

Review of MARVEL pre-Phase-A

*Jean-Michel Lemoine¹, Mioara Mandaia¹, Louise Lopes¹,
Benoit Meyssignac¹, Alejandro Blazquez¹, Georges
Balmino¹, Vincent Costes¹, Sean Bruinsma¹, Etienne
Samain², Stéphane Bourgogne³*

¹ CNES ; ² SigmaWorks ; ³ Stellar Space Studies



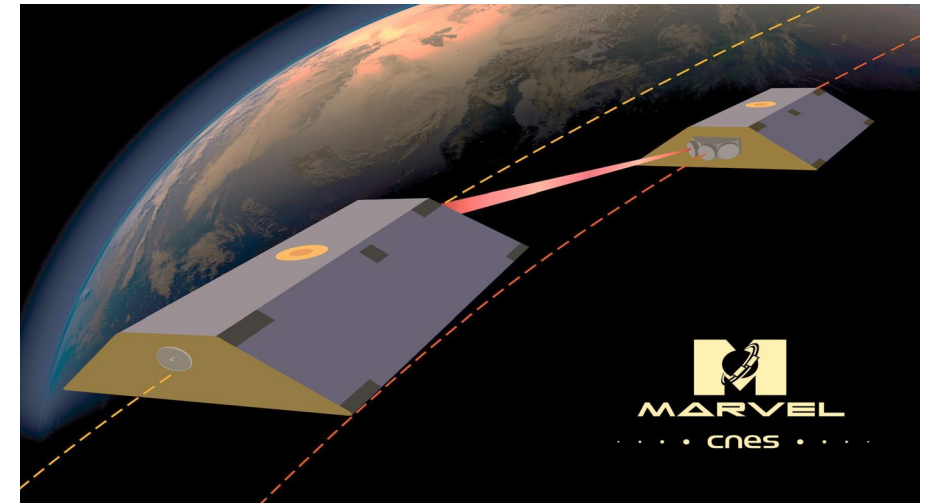
THE MARVEL TEAM



Mioara Manda (CNES)
Solid Earth Program Manager



Louise Lopes (CNES)
Pre-Phase-A leader



Benoit Meyssignac
(CNES/LEGOS)



Alejandro Blazquez (CNES/LEGOS)
Scientific advisers



Jean-Michel Lemoine
(CNES/GET)



Etienne Samain (SigmaWorks)
Instrument experts



Vincent Costes (CNES)



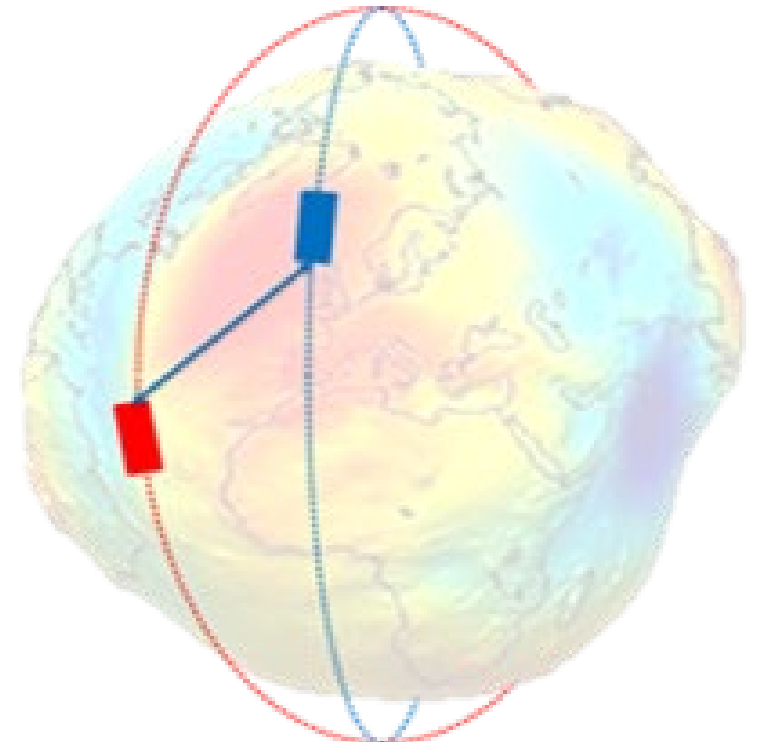
Stéphane Bourgoigne (Stellar Space Studies)
Numerical Simulations expert



Georges Balmino (CNES Emeritus)
Member of the science advisory group

- ❖ Context
- ❖ The laser ranging instrument
- ❖ The numerical simulations of scientific performance
- ❖ Conclusion

- The MARVEL mission proposal for an **improved observation of the time variable gravity field** was submitted to the 2019 CNES Scientific Prospective Seminar and was accepted in September 2019
- The pre-Phase-A study started in January 2020 and ended in February 2022
- The principle of the MARVEL concept is a **pendulum** configuration with 2 (or more) low flying polar satellites
- In pendulum configuration, the 2 satellites are on two similar polar orbital planes, with a slight offset in ascending node and mean anomaly
- The measurements done between the satellites are therefore oriented alternatively to the right and to the left of the orbital track, up to $\pm 45^\circ$
- After a few days, the determination of the gravity field from those measurements becomes almost **isotropic**

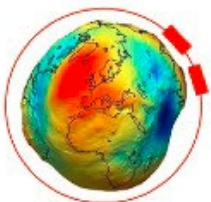


➤ The key to improving the gravity field observation is improving the **geometry** of the observations

CURRENT CONCEPT (IN-LINE POLAR PAIRS)

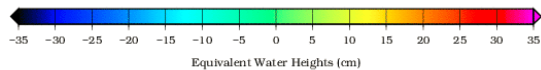
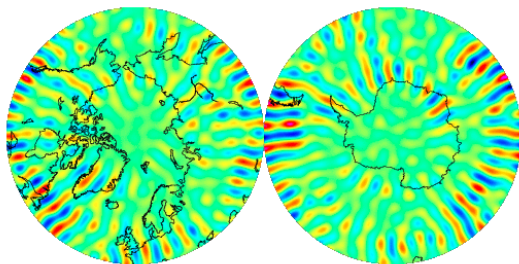
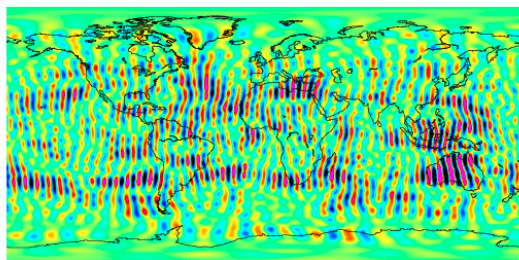
FUTURES CONCEPTS (INCLINED MEASUREMENTS)

GRACE

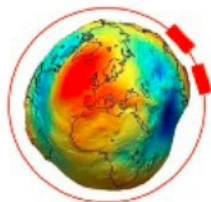


σ SST = 1.e-7 m/s

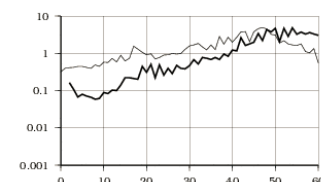
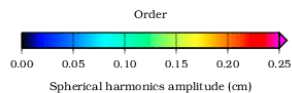
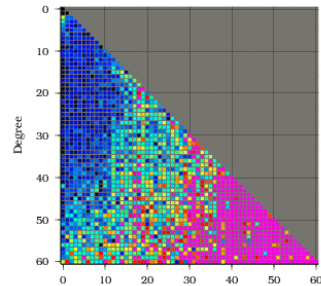
GRACE | Error | Degree 2 to 60 | 200601
 GRACE-A | 3h 20% 0.1mc/s 3mm 3e-10m/s² 0E 31K 31X
 GRACE-B | 3h 20% 0.1mc/s 3mm 3e-10m/s² 0E 31K 31X
 min -72.50 cm / max 79.47 cm / rms 14.52 cm / ocn 14.36 cm / ctn 14.78 cm



GRACE-FO

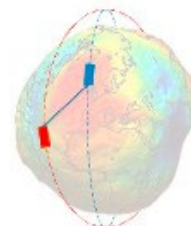


σ SST = 2.5e-9 m/s



Spectrum by degree (bold) and order (cm)

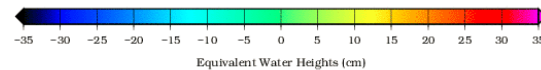
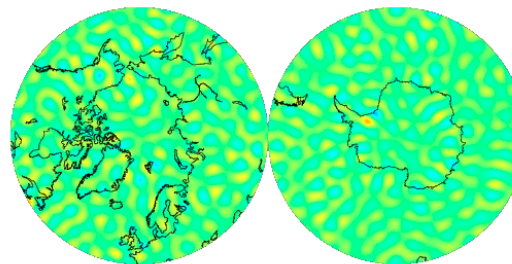
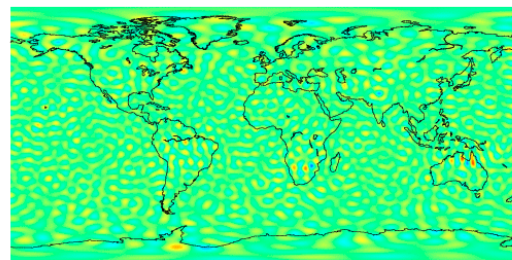
MARVEL



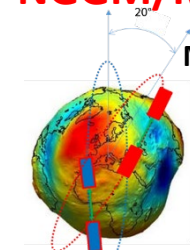
Maximum angle
45°

σ SST = 1.e-7 m/s

PENDULUM | Error | Degree 2 to 60 | 200601
 GRACE-A | 3h 20% 0.1mc/s 3mm 3e-10m/s² 0E 31K 31X
 GRACE-P | 3h 20% 0.1mc/s 3mm 3e-10m/s² 0E 31K 31X
 min -21.12 cm / max 23.71 cm / rms 4.19 cm / ocn 3.88 cm / ctn 4.69 cm

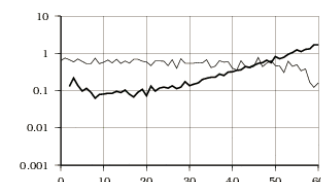
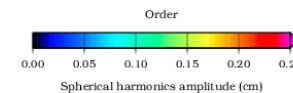
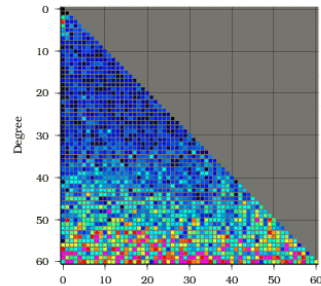


NGGM/MAGIC



Maximum angle
25°

σ SST = 2.5e-9 m/s



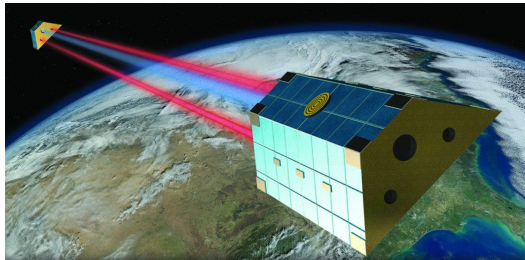
Spectrum by degree (bold) and order (cm)

- MARVEL pre-Phase-A has taken place in a context where next generation gravity missions are under design among many actors



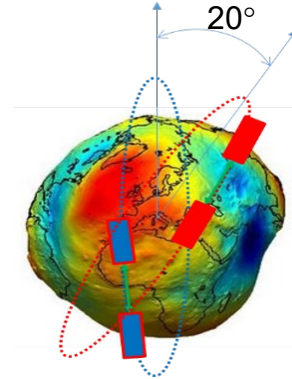
MASS CHANGE

Continuity of GRACE & GRACE FO



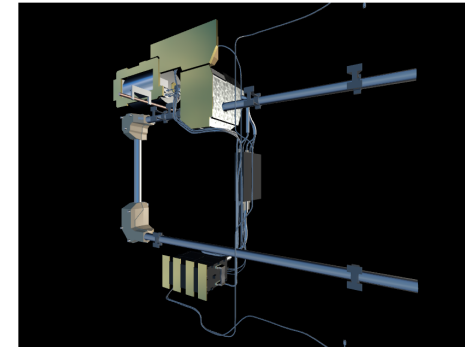
NGGM/MAGIC

Innovation



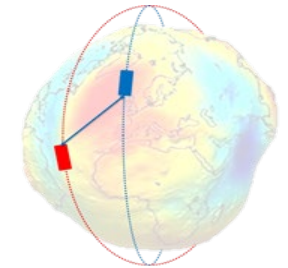
GRACE-I

Continuity

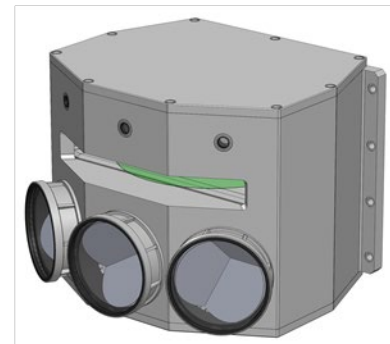
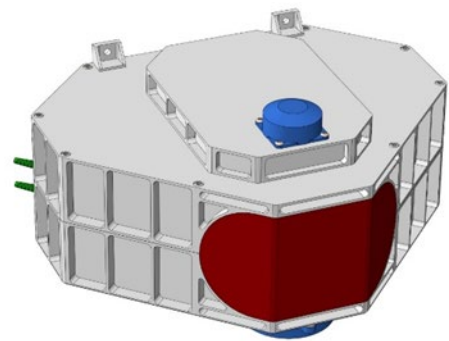


MARVEL

Innovation



THE LASER RANGING INSTRUMENT



❖ The laser ranging instrument

SPECIFICATIONS:

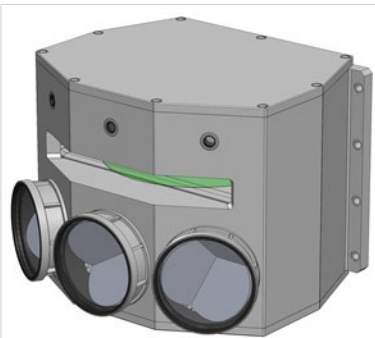
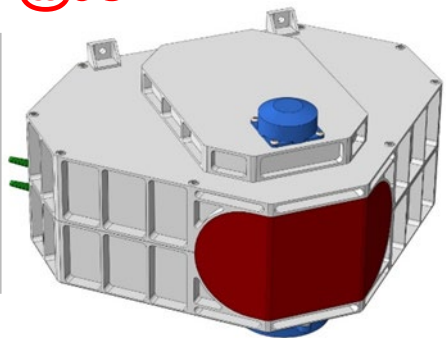
- Average inter-satellite distance ~ **200 km**
- Satellite body in fixed attitude law (→ the laser beam has to be oriented onboard the satellite)
- Instrument-induced dynamical perturbations on the spacecraft $< 10^{-11} \text{ m/s}^2$
- Maximum angle between line-of-sight and satellite reference frames : **+/- 45°**
- Instrument accommodation @ maximum **1.5 m from satellite Center of Mass**
- Ranging accuracy: better than **1 μm @5s** in all circumstances

TECHNOLOGICAL SOLUTION:

- **Chronometric** (i.e. not interferometric) laser link, using proven telecom components
- Line-of-sight angular measurement, for Center-of-Phase/Center-of-Mass correction @ better than **1 μrad @5s**

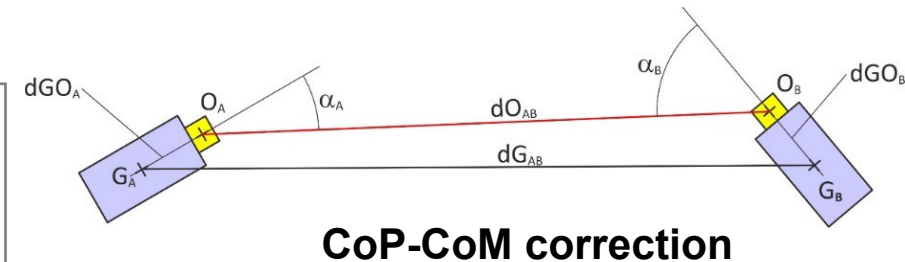
Emitter / Receiver side

with **ultra-light** moving parts

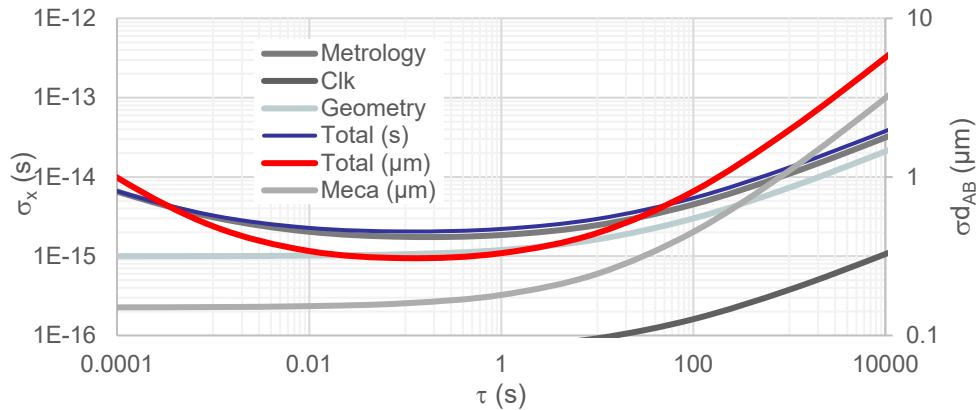


Reflector side

Totally static



Stability curve, expressed in relative stability (left) and ranging noise (right)

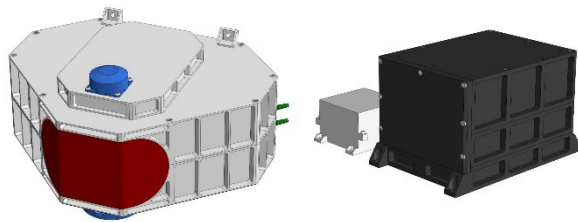


Integrated RMS noise, in distance and velocity, for 1, 5 and 10 s integration time

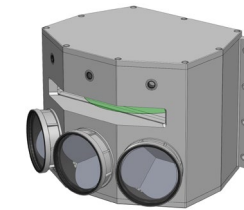
RMS		
	Ranging D	Velocity V
	σ_D RMS	σ_V RMS
	μm	$\mu\text{m/s}$
@ 1 s	1.19	1.24
@ 5 s	1.14	0.26
@ 10 s	1.11	0.13

Mass, consumption and TRL, of the emitter/receiver (left) and reflector (right)

Mass = 13 kg
Power = 55 W



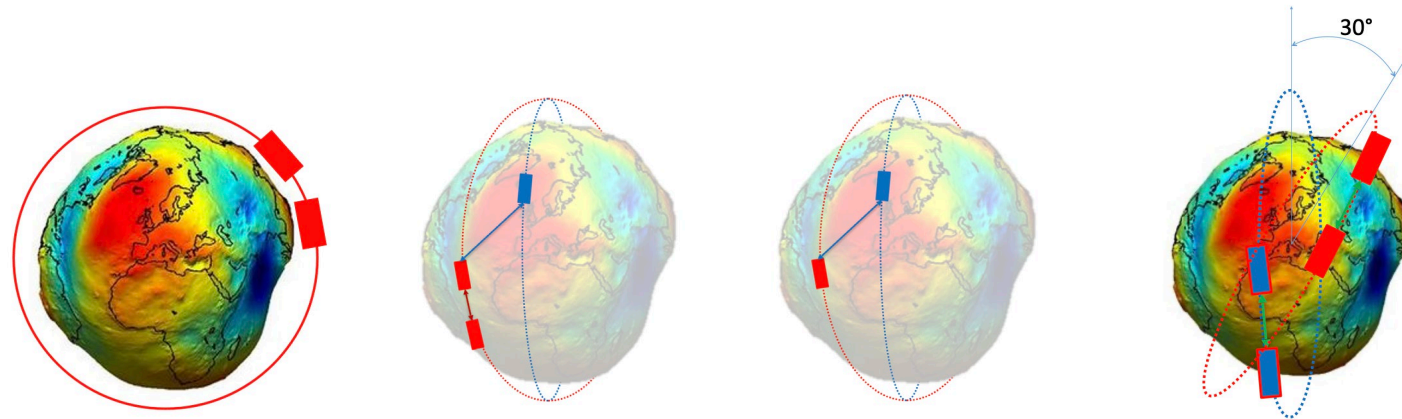
Mass = 8 kg
Power = 14 W



Sub-system	TRL	Heritage
Erbium Laser amplifier	4-5	CW 1,5 μm amplifier
Telecom Transceiver	9	Teleo
Event timer	9	Time Transfer by laser Link T2L2
Femto metrology	4	R&T Program : Metrology bread bord Model
Ultra stable oscillator	9	DORIS Navigation system

Sub-system	TRL	Heritage
Corner cube	9	
Wide field μrad sensor	9	Telescope and algorithm, CNES Patent for the sensor optimization
Ultra stable oscillator	9	DORIS Navigation system

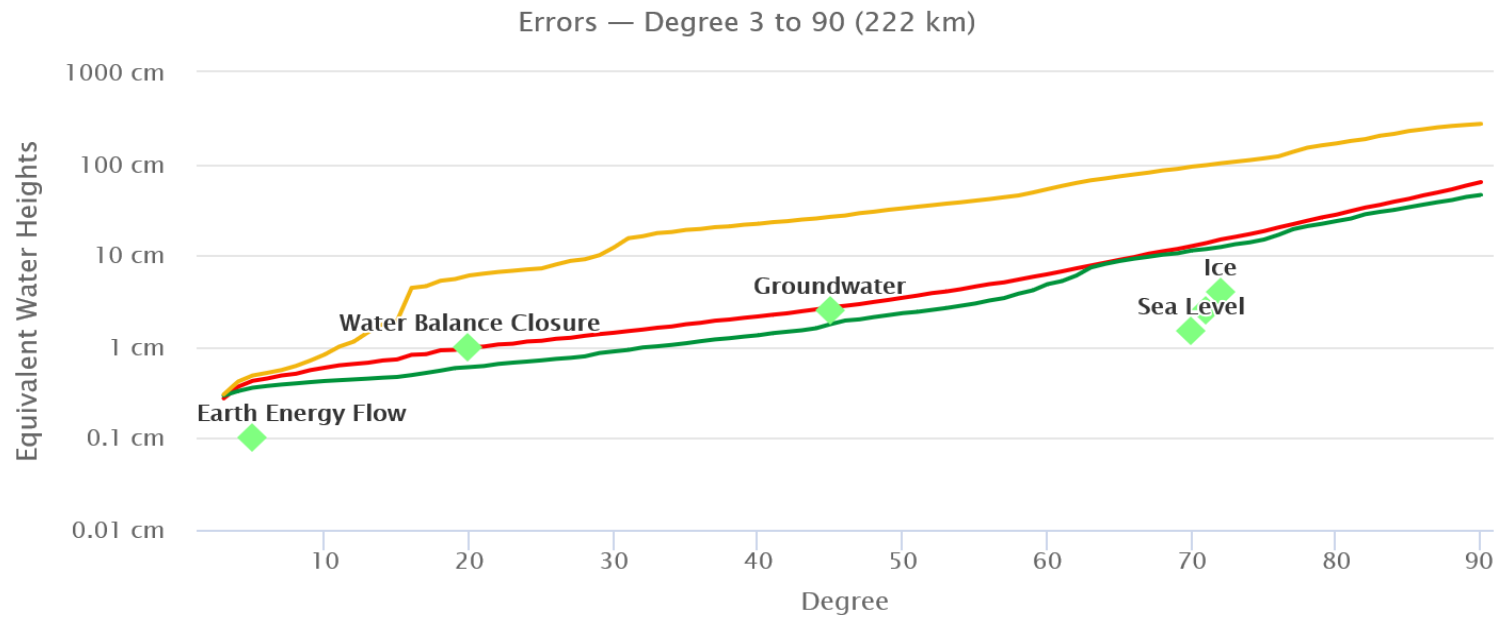
THE NUMERICAL SIMULATIONS



- Most simulations were performed in the “CNES simulation environment” but the latest ones were done in the “ESA Earth System Model” environment for NGGM, in cooperation with GFZ
- The gravity field recovery performances of many different configurations were studied: classic “GRACE-type” single pair, double pair (“Bender”), two-satellite pendulum and three-satellite pendulum (i.e. a “GRACE-type” pair + third pendulum satellite). In each case we explored different altitudes, different inter-satellite separations, etc.
- The main outcome is that Bender and Pendulum achieve comparable performance at the monthly time scale and that they both provide a 6 to 8-fold improvement over current GRACE-type missions
- What was not considered in the simulations, but does have an impact on the results, is the fact that the Bender configuration allows a better time sampling of the gravity field because of its double pair, than the simple “2 satellite” pendulum

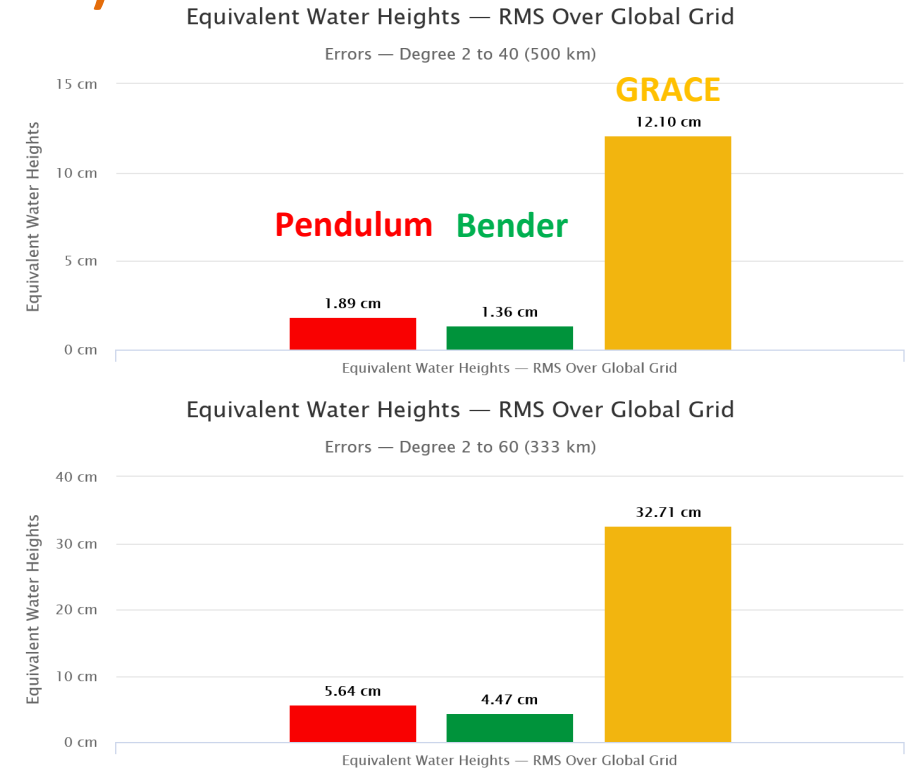
Unconstrained results

for the three mission types (pendulum, bender and grace-like)



- CNES GRACE+PENDULUM [H=(490,490) km, D=(200,200) km, $\alpha=(00,45)^\circ$, KBR=(LRI,Nominal)]
- CNES BENDER [H=(490,490) km, D=(200,200) km, KBR=(LRI,LRI)]
- CNES GRACE [H=490 km, D=200 km, KBR=LRI] ◆ Scientific objectives

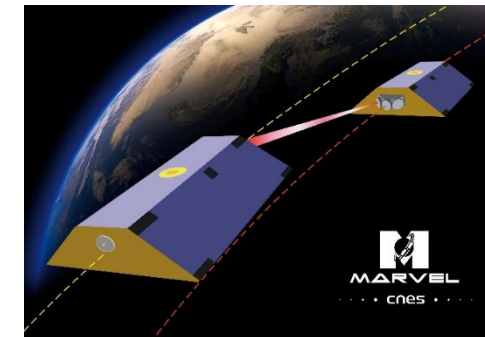
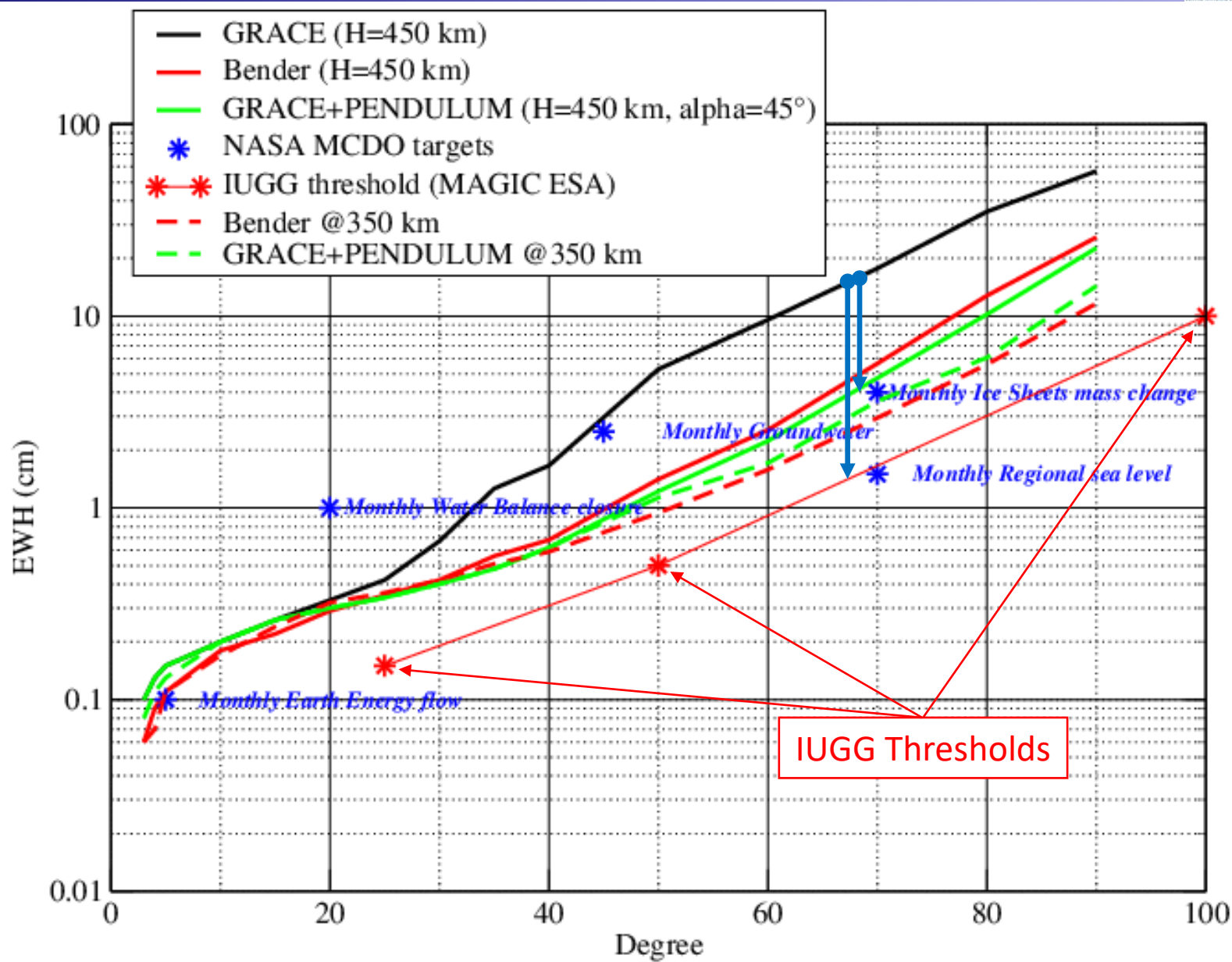
Cumulated error from degree from 3 to 90



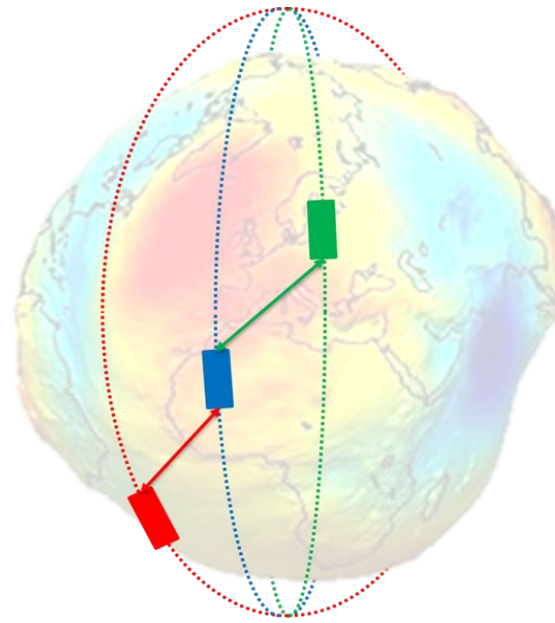
Root mean square (rms) over the global grid

Top panel: degrees 2 to 40
Bottom panel: degrees 2 to 60

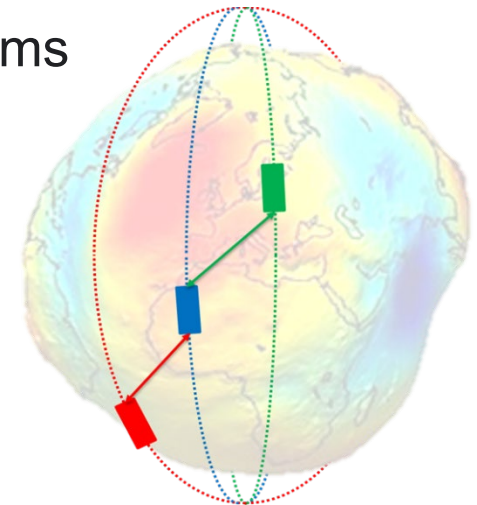
Comparison to NASA and ESA thresholds and targets



CONCLUSIONS



- Plans for future missions, with inclined measurements, show a very clear improvement in scientific results compared to current polar missions (GRACE/GRACE-FO)
- MARVEL pre-phase-A has proven the interest of the pendulum concept
- Chronometric laser ranging instrument → for ~200 km distance:
 - range better than 1 μ meter @5s,
 - angle better than 1 μ rad @5s,
 - lateral scan up to +/- 45°
 - based on several innovative concepts using proven optical telecom subsystems
 - precise, cheap, light and low power instrument
- In the future, it can in particular be envisaged as the ranging instrument of a constellation of 3 or more LEO satellites performing one-to-one pendulum measurements and **covering in a single pass a swath of a few hundred kilometres at the equator with a high temporal repetitivity**



THANK YOU for your attention

