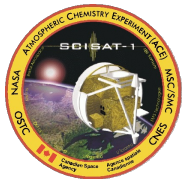


Recent Validation and Science Results from the Atmospheric Chemistry Experiment (ACE)

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ACE on SCISAT

Atmospheric Chemistry Experiment (ACE) Satellite Mission

Mission to measure atmospheric composition: profiles of trace gas species, cloud and aerosol extinction and temperature/pressure

Focusing on investigating:

- Distribution of ozone in upper troposphere and stratosphere
- Effects of biomass burning on the troposphere
- Relationship between atmospheric chemistry and climate change

Size: 1.12 m dia. x 1 m; 152 kg

Total power: 70 W (from single solar panel)

Launch date: 12 August 2003

Orbit: 74° inclined circular orbit at 650 km

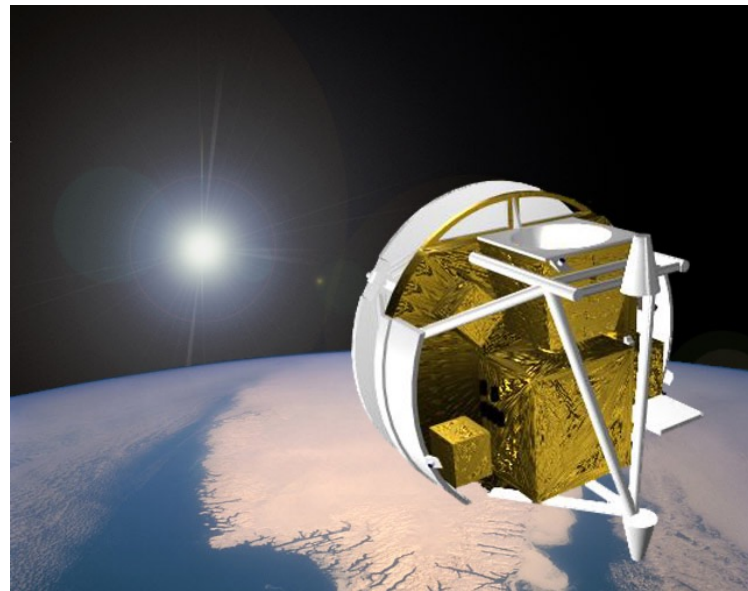
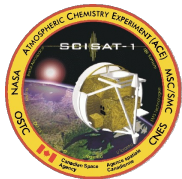


Image: T. Doherty.



ACE Solar Occultation Instruments

ACE-FTS (Fourier Transform Spectrometer):

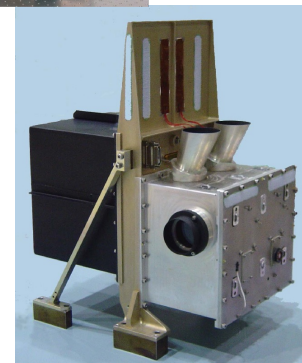
- Mid-infrared FTIR spectrometer – 2-13 microns at 0.02 cm^{-1} spectral resolution and two filtered imagers (Visible and NIR)

ACE-MAESTRO (Measurement of Aerosol Extinction in the Stratosphere and Troposphere Retrieved by Occultation):

- dual UV/Visible/NIR spectrometer – 0.285 to 1.030 microns, with $\sim 1\text{-}2 \text{ nm}$ resolution

Time series: From late February 2004 and on-going

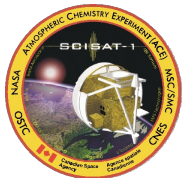
- Up to 30 sunrise and sunset measurements per day
- Designed for 2 year mission lifetime – in 19th year on orbit
 - Starting to see some degradation in ACE-FTS performance
 - ACE-MAESTRO continues to “age gracefully”





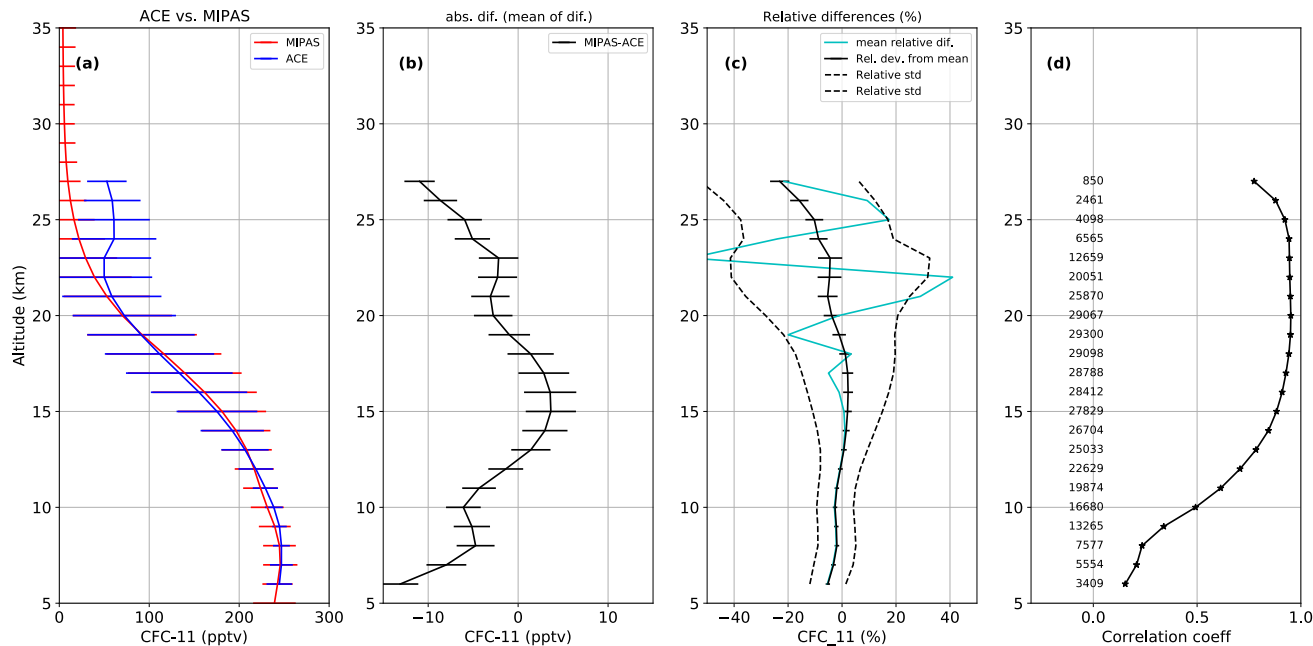
ACE Data Products

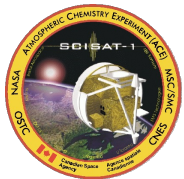
- ACE-FTS profiles (current version 4.1/4.2):
 - Tracers: H_2O , O_3 , N_2O , NO , NO_2 , HNO_3 , N_2O_5 , H_2O_2 , HO_2NO_2 , N_2 , SO_2
 - Halogen-containing gases: HCl , HF , ClONO_2 , CFC-11 , CFC-12 , CFC-113 , COF_2 , COCl_2 , COFCl , ClO , CF_4 , SF_6 , CH_3Cl , CCl_4 , HCFC-22 , HCFC-141b , HCFC-142b , HFC134a , HFC-23
 - Carbon-containing gases: CO , CH_4 , CH_3OH , H_2CO , HCOOH , C_2H_2 , C_2H_4 , C_2H_6 , OCS , HCN , acetone, CH_3CN , peroxyacetyl nitrate (PAN), CO_2 (5-18 km and >60 km), pressure / temperature from CO_2 lines
 - Isotopologues: Minor species of H_2O , CO_2 , O_3 , N_2O , CO , CH_4 , OCS , NO_2 , HNO_3
- MAESTRO profiles (current version 3.13):
 - O_3 , NO_2 , optical depth, aerosol and water vapor (v31)
- IMAGERS profiles (current version 4.1/4.2):
 - Atmospheric extinction & aerosol extinction at 0.5 and 1.02 microns



MIPAS and ACE-FTS CFC Comparisons

- ACE-FTS v4.1 and new MIPAS IMK-IAA v8 retrievals between 2005 and 2012, global comparisons within 24 hours and 1000 km for CFC-11 and CFC-12
- Max. ACE-FTS retrieval altitude varies with lat. (~24-27 km from polar to tropics)
- Typical differences of less than 10 pptv and within 10%
- Improved agreement – previous comparison differences showed MIPAS ~10-20 pptv higher than ACE-FTS

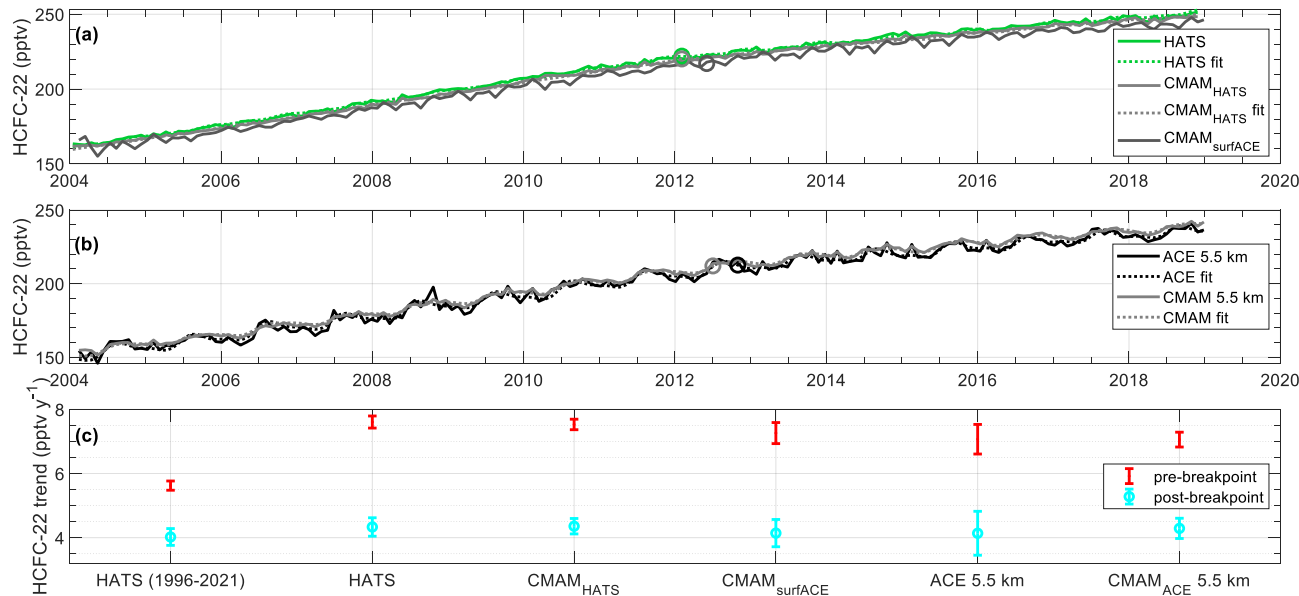


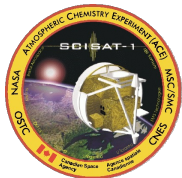


Validation of ACE-FTS HCFC-22: Trend

In addition to comparisons with MIPAS, MkIV and BONBON balloon, and CARIBIC results

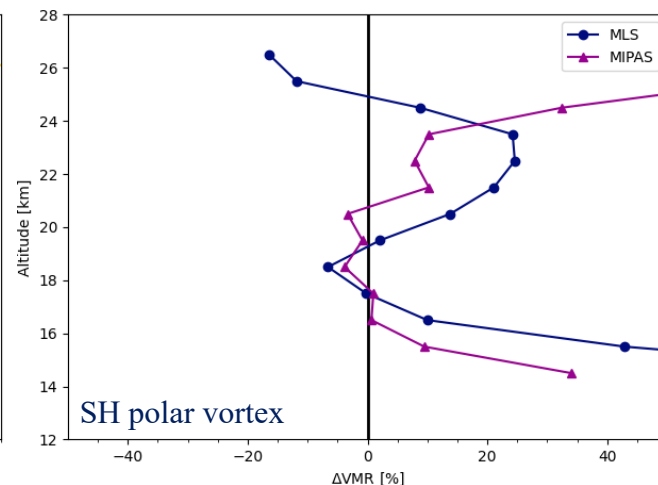
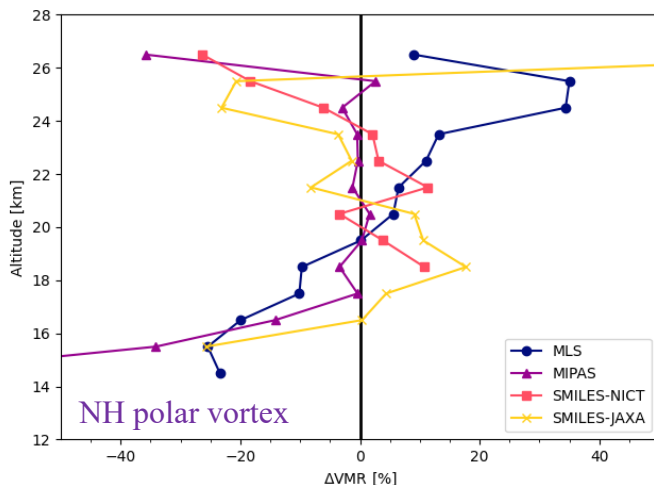
- Comparing trends in HCFC-22 from ACE-FTS at 5.5 km, NOAA HATS surface flask measurements and subsampled CMAM39-SD at ACE-FTS and HATS locations
- Fits of global monthly means using multiple linear regression
- Consistent change in trend at 2012 across datasets
- Significant decrease in trend at breakpoint
Pre: 7.2 ± 0.6 pptv/year;
Post: 4.2 ± 0.7 pptv/year





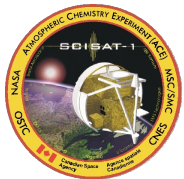
v4.2 ClO versus MIPAS, Aura-MLS and SMILES

- Comparison of daytime profiles taken within 300 km of ACE-FTS, within polar vortex
- Used PRATMO chemical box model to scale comparison profiles to ACE local solar time
- Time periods for comparison vary; several months for SMILES to 2004-2020 for MLS
- Differences found from 17 - 24 km are within approx. $\pm 10 - 20\%$ for NH and between -10 and +25% for SH



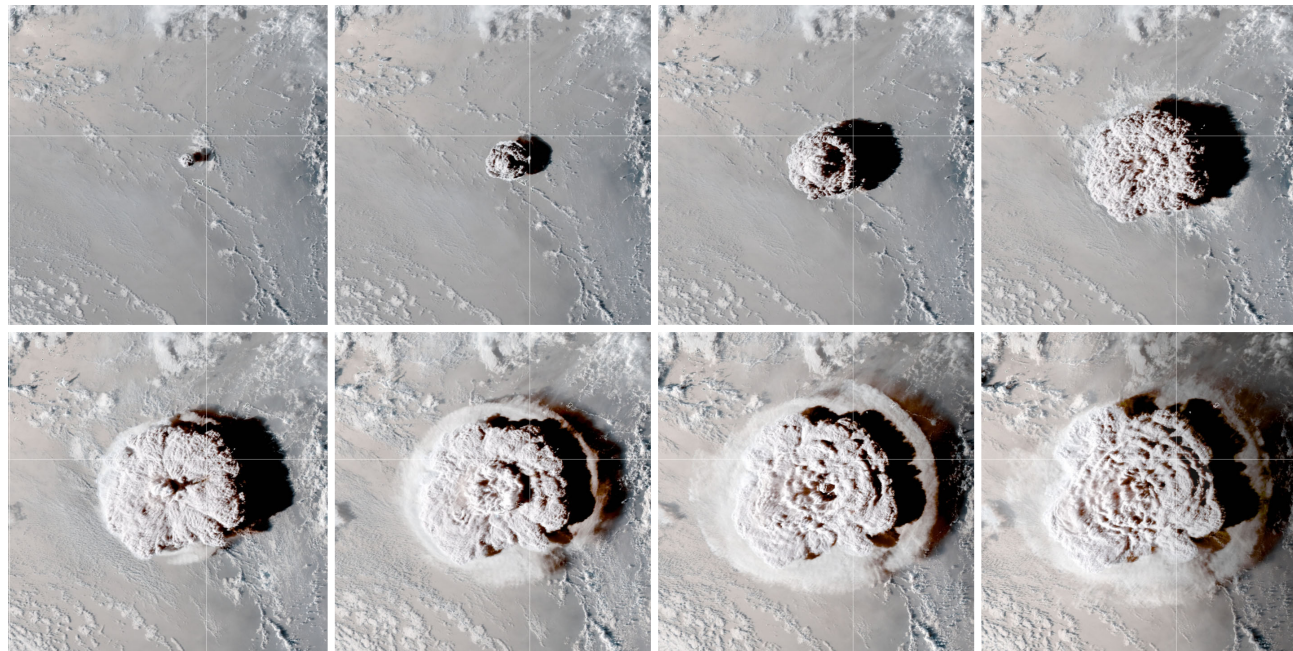
Medians of differences (ACE-FTS – comp.) shown for comparisons.

L. Saunders, N. Ryan et al., in preparation.



Hunga Tonga–Hunga Ha‘apai Eruption

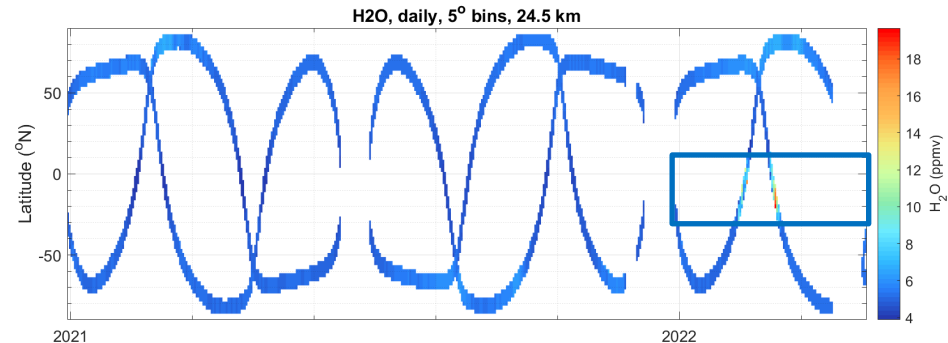
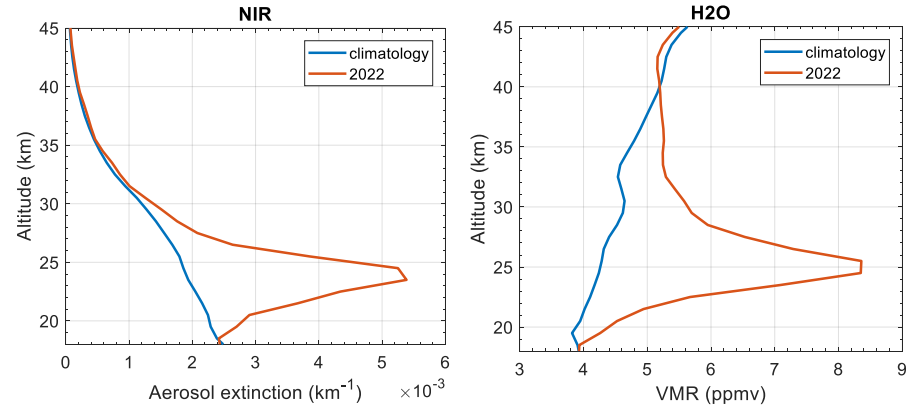
- 15 January 2022:
Explosive eruption from submarine volcano (20.54° S, 175.38° W)
- Impact seen in tsunami alerts and ionospheric disturbances, and plume reached mesosphere
- Stereoscopic images from GOES-17 and Himawari-8 (10 min. interval for images)

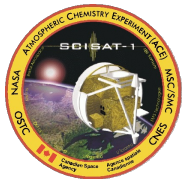




Hunga Tonga–Hunga Ha‘apai Eruption

- ACE-FTS results from 30° S–10° N, comparing February 2022 to zonal mean from all of 2004–2021
- Aerosol extinction from ACE Imagers and SO₂ and H₂O from ACE-FTS all show strong enhancement at ~22–25 km
- Estimate that injection of water into stratosphere is equal to ~10% of total





Timeseries analysis of mesospheric water vapour

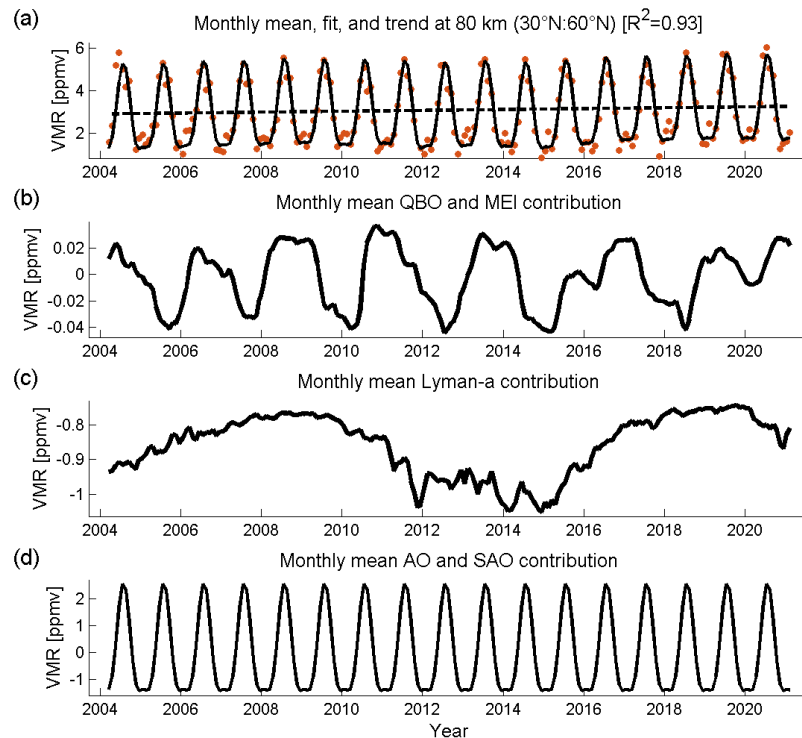
- Using 2004-2020 time series from ACE-FTS to examine behaviour of middle atmosphere water vapour over Solar Cycle 24 (2008-2019)
 - Monthly zonal means for 30 degree regions and all latitudes at each 1-km from 40 to 90 km

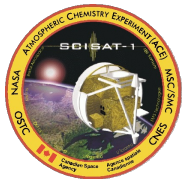
- Fitting using multi-variable linear regression:

$$\begin{aligned}
 H_2O(t) = & a_0 + a_t t + a_{QBO} QBO(t) \\
 & + a_{MEI} MEI(t) + a_{Ly\alpha} Ly\alpha(t) \\
 & + \sum_1^2 \left(a_{1,n} \sin\left(\frac{2\pi t}{1/n}\right) + a_{2,n} \cos\left(\frac{2\pi t}{1/n}\right) \right)
 \end{aligned}$$

QBO(t): radiosonde winds at 30 hPa over Singapore - Free University of Berlin
MEI(t): multivariate ENSO index - NOAA Physical Sciences Laboratory
Ly α (t): Lyman- α radiation using LISIRD composite of instrument and model time series from LASP U. Colorado

Results for 30°-60° N at 80 km





Summary

- ACE has evolved from an ozone-focused mission to a climate-focused mission
 - Advantage of large number of species being measured simultaneously
 - Long time series with stability needed to meet climate data record requirements
- Data availability for ACE:
 - ACE-FTS and MAESTRO from <https://database.scisat.ca/level2> (registration required)
 - Data quality flags are being produced separately [doi:10.5683/SP2/BC4ATC](https://doi.org/10.5683/SP2/BC4ATC)
 - Zonal mean climatologies available through SPARC Data Initiative or talk to me!

Funding for ACE and this work provided by:

- Canadian Space Agency (CSA) and Natural Sciences and Engineering Research Council of Canada

Thanks to:

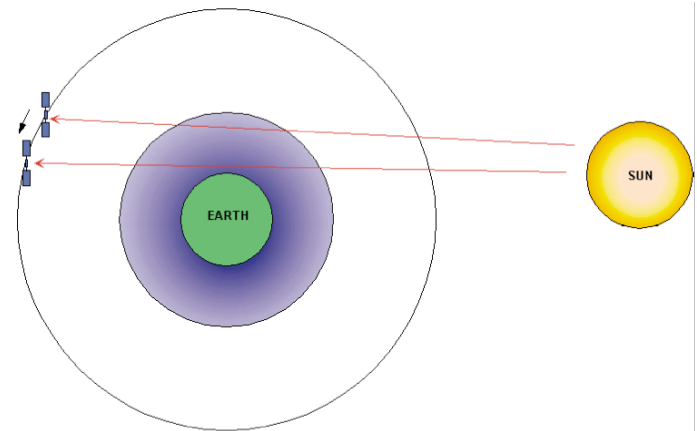
- SCISAT Science Operations Centre, Peter Bernath and Chris Boone
- MIPAS, MLS, SMILES teams: Gabriele Stiller, Thomas von Clarmann, Alexandra Laeng, Michelle Santee, Gloria Manney, Luis Milan, Makoto Suzuki, Masato Shiotani, Yasko Kasai
- CARABIC, HATS, MkIV teams: Andreas Engel, Tanja Schuck, Stephen A. Montzka, David E. Oram, Geoff Toon

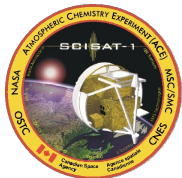


Technique: Solar Occultation

Measure absorption spectra during sunrise and sunset

- Advantages:
 - High signal-to-noise ratio and long path to allow measurements of species with low concentrations as compared to limb emission or nadir
 - Simultaneous measurements of many species when using infrared “fingerprint” region
 - Less susceptible to changes in instrument sensitivity because of reference exo-atmospheric measurement
 - Produces very accurate and reliable time series
- Limitation of technique
 - Geographic coverage limited to 30 occultations per day
 - Samples only free troposphere





Validation of ACE-FTS HCFC-22

- Comparisons with ACE-FTS v4.2 with MIPAS IMK/IAA v8
- ACE-FTS biased low up to ~5-10% from 10-20 km
- Similar results with MkIV balloon FTIR and CARBIC aircraft in situ measurements

