



CNES Experience in Image Quality and Illustration with Pleiades Very High Resolution Data

Patrice Henry

Centre National d'Etudes Spatiales (CNES), 18 avenue Edouard Belin, 31401 Toulouse Cedex 9, France

35 Years of Experience in 'Image Quality' of Optical Satellite Imaging Systems: Ground Characterization, Calibration, Validation, Performances Assessment, Processing Algorithms... Through:

- **SPOT 1, 2 & 3** (10m Panchro and 20m Multispectral - VNIR)
- **SPOT 4** (10m Panchro and 20m Multispectral – VNIR + SWIR)
- **SPOT 5** (2.5m Panchro and 10m Multispectral – VNIR + SWIR – F/B Stereo)
- **VGT** (1km Daily Global Coverage - VNIR + SWIR)
- **POLDER/PARASOL** (Polarization and Multiangle Viewing – VNIR)
- **Pléiades** (0.7m Panchro and 2.8m Multispectral - VNIR)
- **Sentinel 2** (10m Multispectral – VNIR + SWIR)
- **Venµs** (5m Multispectral – VNIR – 2 days revisit)

Experience extended to other Earth Observation Systems: TIR Sounding (IASI), Thermal IR Imaging (Calipso/IIR), Passive Microwave (Megha-Tropiques)

CNES concept of 'Image Quality':

- Image performance considered at system level (board+ground)
 - Satellite / instrument / ground processing...
 - Image quality budget performed at system level
 - End to end simulations
- Continuity in the responsibility of image performance for all the project phases
 - Design / development / commissioning / exploitation
 - Image quality manager member of the project team
- Operational image quality monitoring during exploitation phase
 - Dedicated team, dedicated tools, dedicated environment
 - Strong link between expertise/monitoring

In CNES we prefer 'Image Quality Commissioning' than 'Calibration/Validation'.
The same but even more...

Development of generic tools, data bases, sites catalogs

Both for radiometry and geometry

- **SADE**: measurements data base over calibration sites (desert, snow, ocean ...)
- **MUSCLE**: calibration tools environment
- **ROSAS**: calibration over instrumented site (La Crau, Gobabed...)
- High accuracy GCPs data base
- **AGILE**: Geometrical performances assessment workshop
- **SIGMA/Euclidium**: viewing modelization and geometrical corrections
- Images correlation
- Optimized resampling
- MTF assessment
- Denoising...

But also...

Dedicated tools and methods developed to deal with satellite/instrument/mission specificities

- **Very high resolution image location assessment (Pléiades)**
- **Microvibrations assessment (SPOT5, Pléiades)**
- **Stereo performances assessment (SPOT5/HRS, Pléiades, Venµs)**
- **Polarization parameters determination (POLDER/PARASOL)**
- **Stray light correction assessment (POLDER, Venµs)**
- **Moon calibration thanks to satellite agility (Pléiades, Venµs)**
- **Radiometric calibration over clouds (VGT, POLDER, MERIS)**
- **...**

Illustration with Pléiades

How to use satellite agility to assess Image Quality?

Spatial resolution:

- **Panchromatic: 0.70 m**
(product resampled @ 0.5m)
- **XS (B, G, R, NIR): 2.80 m**

Swath: 20 km

- **Revisit Capability: 1 day with 2 satellites**



- **PHR1A launch: December 17, 2011**
- **PHR1B launch: December 2, 2012**

Mode	Band	Spectral characteristics
Multi spectral	1	430 – 550 nm
	2	490 – 610 nm
	3	600 – 720 nm
	4	750 – 950 nm
Panchromatic	P	480 – 830 nm

New Concept

A small (<1000 kg) and very agile satellite to improve operational capability and minimize the conflicts between users

Designed for high agility

Compact and rigid satellite with low inertia and fixed solar array

Attitude control system with 4 powerful CMG (Control Moment Gyros) actuators

☞ agility (roll and pitch):

5° in 8 seconds

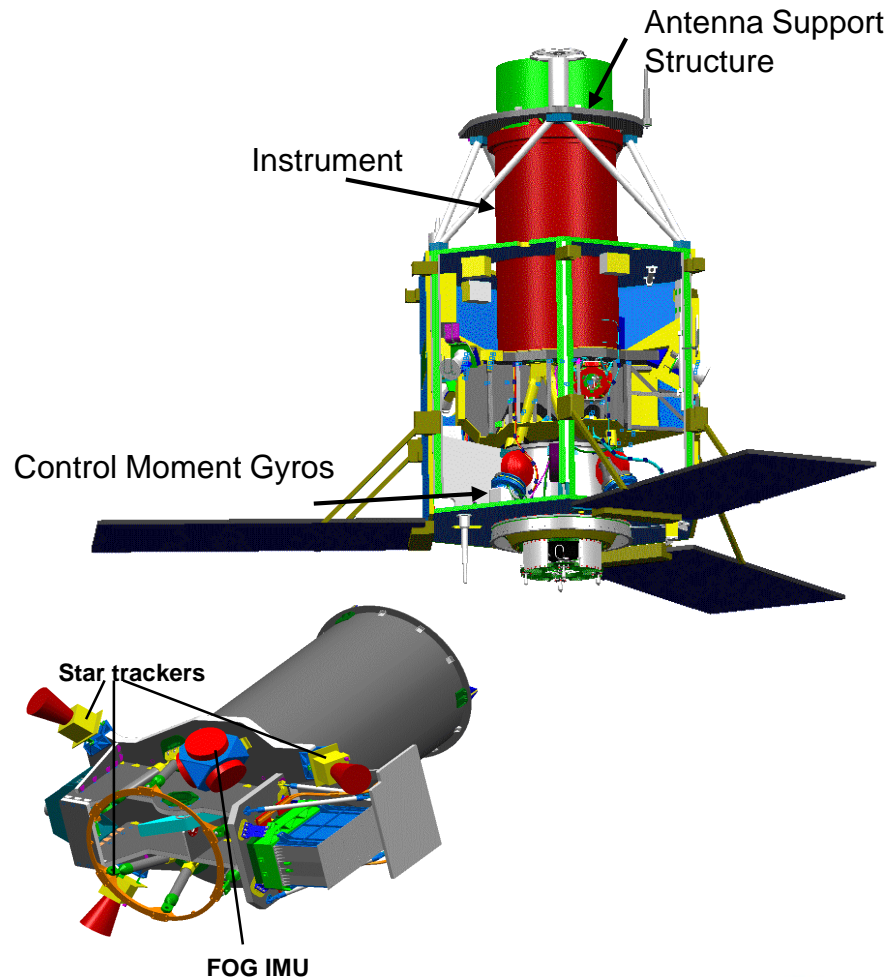
10° in 10 seconds

60° in 25 seconds

Designed for high geometric image quality

Very stable instrument with very accurate sensor heads mounted on the optical bench (3 star trackers and 4 fiber optical gyrometers)

Autonomous navigation with DORIS/DIODE system



Korsch camera:

focal length = 13 m

input diameter = 0.65 m

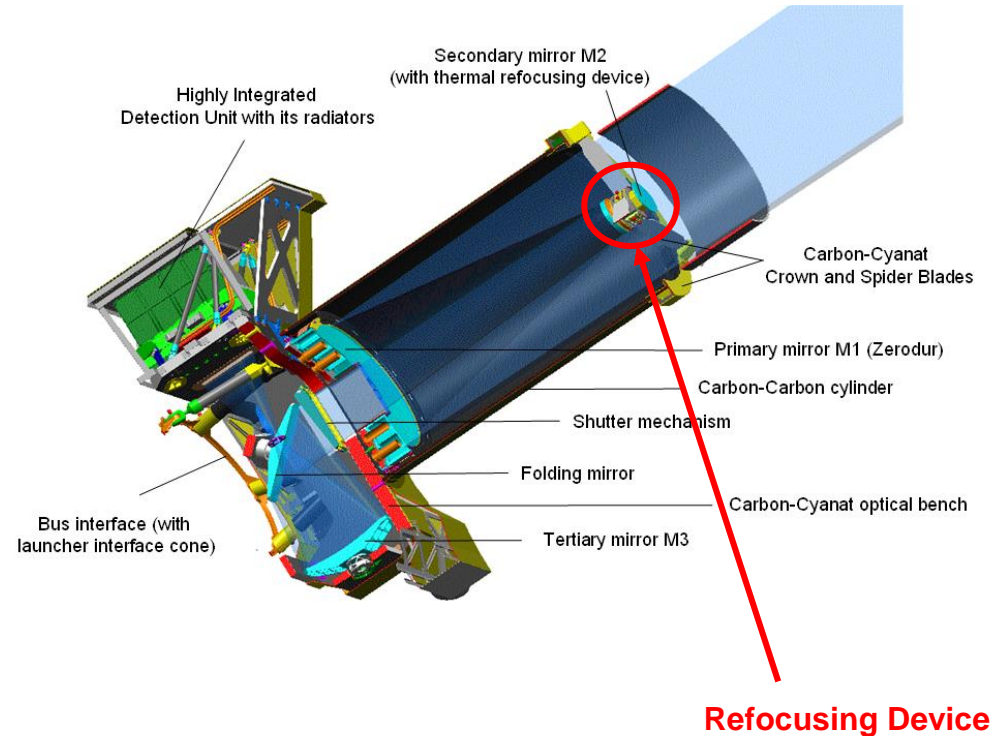
PA retina : 5 TDI detectors

XS retina : 5 four-color CCD

Radiometric resolution: 12 bits

On-board detectors normalization

Wavelet compression: from 1.4 to 3.33 bits/pixel



Mecca "tower clock" viewed by PLEIADES 1B every 90 sec. in a single pass to see 3 faces of the building...

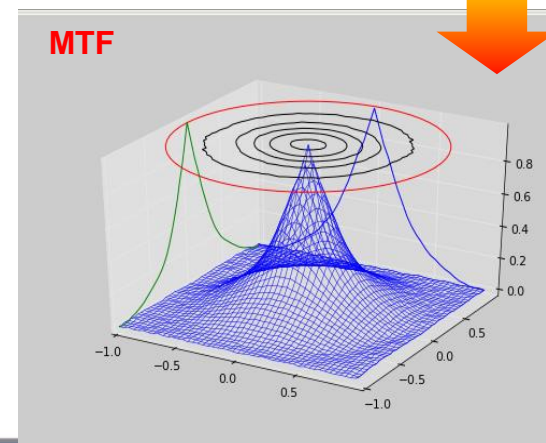
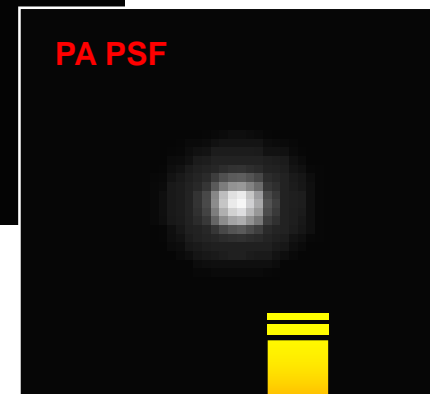
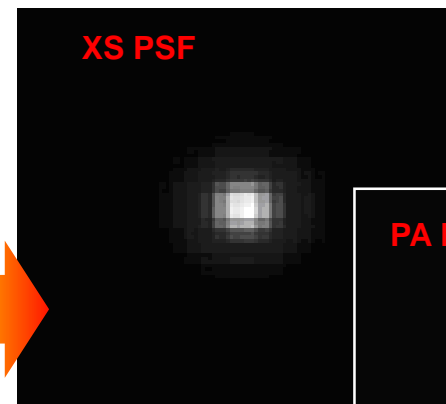


...and the minutes
needle moving!

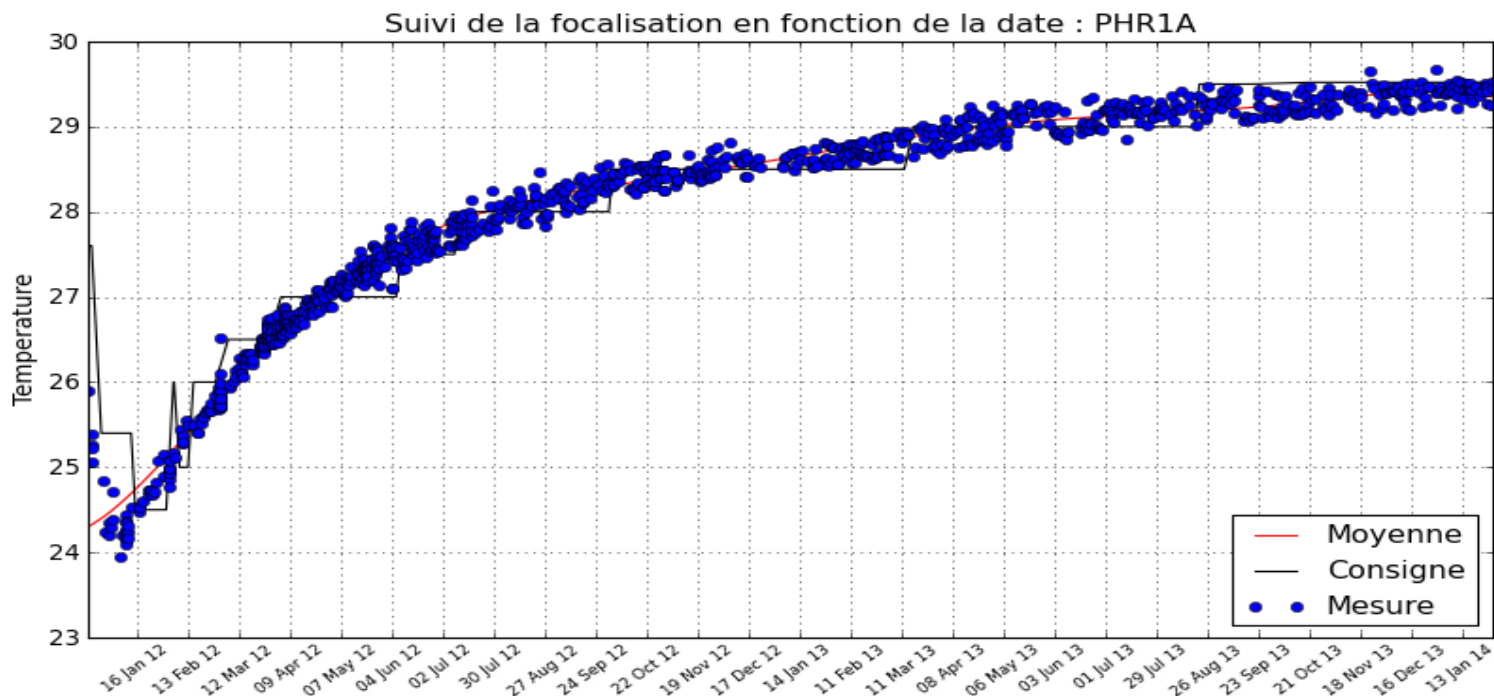
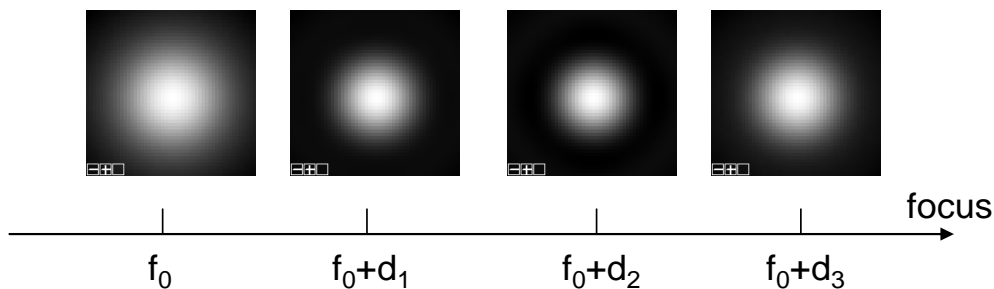
Viewing the Stars for MTF Assessment and Refocusing (1)



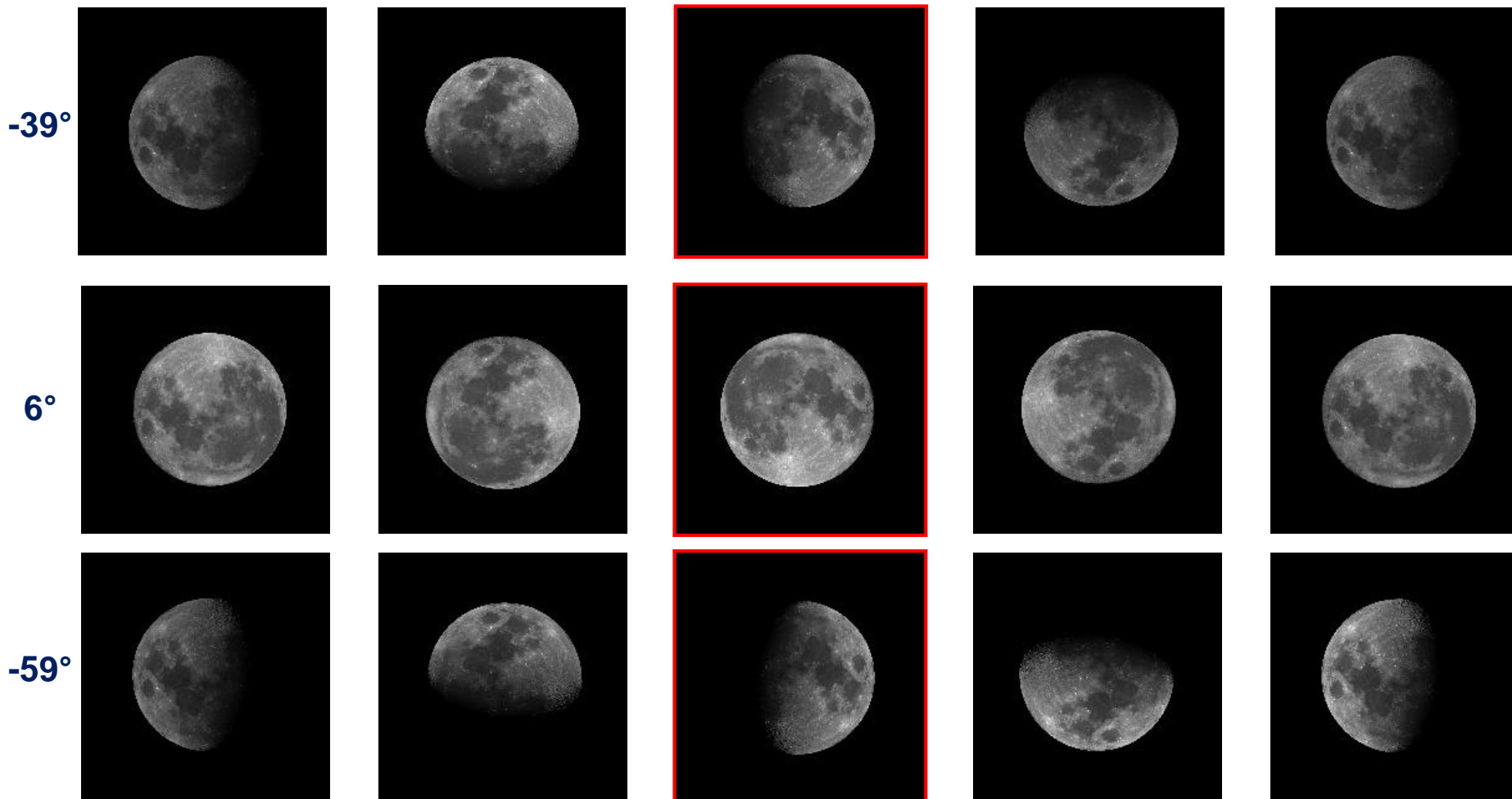
- Few images are needed
- No atmosphere
- Good accuracy
- 2D MTF assesment



Monitoring of the best focus

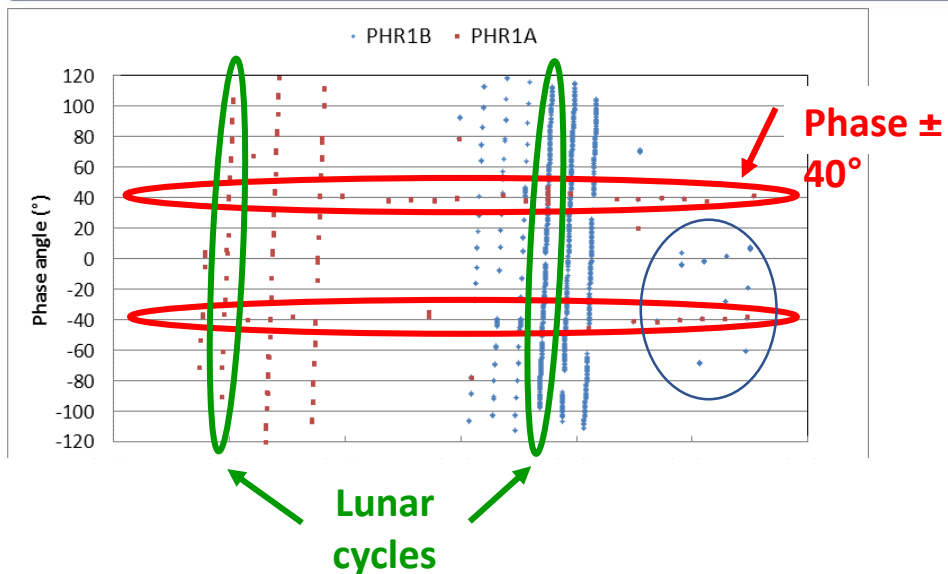


Viewing the Moon for Radiometric Calibration (1)

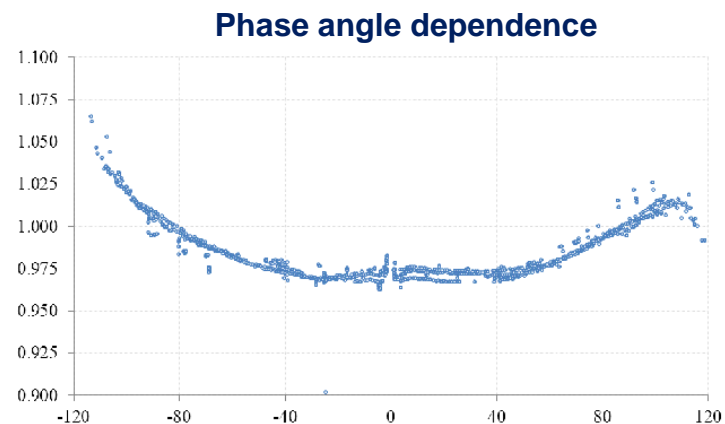
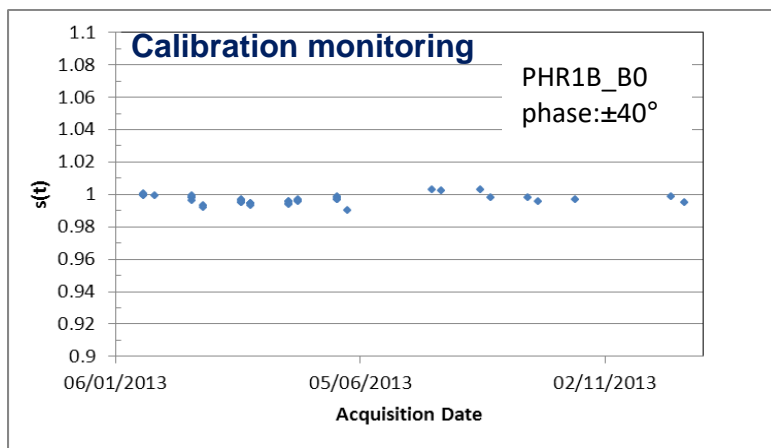


**PLEIADES
dataset :
more than
1200 images
of the Moon
with multiple
phase angles
and
orientations**

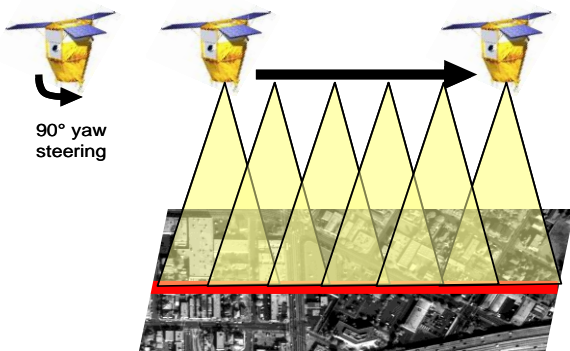
Viewing the Moon for Radiometric Calibration (2)



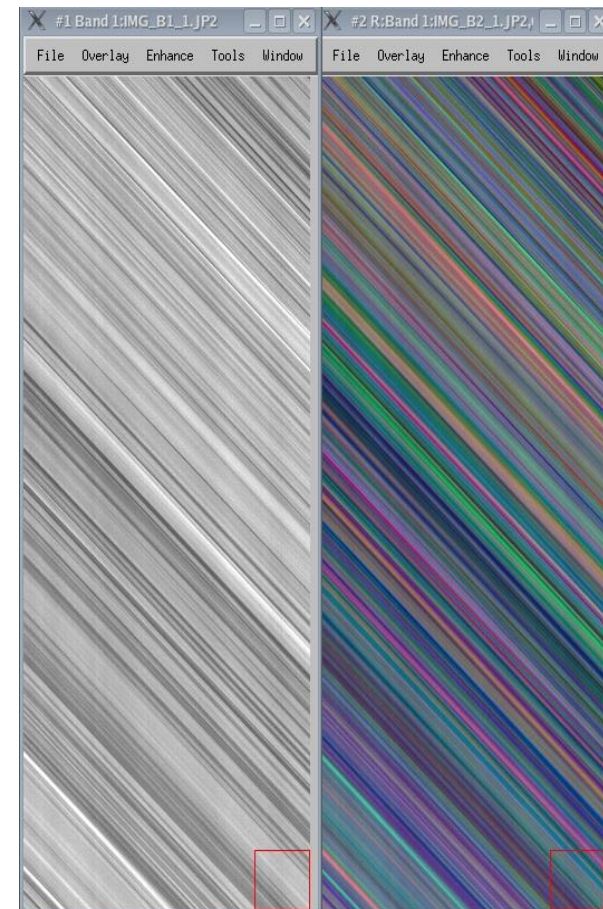
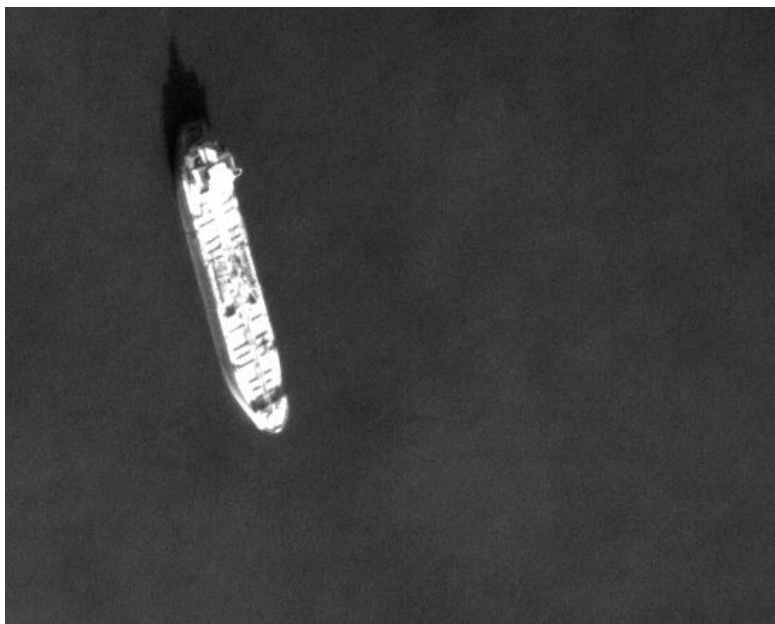
PLEIADES Moon dataset used to enhance the Lunar radiometric calibration model (ROLO)



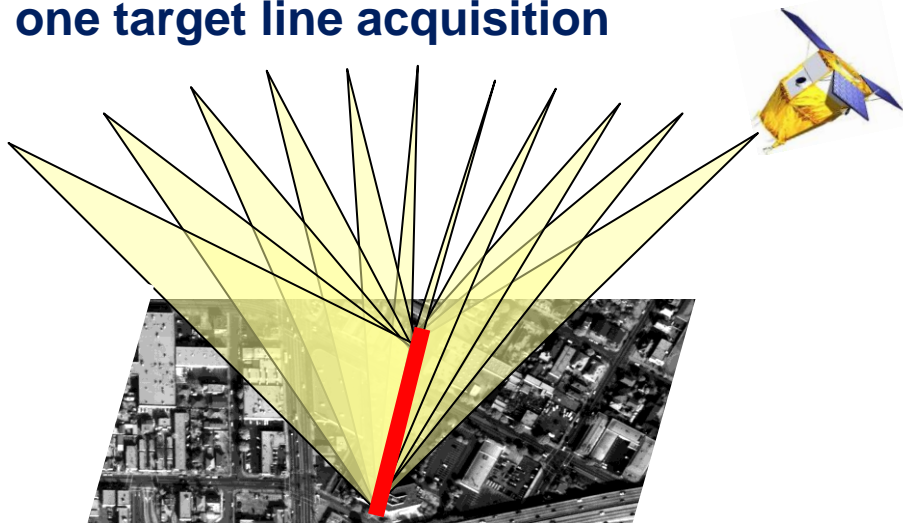
90° Rotation of the Satellite for Uniformity Calibration



AMETHIST Mode:
each pixel successively
views the same landscape



Steady Mode:
one target line acquisition

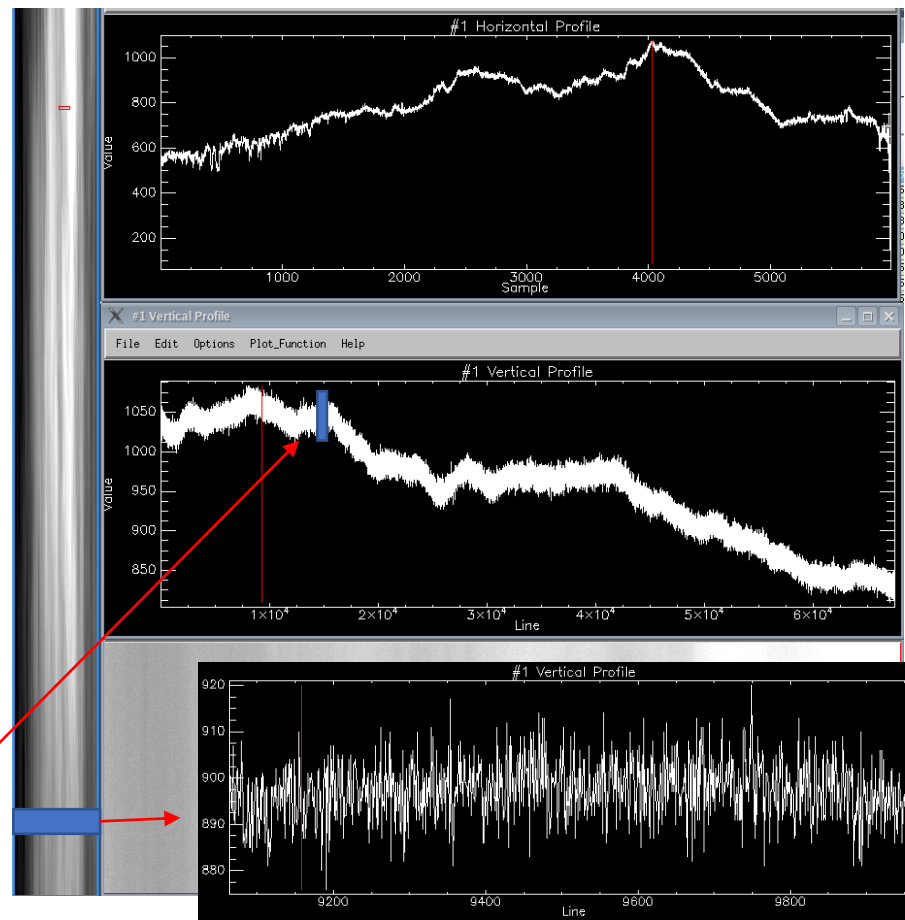
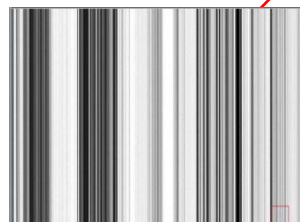


Acquisition duration

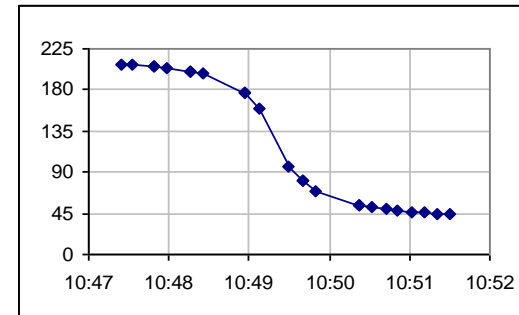
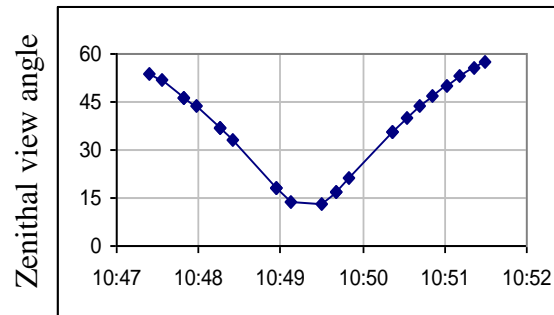
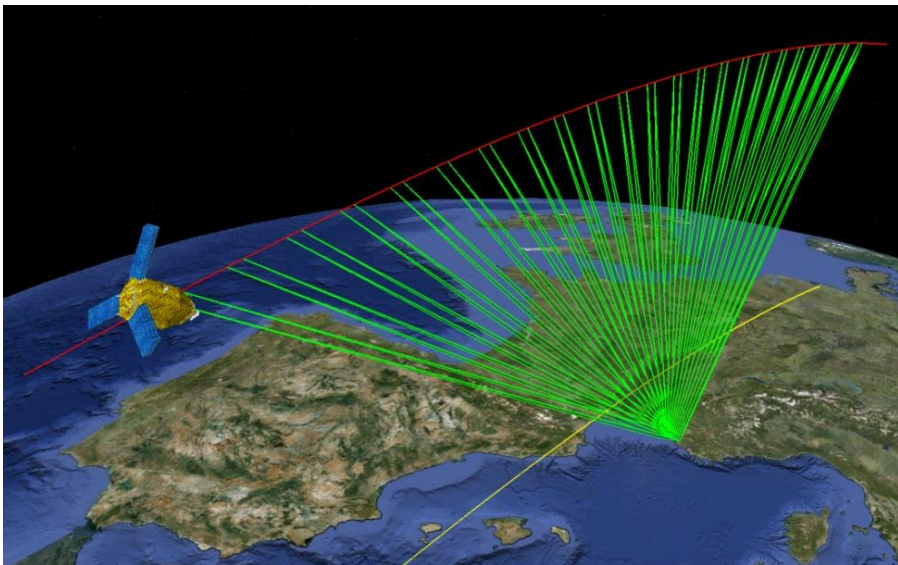
- *a few seconds*
- *about 60000 rows*

Relative good line-of-sight stability

- *use of 1000 rows*



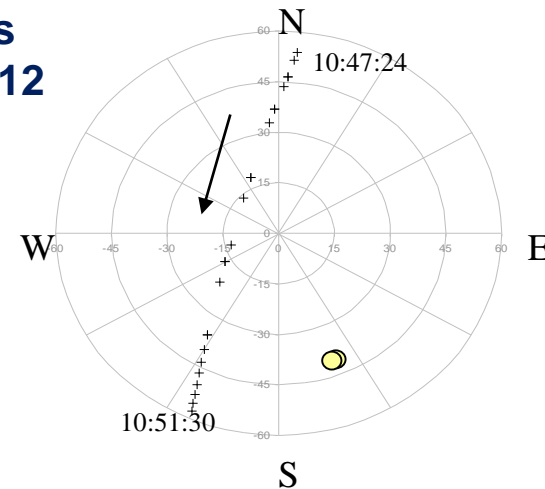
Vertical profile of 1 column and for 1000 rows

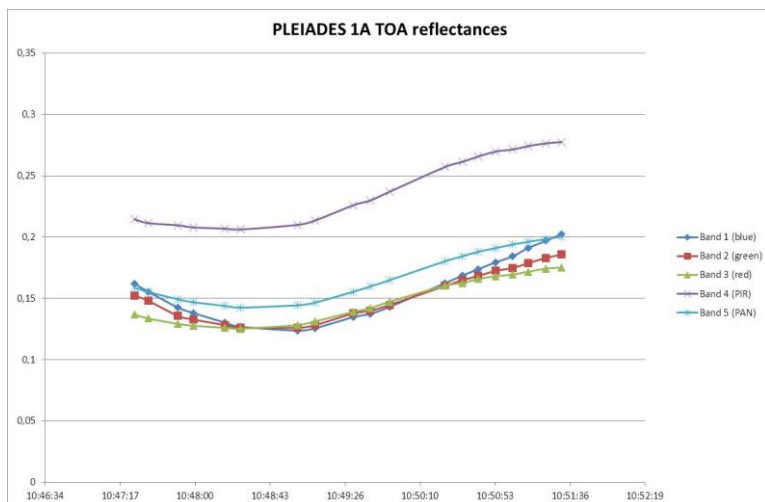
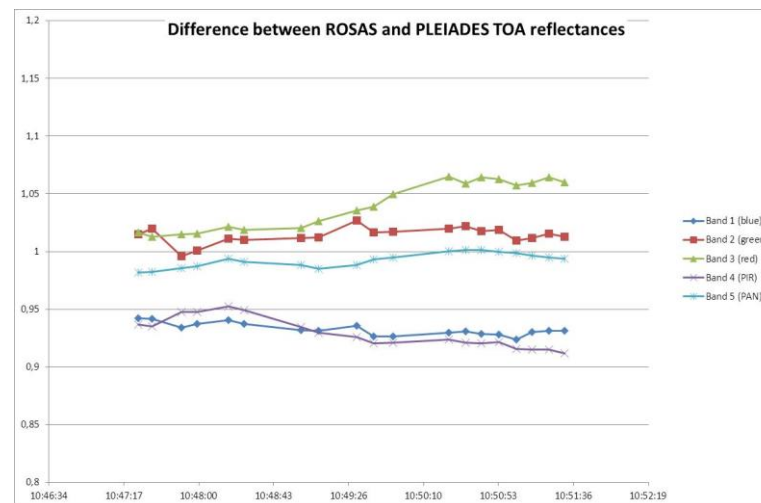
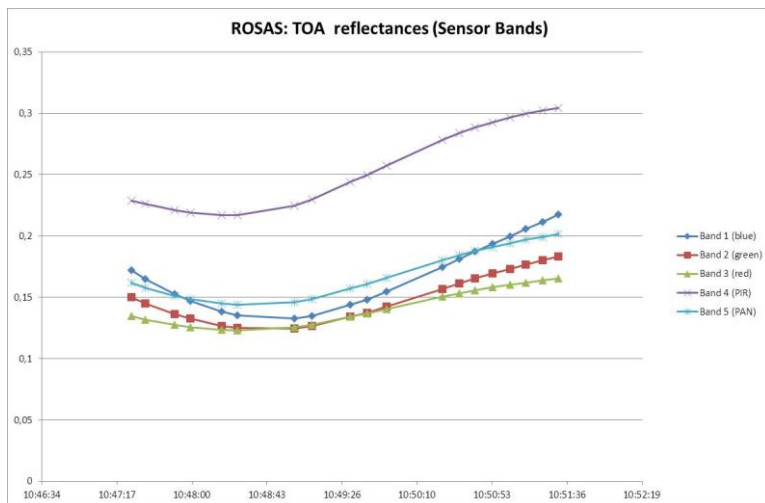


Relative azimuth angle

PHR1-A Video acquisitions over La Crau - March 28, 2012

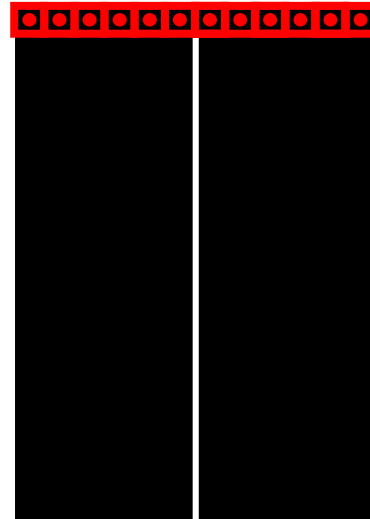
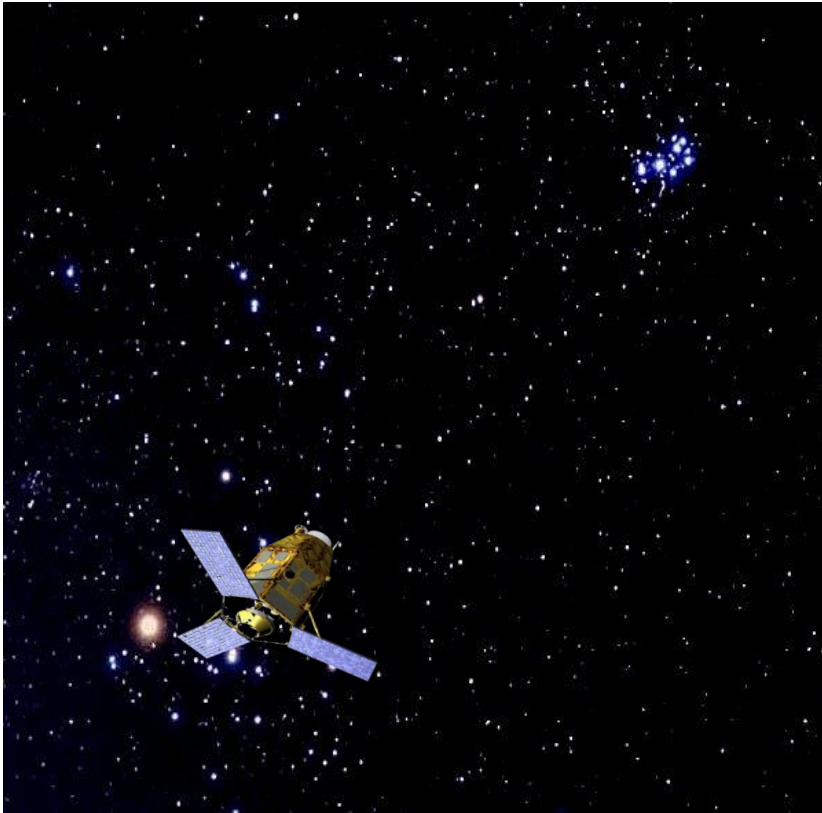
19 measurements
 4 min. total duration
 $\theta_s \approx 42$ deg
 $\theta_v \text{ min} = 13.2$ deg
 $\theta_v \text{ max} = 57.7$ deg



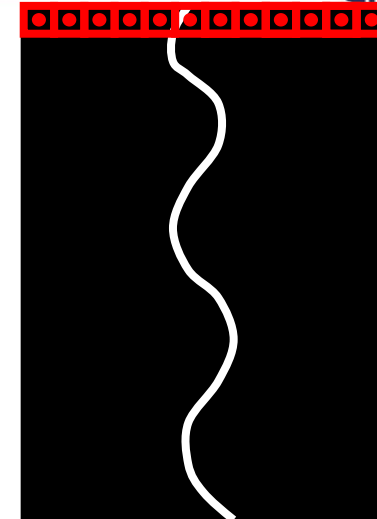


BRDF modelling and correction

Bands	Ak ROSAS / Ak Ground (mean)	Standard deviation	Max-Min	Ak Official / Ak Ground	Ak ROSAS / Ak Official
B1 (blue)	0.92	0.0053	0.0183	1.05	12.5%
B2 (green)	1	0.0060	0.0308	0.95	5%
B3 (red)	1.025	0.0031	0.0522	0.96	6.5%
B4 (PIR)	0.915	0.0029	0.0406	0.94	2,5%
B5 (PAN)	0.975	0.0045	0.0198	0.95	2,5%

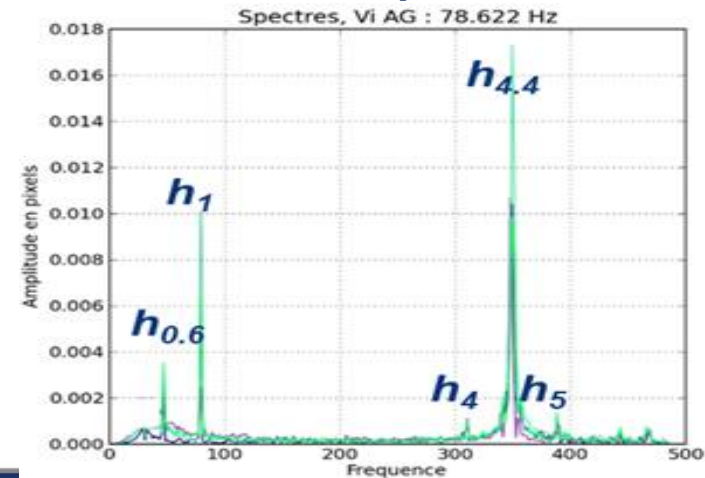


without μ vib

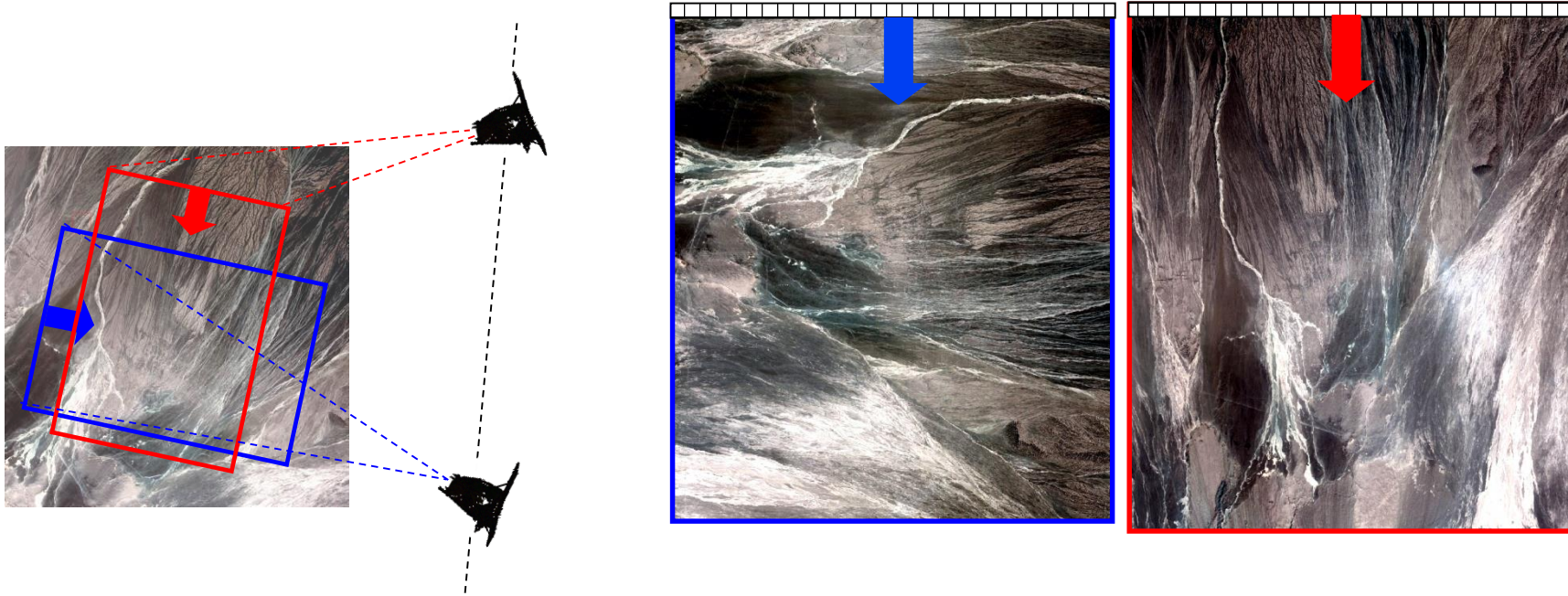


with μ vib

- 👉 Microvibration amplitude < 0.02 pixels
- 👉 Accuracy: few nrad



Use of images taken in a crosswise mode (LOS orientation variation = 90°) to separate dynamic and static phenomenon which are specific to the push-broom acquisition principle



Crosswise Viewing for Focal Plane Cartography (2)

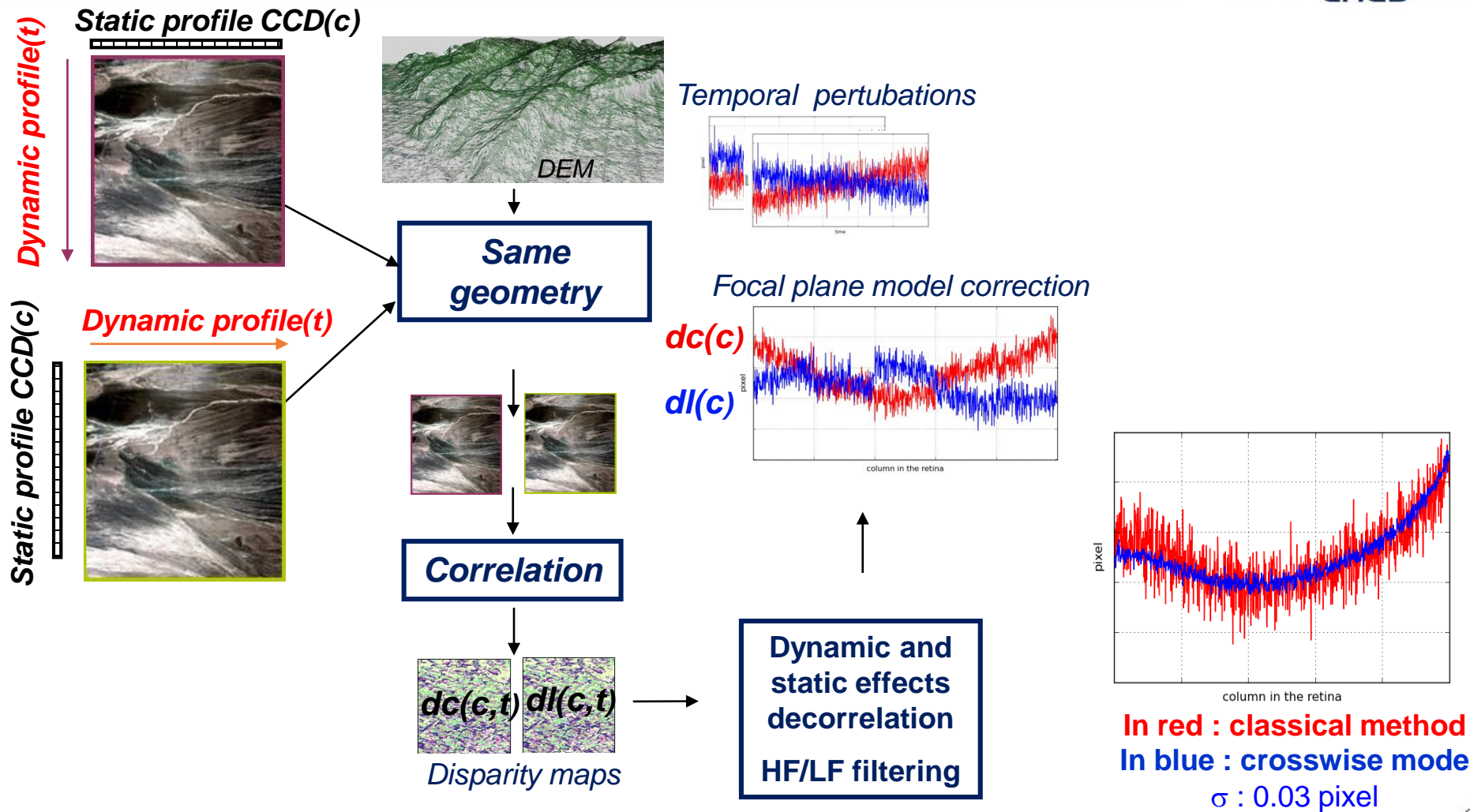


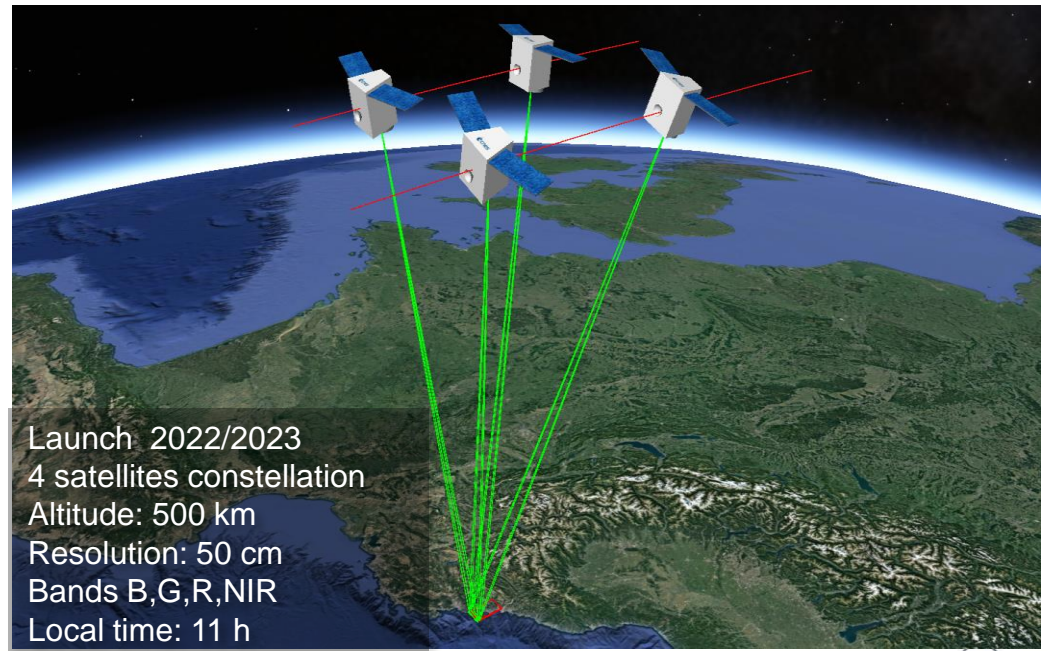
Image Quality is considered at CNES as a specific skill

- Sharing experience from one mission to other ones
- Knowledge transmission ensured thanks to a dedicated organization
- 3 generations of 'Image Quality' experts since the 80's (SPOT1) !!!

Ensuring high Image Performances and Efficiency

Next one is : CO3D !!!

- 3D global coverage in 5 years
- Relative altimetric accuracy: 1 m (B/H~0,3)
- Synchronous acquisition per satellites pair
- Capacity for urban quadri-stereo every 2 days
- Massive 3D production based on cloud facilities









Thank you
for your attention