

A pulsed Tuneable LAser system for the characterisation of Spectrometers (ATLAS).

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Project Deliverables

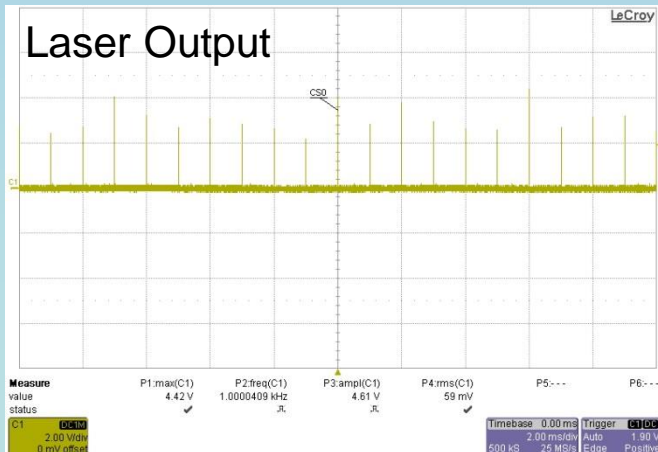
Project Duration : March 2015 – June 2016

- **WP1:** Commissioning of tuneable laser system.
- **WP2:** Radiometric characterisation of 2 array spectroradiometers using tuneable laser sources
- **WP3:** Development of correction methodologies for stray light and linearity

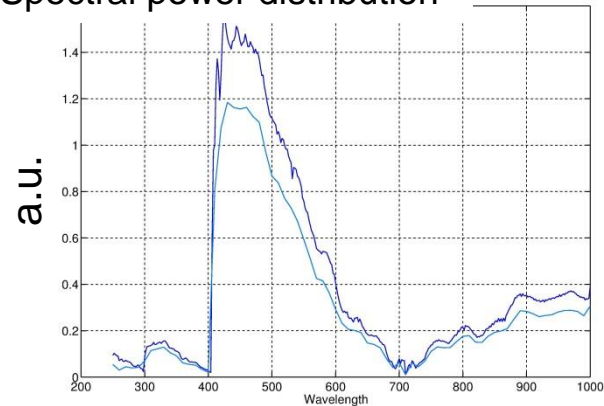
WP1 Commissioning of tuneable laser

Delivered August 2015

- Tuning range 210 -2600 nm
- Pulse 5 ns
- repetition rate 1000 Hz
- Pulse power at 450 nm ~300 mW



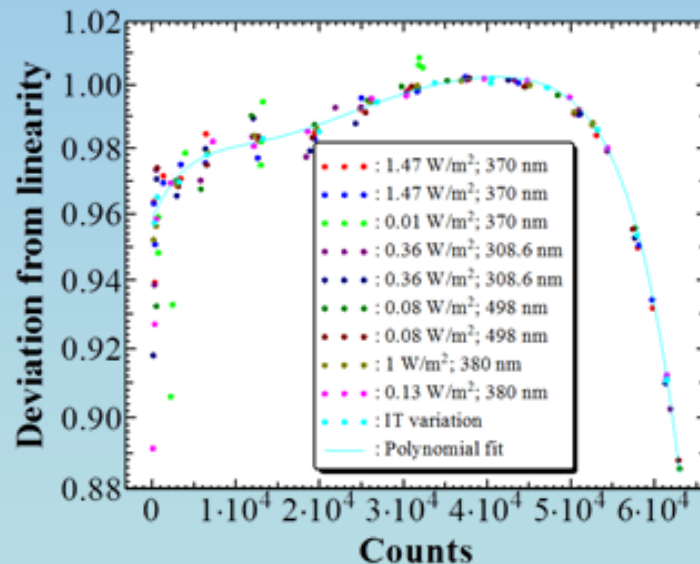
Spectral power distribution



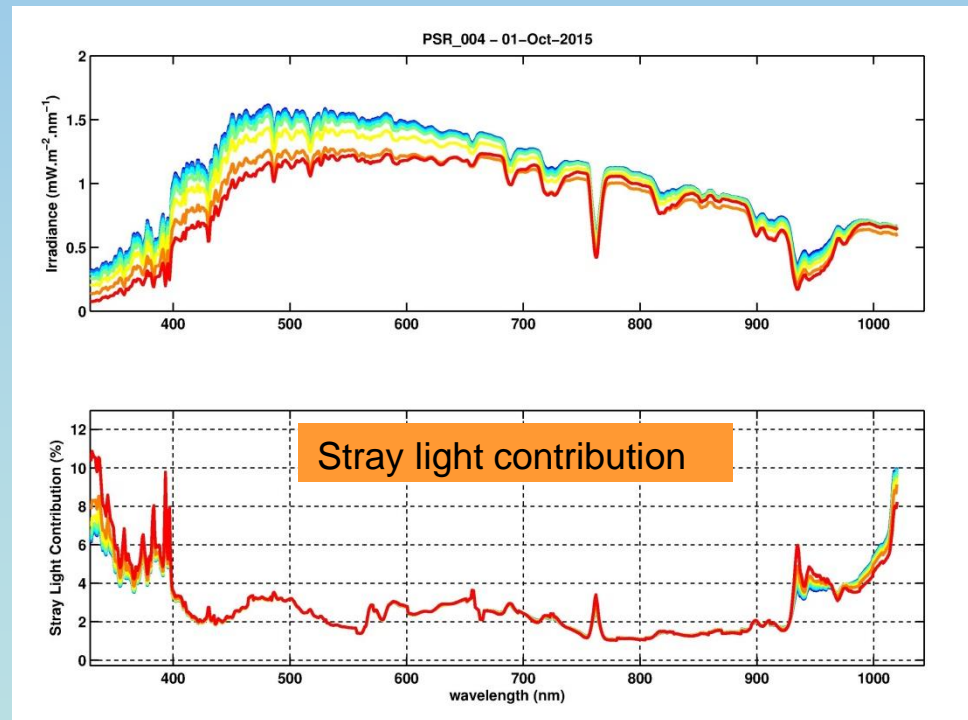
Characterisation of array spectroradiometers

Main radiometric parameters to determine are:

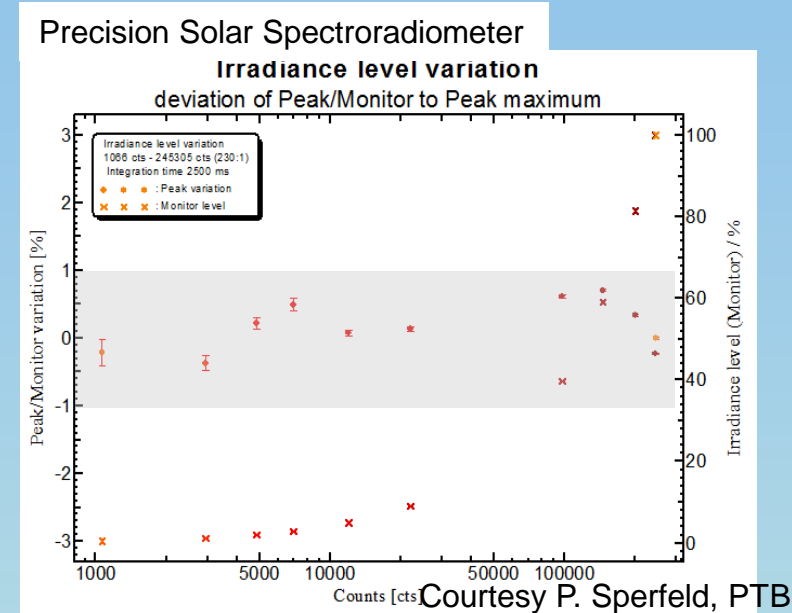
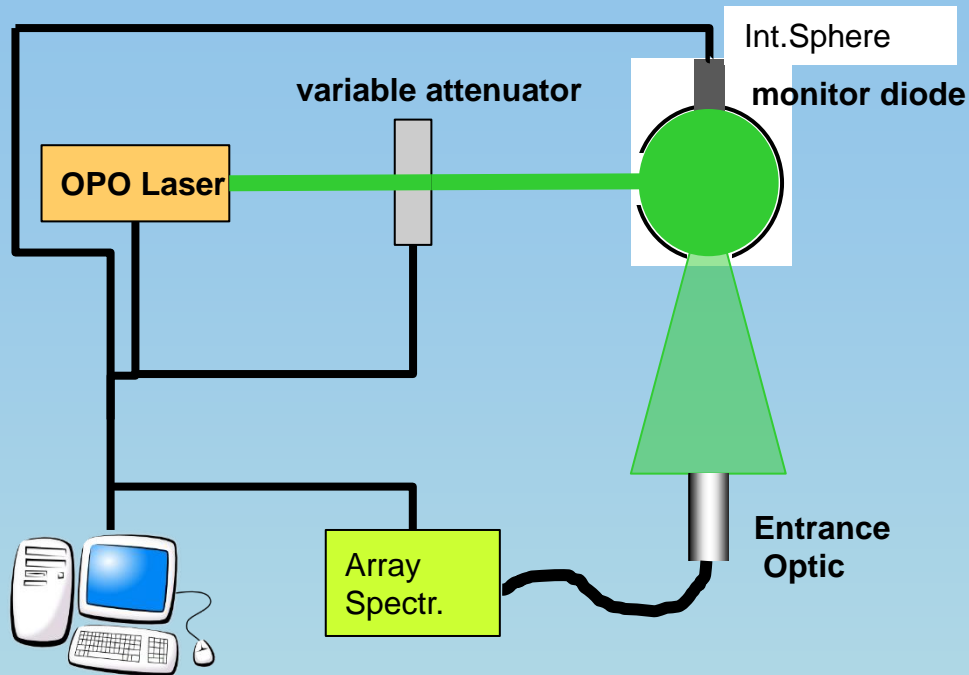
- linearity of system (detector and data acquisition)
- Stray Light, in-range and out-of-range (if applicable)



Significant nonlinearity of up to 12%

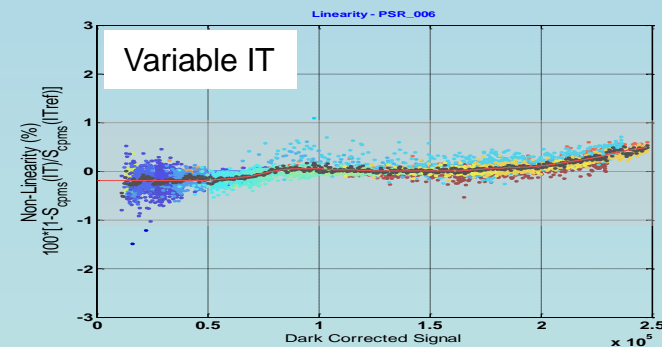


Tuneable laser setup at PMOD/WRC



Linearity Characterisation

- 1) Versus Intensity change
- 2) Versus integration time



Stray-Light correction of array spectroradiometers

Problems

- Spectral stray light is a major challenge for array spectroradiometers measuring solar UV radiation (300-400 nm), and for deep absorption lines (e.g. O₂ at 762 nm, water vapour).
- Out-of-range (Oor) stray light generates background radiation across the whole sensitivity range of the detector (up to 1100 nm for Si-Detector)

Solutions

- Matrix method (Zong et al., AO 2006), based on measured Line spread function
 - corrects SR up to 2 orders of magnitude
- Oor stray light can be corrected but is very complex and needs knowledge of spectrum which is not measured by the instrument.
- Hardware solution: Use of cut-off filters to reduce Oor radiation

Stray-Light correction of array spectroradiometers

Methodology for in-range and out-range stray light correction from Zong et al., 2006, and Nevas et al., 2014.

$$\mathbf{Y}_{s_spec} = \mathbf{D} \cdot \mathbf{Y}_{IB} + \Delta,$$

- \mathbf{D} stray light distribution function
- \mathbf{Y}_{IB} the in-band signal
- Δ Oor contribution

$$\Delta = \mathbf{s}_{OoR} \cdot \mathbf{E}_{OoR} \cdot \delta\lambda,$$

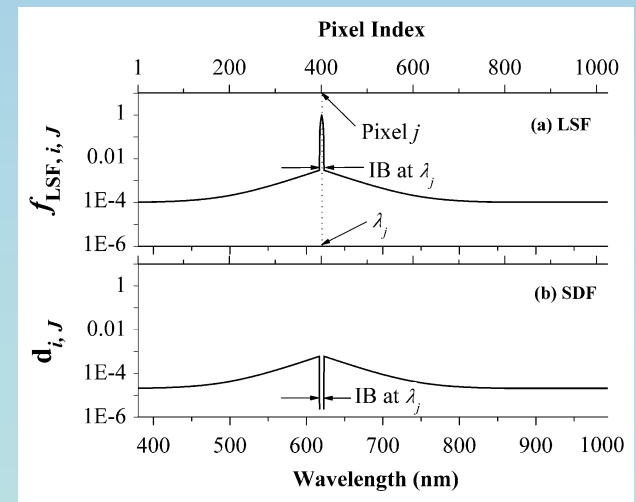
Nevas et al., 2014

\mathbf{E}_{OoR} is the spectral radiation outside the spectral range of the spectroradiometer

Use of matrix inversion to retrieve the in-band signal

$$\mathbf{Y}_{IB} = [\mathbf{I} + \mathbf{D}]^{-1} \cdot [\mathbf{Y}_{meas} - \Delta] = \mathbf{A}^{-1} \cdot [\mathbf{Y}_{meas} - \Delta].$$

Zong et al., 2006

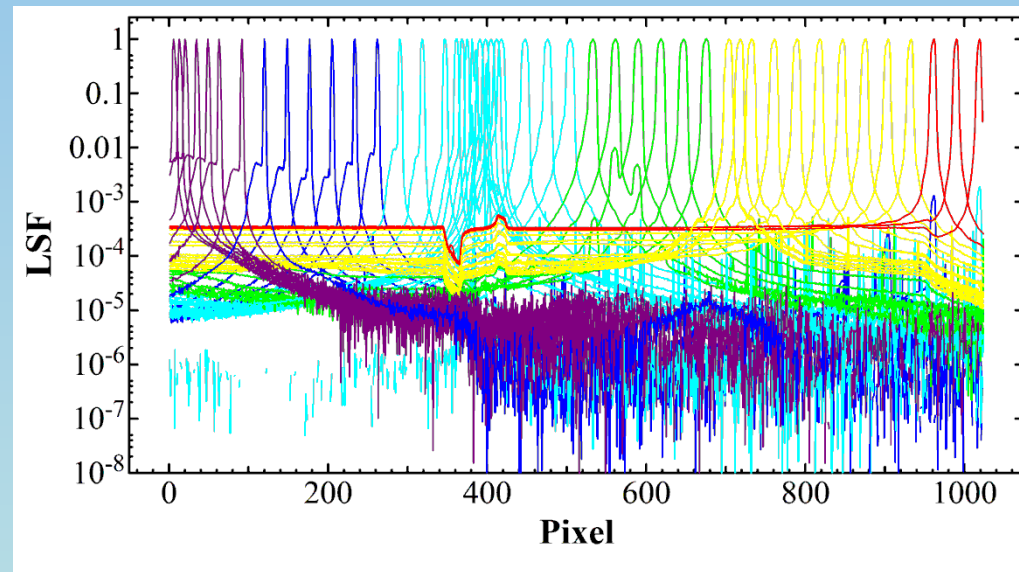


In-range stray light

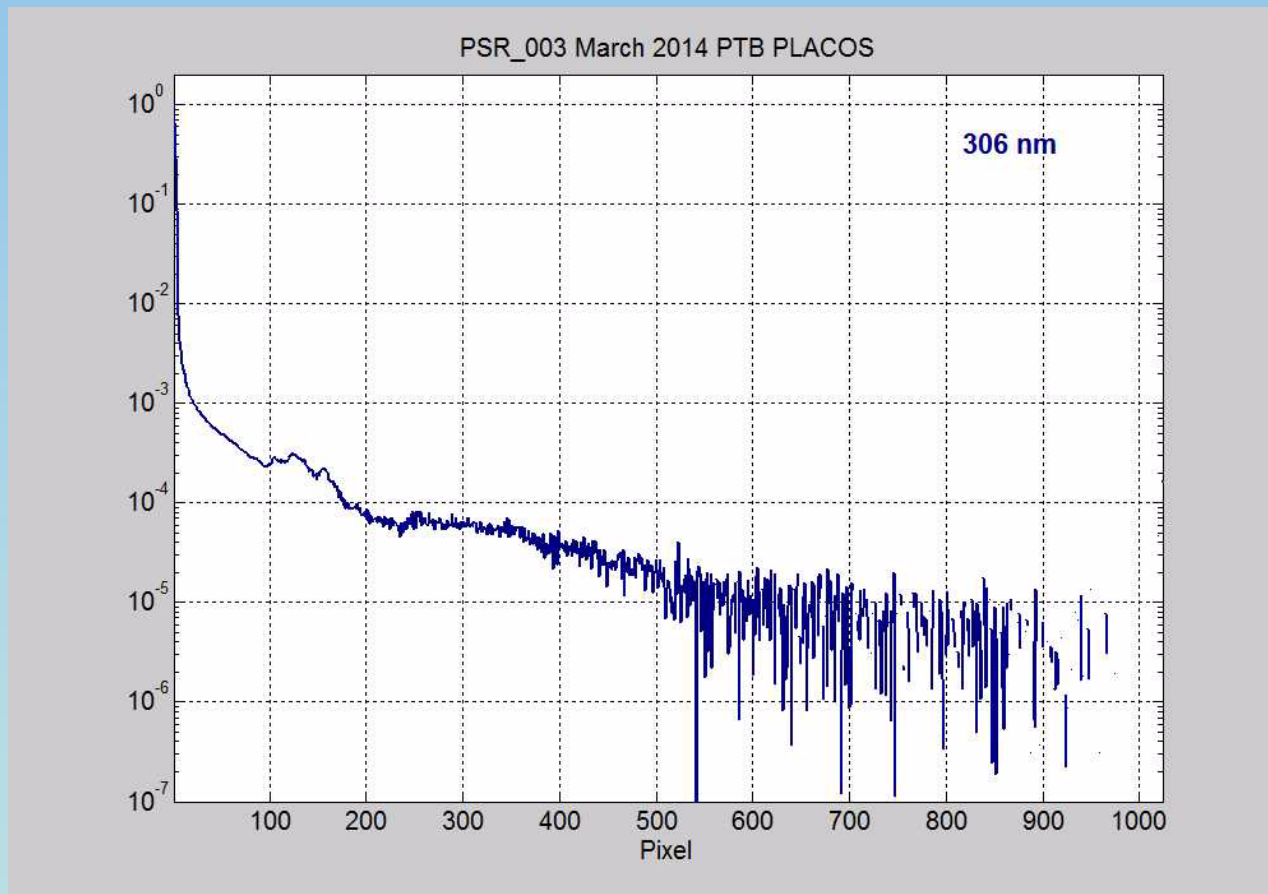
The line spread functions (LSF) are determined using tuneable laser facilities in the operational range of the spectroradiometer.

The stray-light corrected signal Y_{IB} can be retrieved from the measurement by applying a stray-light correction matrix A^{-1} to the measured signal, following the method of **Zong, 2006**.

$$Y_{IB} = A^{-1} Y_{meas}$$

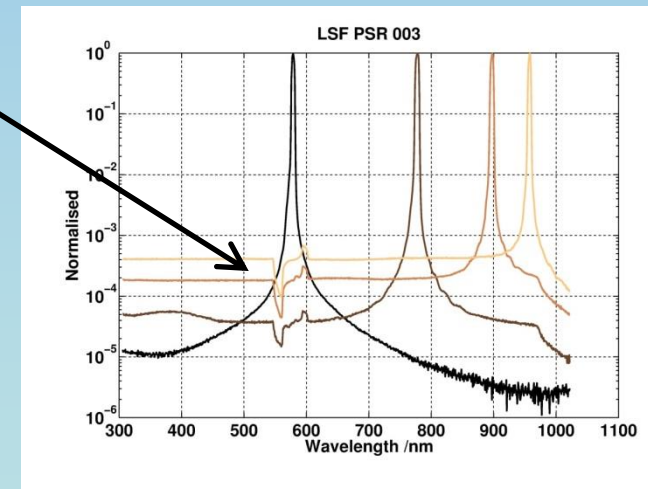
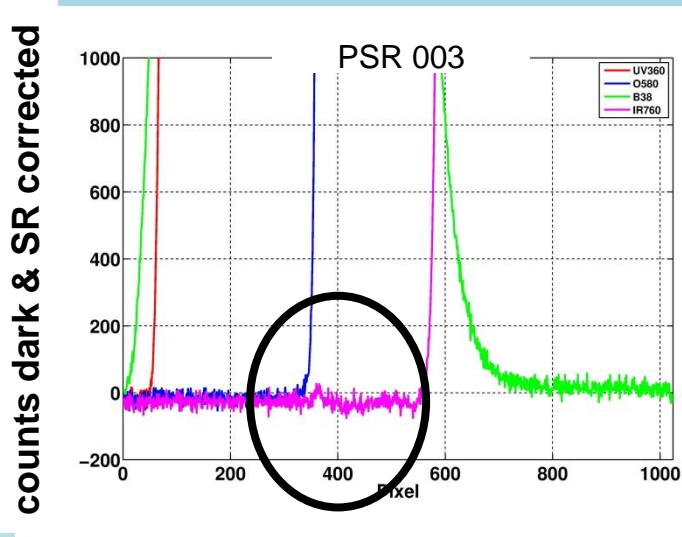
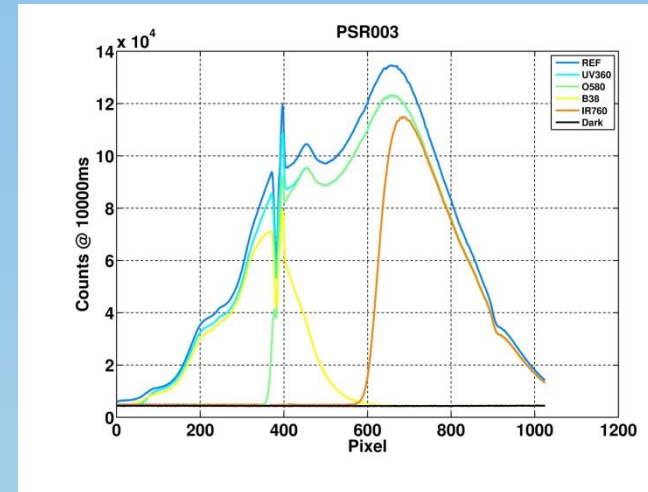
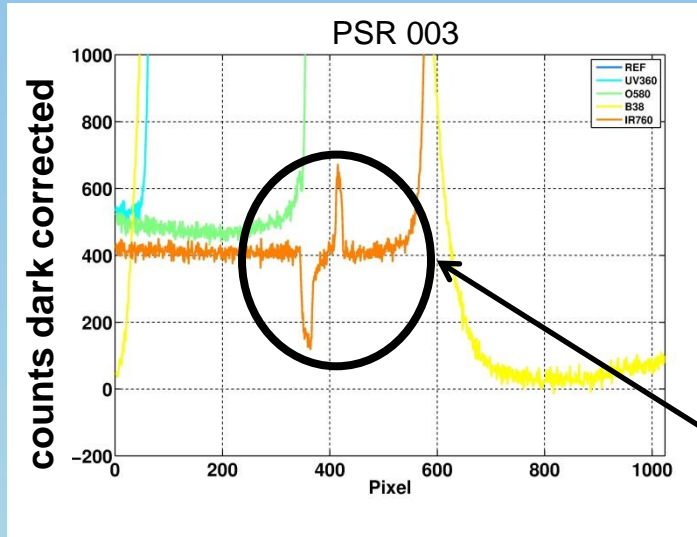


Example line spread functions of PSR 003 measured at the tuneable laser facility PLACOS of PTB



Stray-light correction validation

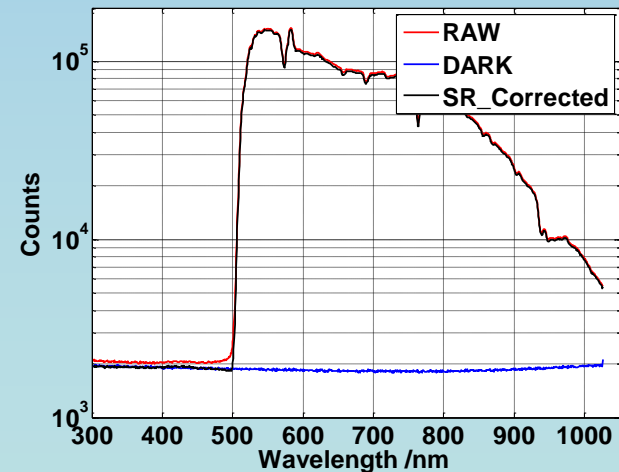
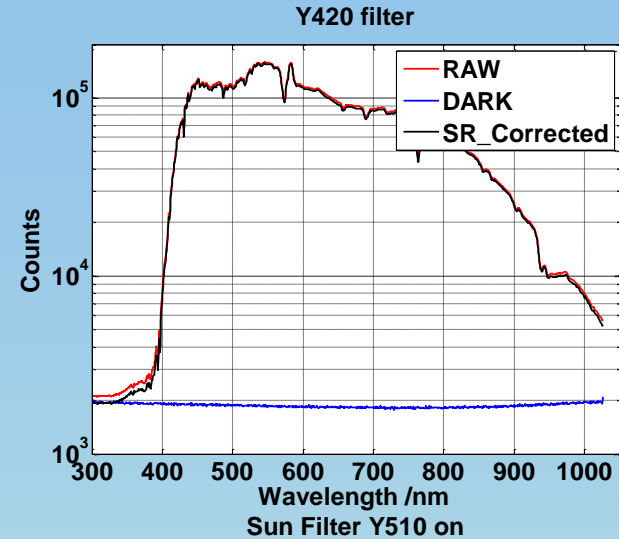
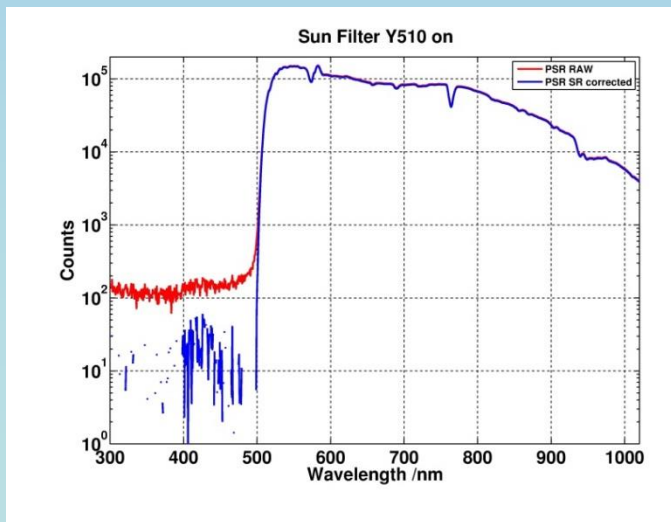
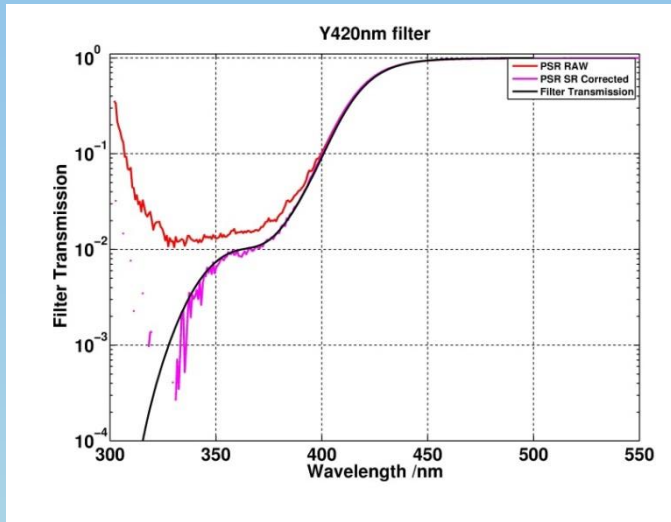
1) Using band-pass filters and halogen lamp



The Stray-light correction using the appropriate SR-Matrix is effective in removing more than 95% of the SR-Signal

Validation of stray-light correction

Using solar measurements and cut-off filters



The Stray-Light can be reduced by ~ factor 10

Project schedule Kick-Off (KO) 22 June 2015

Deliverable number	Deliverable Description	Due Date (KO+#month)	Work Packages	Progress
TD0	Brief Monthly Reports	Monthly	all	✓
TD1.1	Tuneable laser system operational	12.2015	WP1	completed
TD1.2	SOP for LSF and Oor stray light	2.2016	WP1	in progress
TD1.3	SOP nonlinearity	3.2016	WP1	in progress
TD2.1	LSF, Oor stray light and linearity determined for System 1.	6.2016	WP2	scheduled 1.2016 (Pandora)
TD2.2	LSF, Oor stray light and linearity determined for System 2.	9.2016	WP2	scheduled 3.2016 (Phaeton)
TD3.1	Report on nonlinearity correction method	6.2016	WP3	5.2016
TD3.2	Report on In-range and out-of-range stray light method	10.2016	WP3	5.2016
TD4	Final report, Presentation	10.2016	all	6.2016