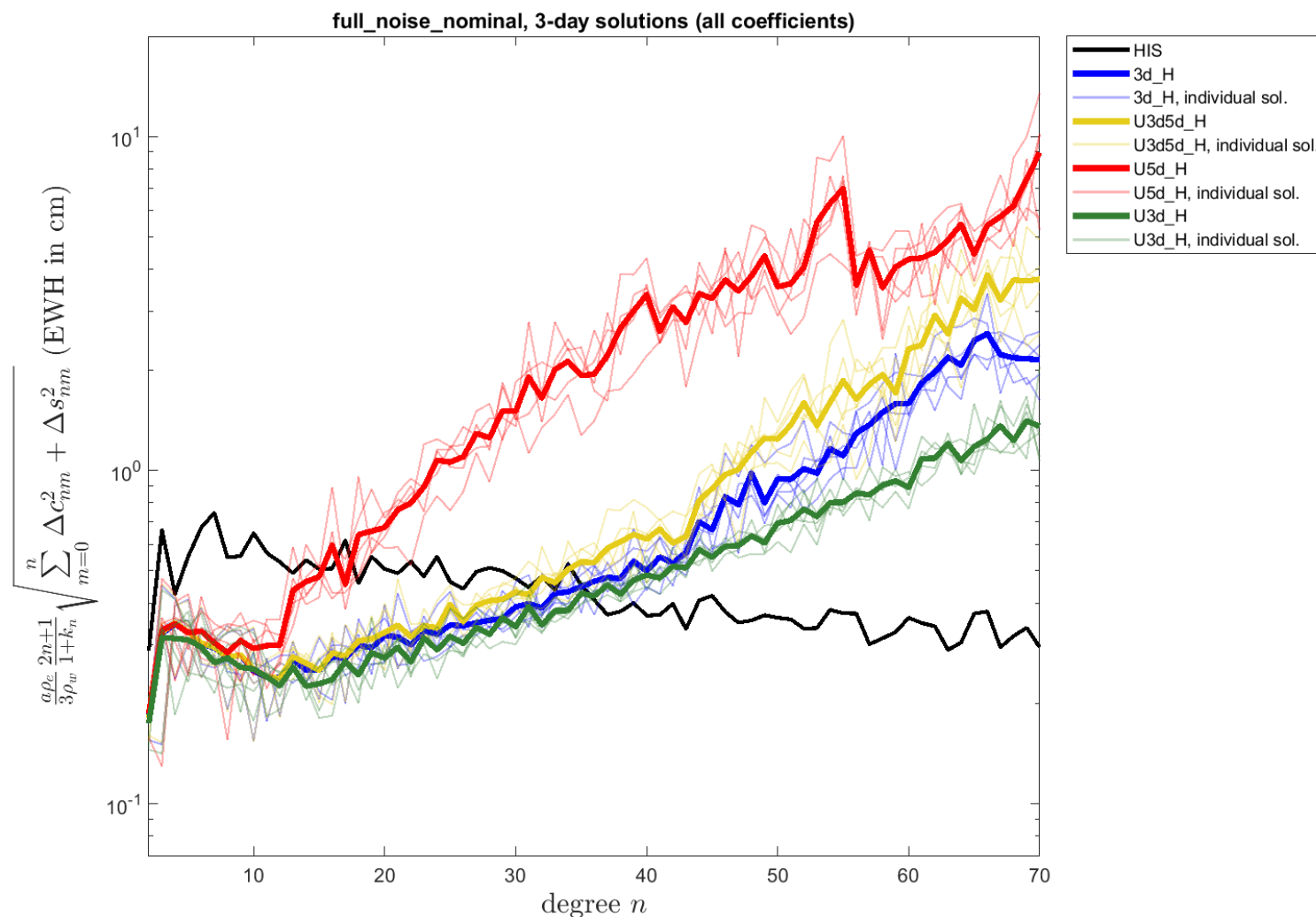
Subcycles of satellite orbits - test scenarios

- **double-pair low-low satellite-to-satellite constellations**
 - polar pair: 89° inclination
 - inclined pair: 70° inclination
- **numerical closed-loop simulations** including
 - AO model errors
 - ocean tide model errors
 - HIS signal
 - orbit, ACC, LRI sensor noise
- **ACC noise** assumption:
 - polar pair: GRACE-type noise ($\sim 10^{-10}$ m/s²/√Hz)
 - inclined pair: NGGM-type noise ($\sim 10^{-11}$ m/s²/√Hz)
- 5 subsequent d/o 120 **3-day solutions**

2. Orbit design aspects for 3-day gravity retrieval

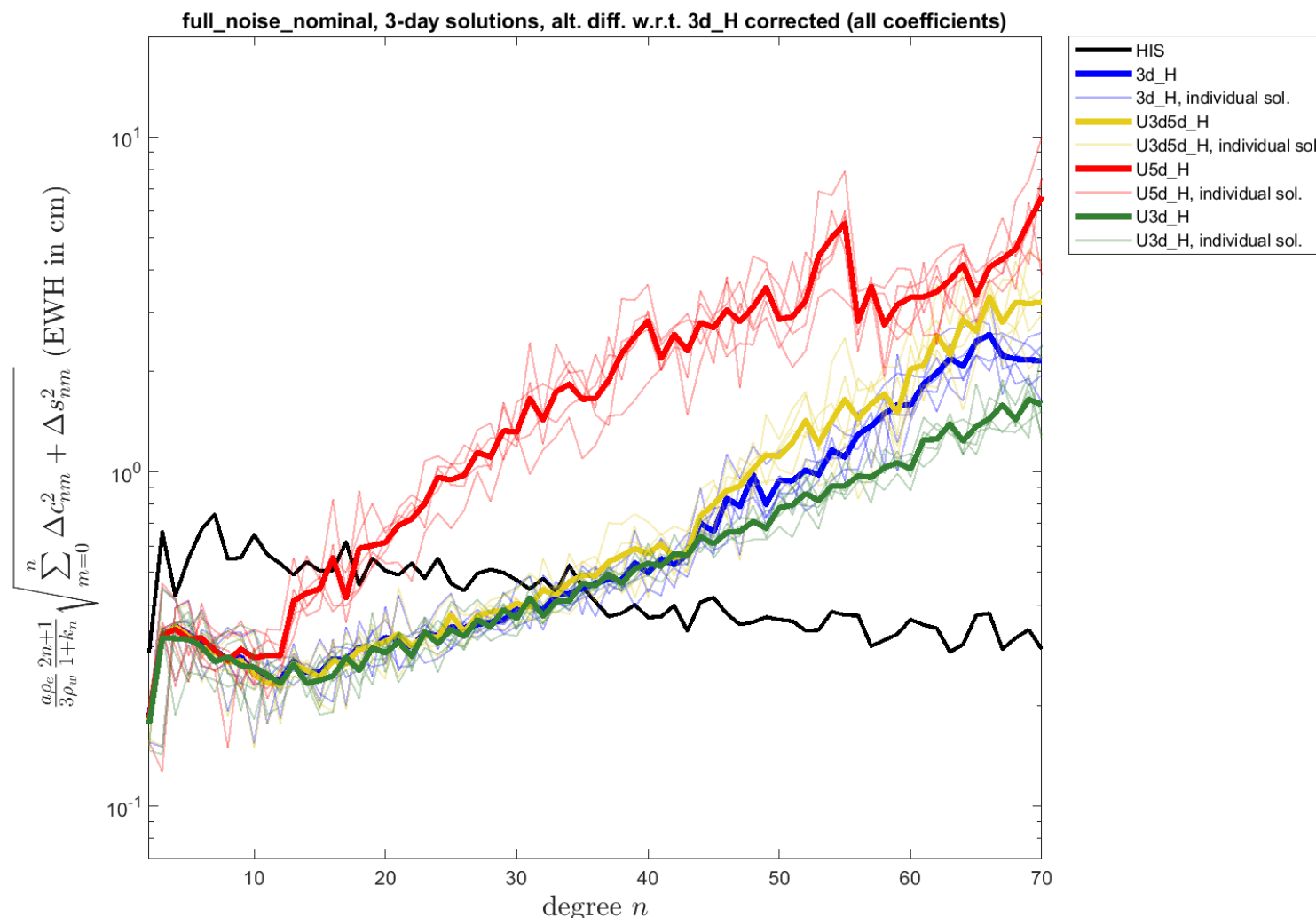
ID	Sats 1 (IP)		Sats 2 (PP)		h_{l_1} [-]	h_{l_2} [-]	$\Delta(\text{Lon})_1$ [deg]	$\Delta(\text{Lon})_2$ [deg]	Sub-cycles [days]	Retrieval period [days]	MRD
	Alt. [km]	Incl. [deg]	Alt. [km]	Incl. [deg]							
Coordinated											
3d_H	432	70	463	89	1.451	1.449	-3.076	-3.067	3, 7, 31	3	MRD-040 – Fulfilled MRD-050 – Fulfilled MRD-060 – Fulfilled
Uncoordinated											
U3d5d_H	432	70	492	89	1.451 (3d)	1.172 (5d)	-3.076 (3d)	-0.790 (5d)	IP: 3, 31 PP: 5, 31	3	MRD-040 – Violated MRD-050 – Violated MRD-060 – Violated
U5d_H	460	70	492	89	1.061 (5d)	1.172 (5d)	-0.284 (5d)	-0.790 (5d)	IP: 5 PP: 5, 31	3	MRD-040 – Violated MRD-050 – Violated MRD-060 – Violated
U3d_H	402	65	463	89	1.382	1.449	2.380	-3.067	IP: 3, 29-30 PP: 3, 7, 31	3	MRD-040 – Fulfilled MRD-050 – Violated MRD-060 – Violated

Degree amplitudes (not corrected for altitude):



2. Orbit design aspects for 3-day gravity retrieval

Degree amplitudes (corrected for altitude):



Altitude correction relative to 3d_H
(here: in blue)

- Apply degree-dependent factor $\left(\frac{r_{ref}}{r}\right)^{n+2}$ to the SH coefficients
- r : mean radius of polar and inclined pair of the considered scenario
- r_{ref} : mean radius of polar and inclined pair of scenario 3d_H
- Background: measurement quantity is range rate, related to gravity accelerations, proportional to $\left(\frac{1}{r}\right)^{n+2}$
- Effect: scenarios U3d_H, 3d_H, U3d5d_H move together for $n < 40$

2. Orbit design aspects for 3-day gravity retrieval

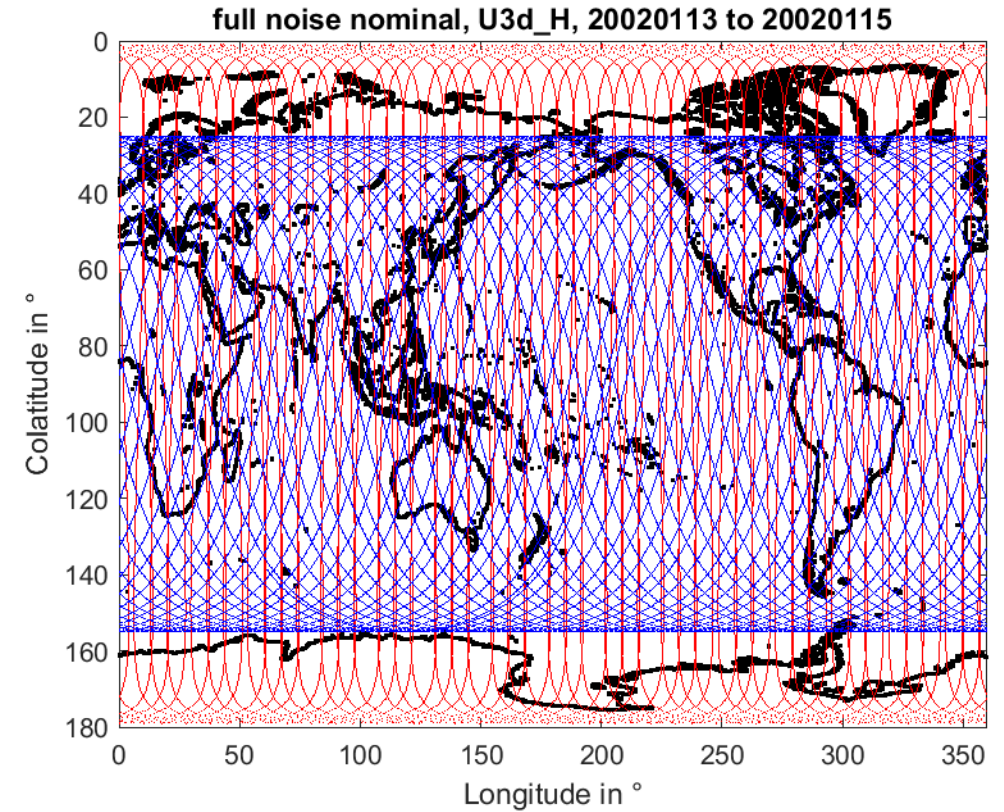
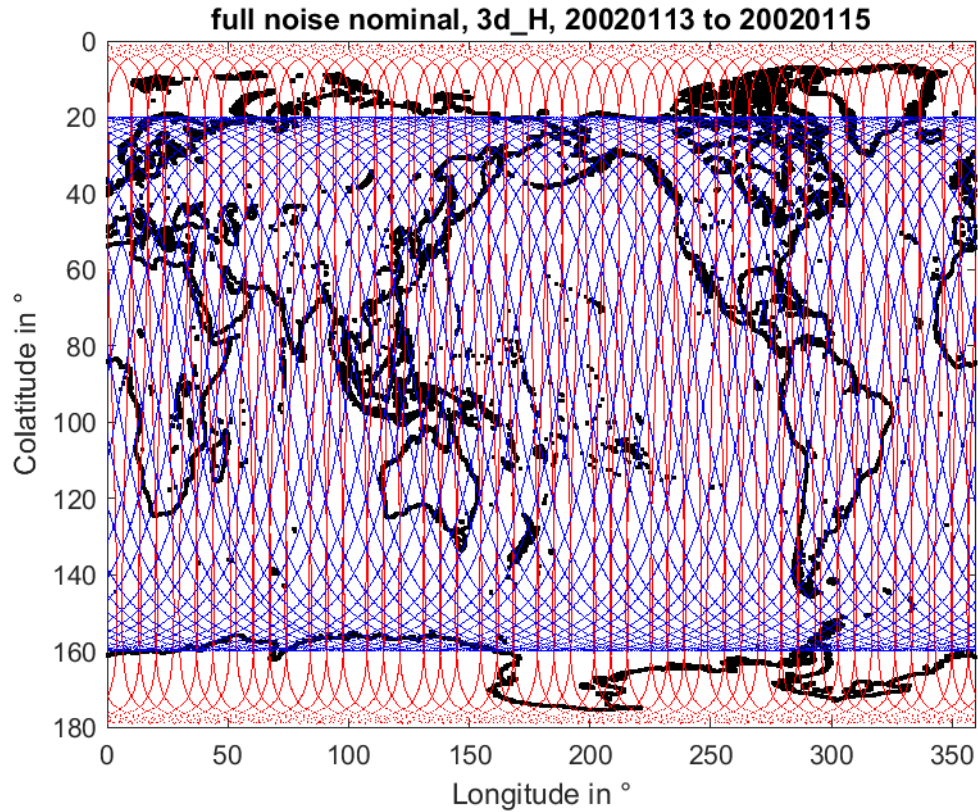
Aspect 3: Impact of common drift rate of polar and inclined pair

3d_H (PP 3d, IP 3d)

 -3°/3 days

U3d_H (PP 3d, IP 3d)

 -3°/3 days
 2°/3 days

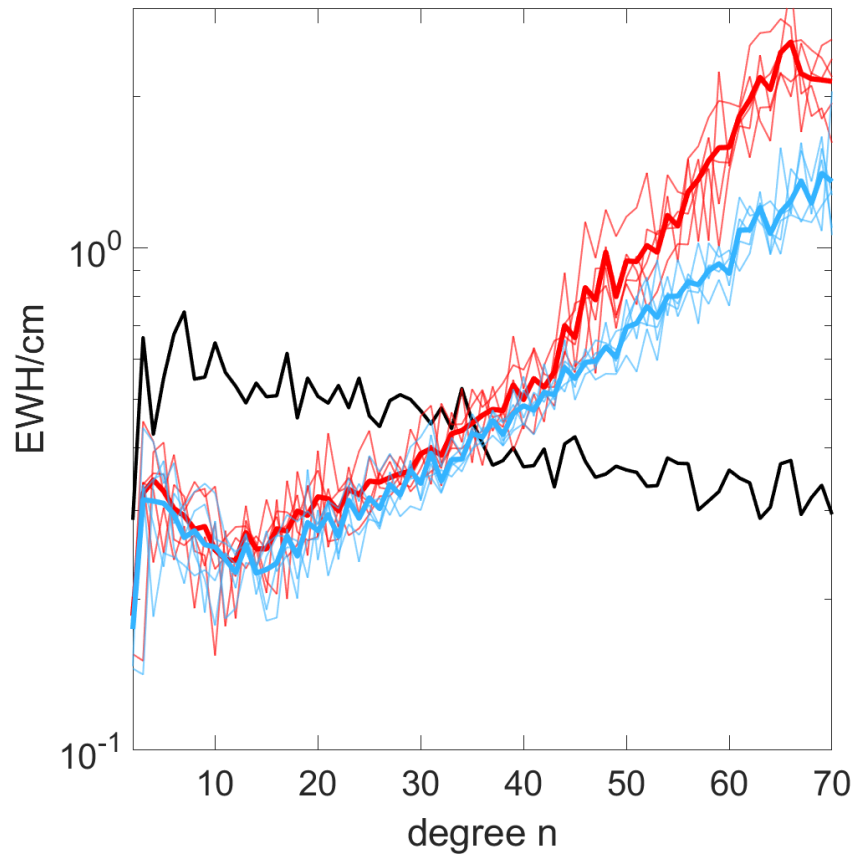


2. Orbit design aspects for 3-day gravity retrieval

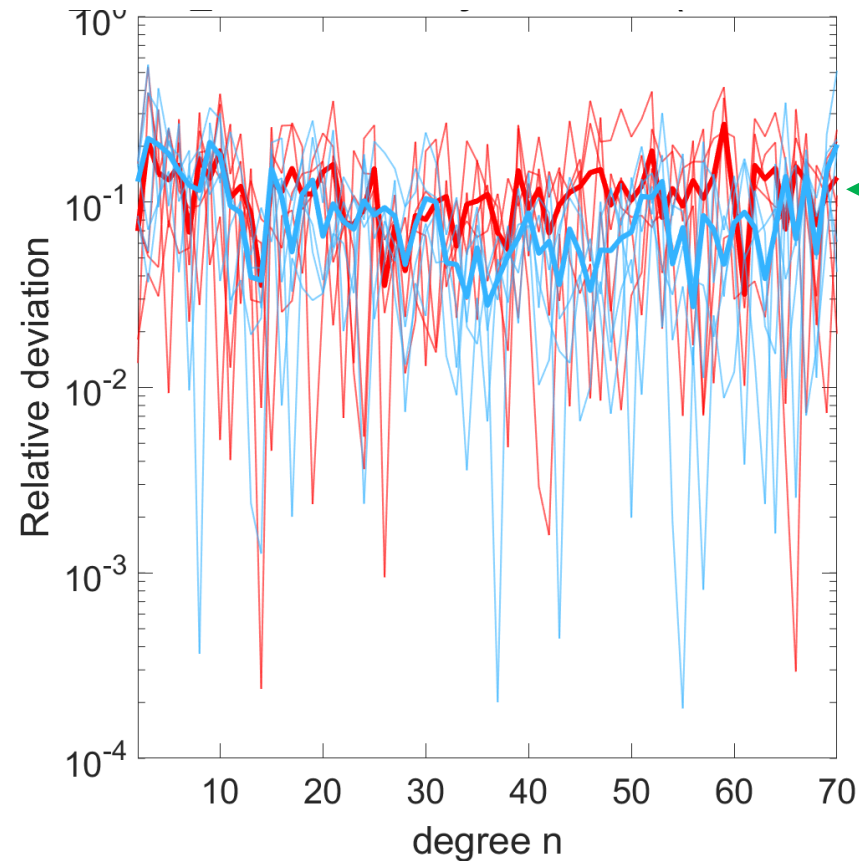
Aspect 3: Impact of **common drift rate** of polar and inclined pair

3d_H (common drift rate)

U3d_H (uncommon drift rate)



- thin: error of individual 3-day solutions
- thick: mean of thin curves



- thin: relative deviation of ind. curves from their mean
- thick: mean of thin curves

no impact of
(un)common drift on
error spread of
subsequent 3-day
solutions visible