HAROKOPIO UNIVERSITY

→ 4th ADVANCED TRAINING COURSE IN LAND REMOTE SENSING

SAR Interferometry: Theory

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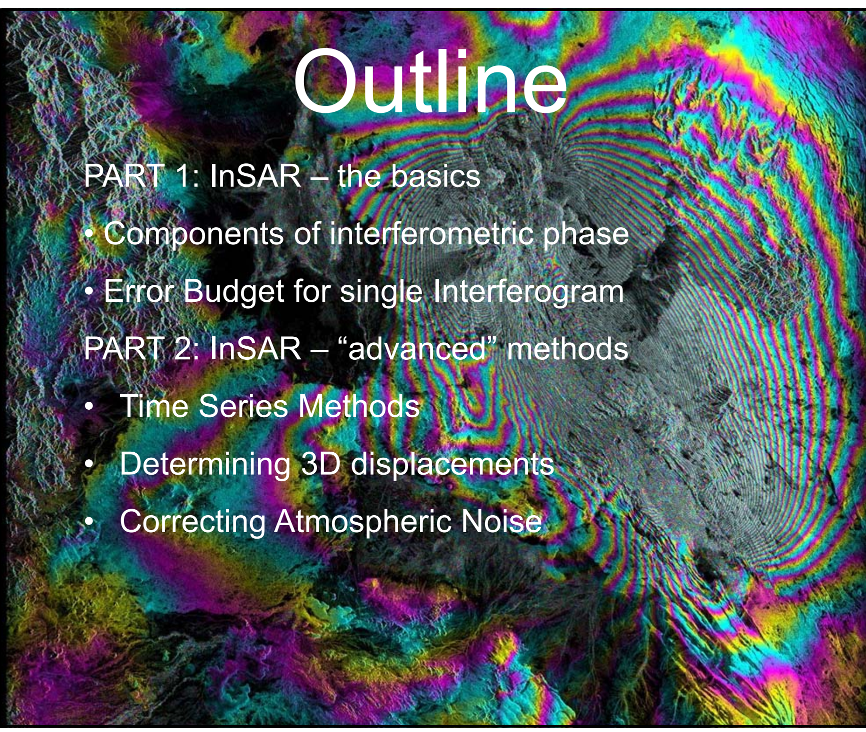


Outline

PART 1: InSAR – the basics

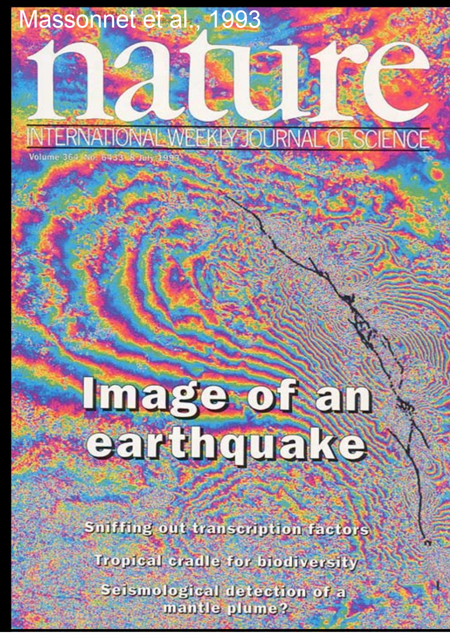
- Components of interferometric phase
- Error Budget for single Interferogram

PART 2: InSAR – "advanced" methods

- Time Series Methods
- Determining 3D displacements
- Correcting Atmospheric Noise

InSAR – how it works

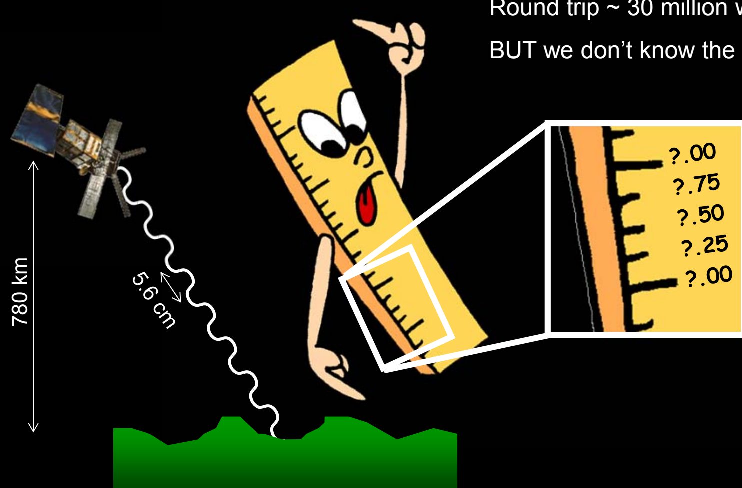
- Actively illuminate ground with radar waves.
- Operates day and night, can see through clouds
- ERS, Envisat (1991): very stable orbits and pointing ⇨ InSAR
- Followed by ERS-2 (1995) and Envisat (2003) for ~ 20 year time series



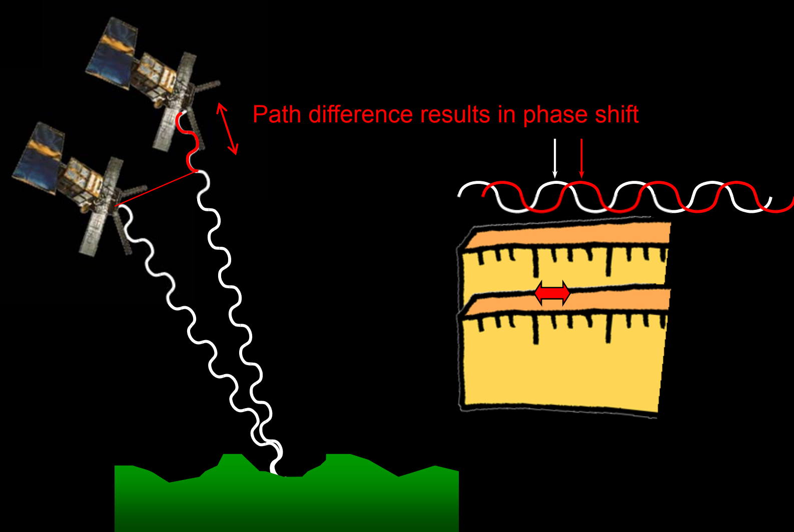
InSAR – how it works

Round trip ~ 30 million wavelengths

BUT we don't know the exact number



InSAR – how it works



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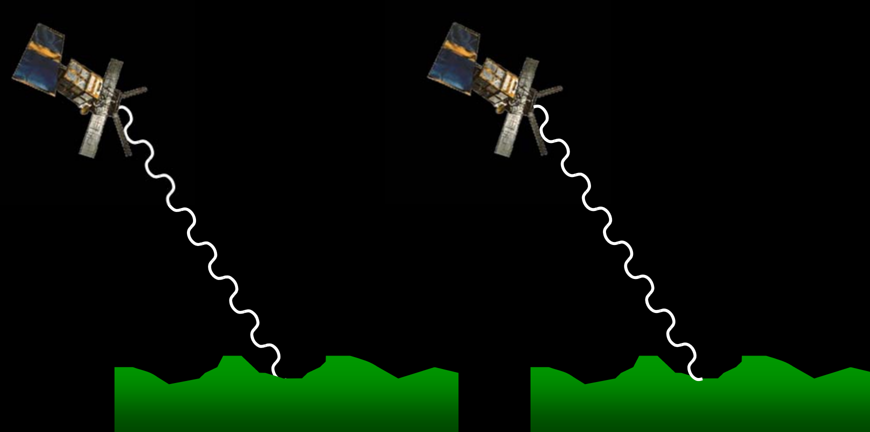
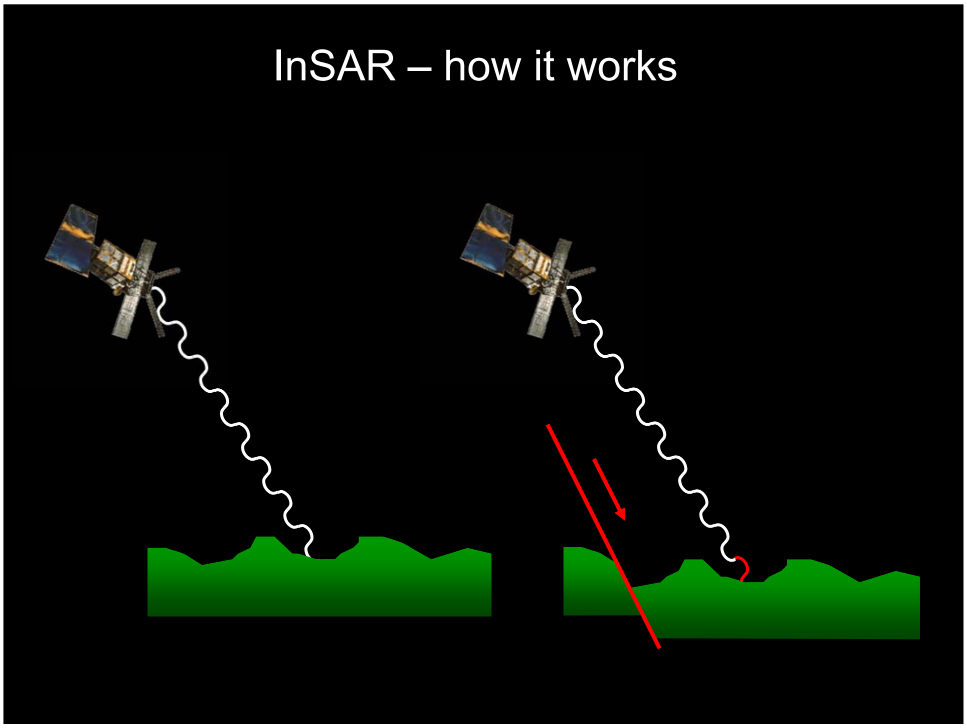



InSAR – how it works

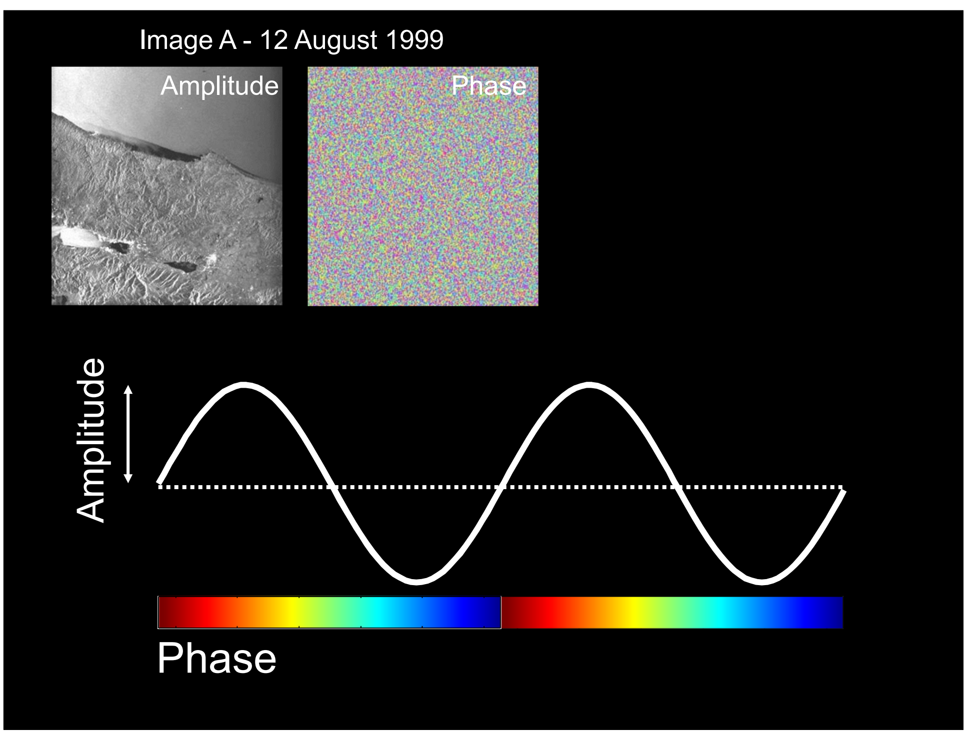


Image A - 12 August 1999
Amplitude
Phase
Amplitude
Phase

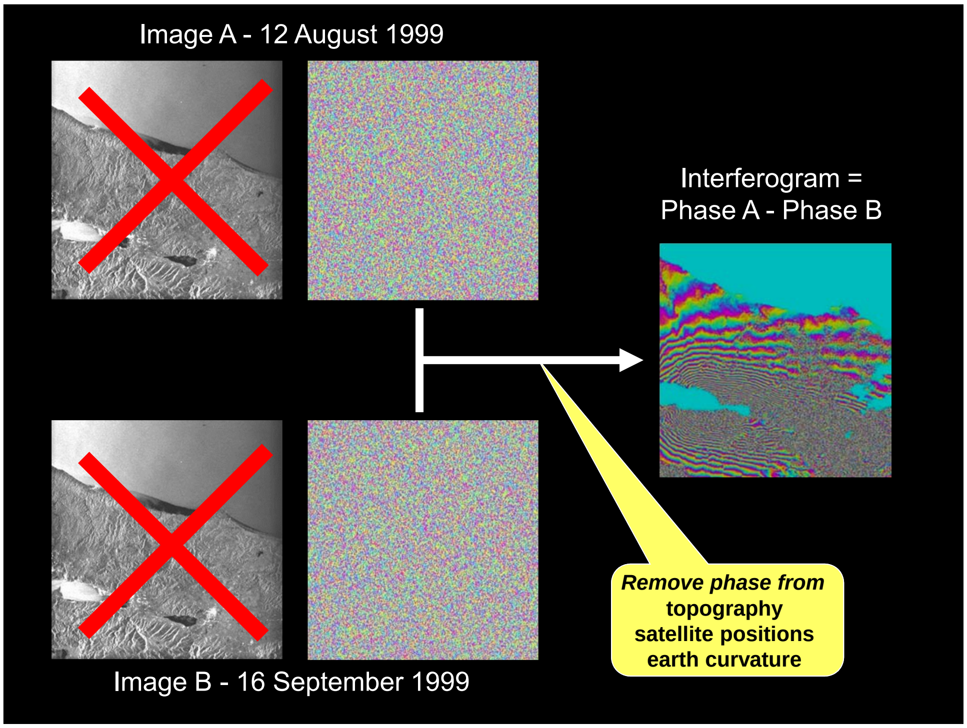
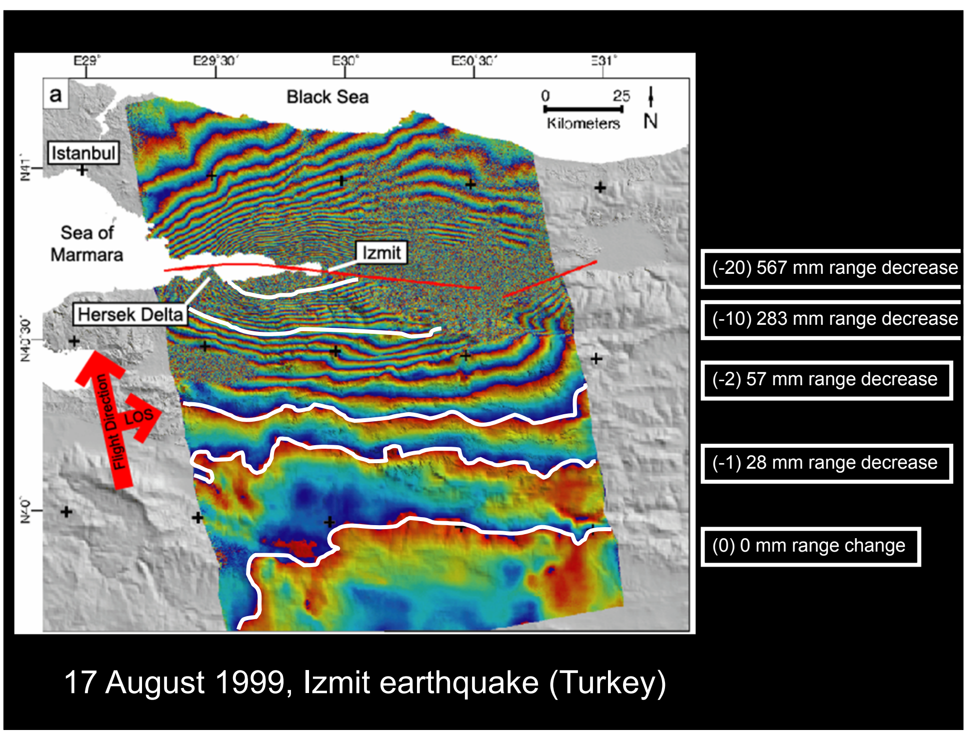
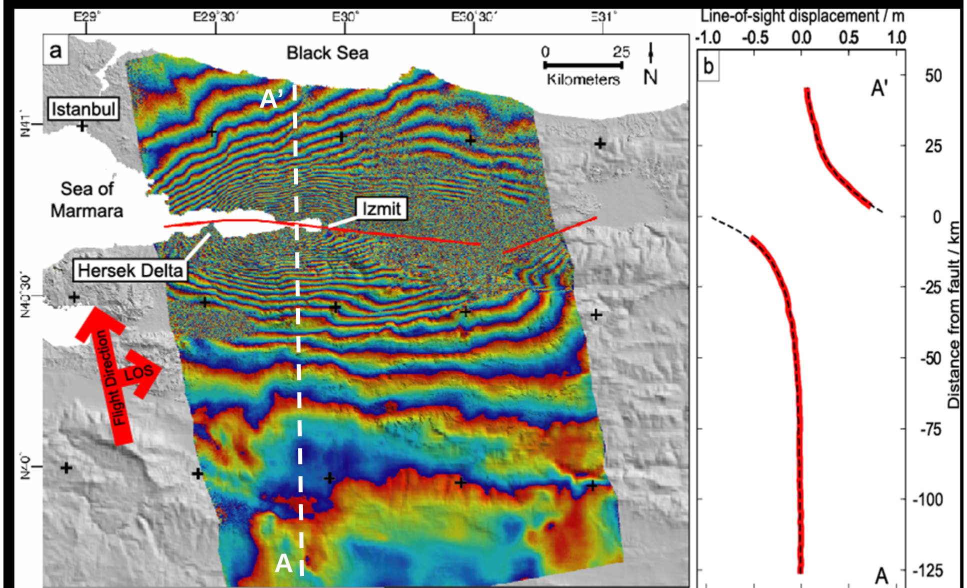


Image A - 12 August 1999
Interferogram =
Phase A - Phase B
Remove phase from
topography
satellite positions
earth curvature
Image B - 16 September 1999


a
Black Sea
0 25
Kilometers
N
Istanbul
Sea of
Marmara
Izmit
Hersek Delta
Flight Direction
LOS
(-20) 567 mm range decrease
(-10) 283 mm range decrease
(-2) 57 mm range decrease
(-1) 28 mm range decrease
(0) 0 mm range change
17 August 1999, Izmit earthquake (Turkey)



17 August 1999, Izmit earthquake (Turkey)

Components of interferometric phase

$$\boxed{\Delta\phi_{int} = \cancel{\Delta\phi_{geom}} + \cancel{\Delta\phi_{topo}} + \cancel{\Delta\phi_{atm}} + \cancel{\Delta\phi_{noise}} + \Delta\phi_{def}}$$

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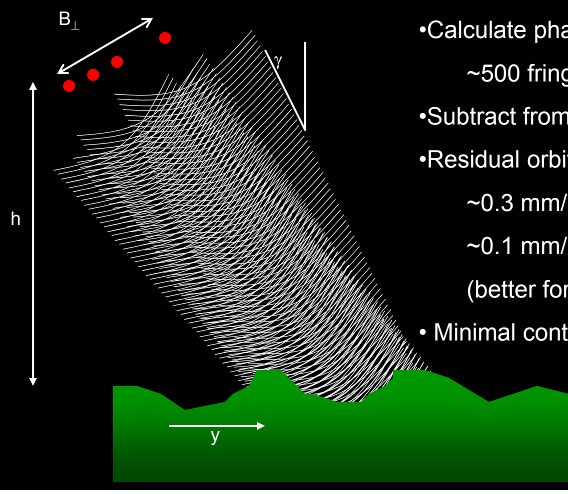

- Calculate phase ramp from satellite orbits
  - ~500 fringes across typical frame
- Subtract from interferogram
- Residual orbital errors:
  - ~0.3 mm/km (north, ERS)
  - ~0.1 mm/km (east, ERS)
  - (better for Envisat)
- Minimal control on v. long wavelength

Components of interferometric phase

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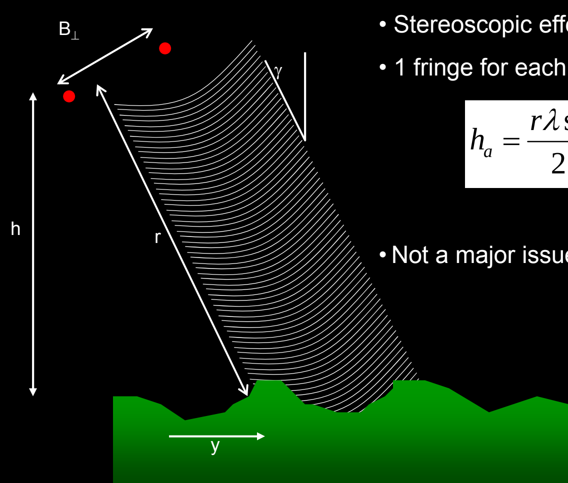


- Stereoscopic effect ⇒ topographic fringes
- 1 fringe for each change in elevation h_a

$$h_a = \frac{r\lambda\sin\gamma}{2B_\perp} \approx \frac{10{,}000}{B_\perp}$$

- Not a major issue since SRTM

Components of interferometric phase

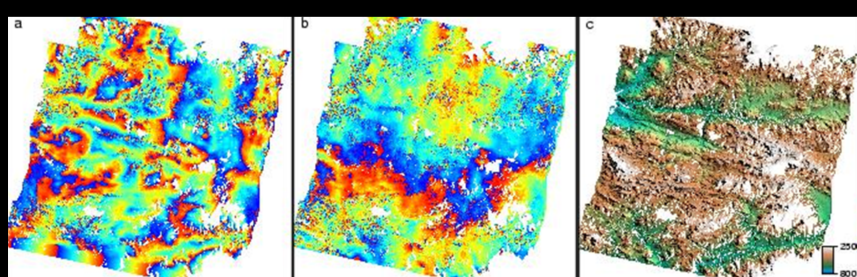
$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{geom}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{noise}} + \Delta\phi_{\text{def}}$$

A foggy morning, near ancient Mycenae, Greece

Components of interferometric phase

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{geom}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{noise}} + \Delta\phi_{\text{def}}$$

Layered atmosphere

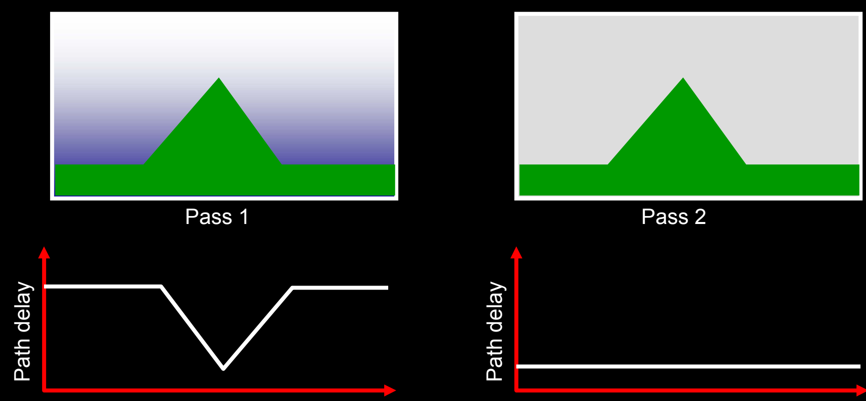


29/8/1995 to 29/7/1997 | 30/8/1995 to 29/7/1997 | Topography

Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

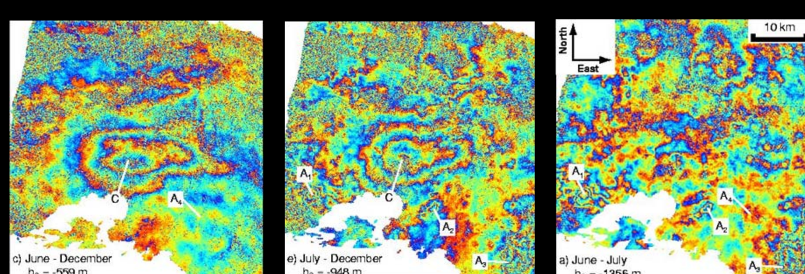
Layered atmosphere



Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

Turbulent atmosphere



June to December July to December June to July

Athens Earthquake – September 1999

Components of interferometric phase

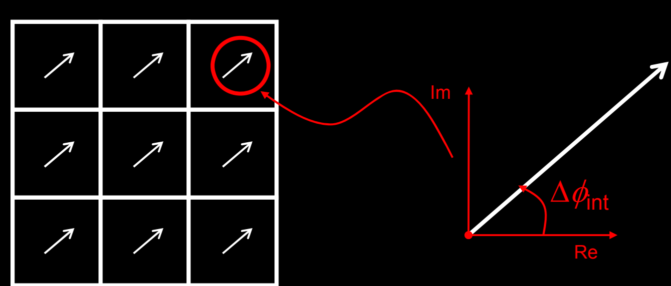
$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

- Size of $\Delta\phi_{atm}$ (at sea level) scales with distance, but can be +/- 10 cm or more.
- Methods for dealing with $\Delta\phi_{atm}$
  - Ignore (most common)
  - Quantify
  - Model based on other observations (e.g. GPS, meteorology…)
  - Increase SNR by stacking or time series analysis

Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

- Biggest source of noise is due to changing ground surface
- *Coherence* is convenient measure



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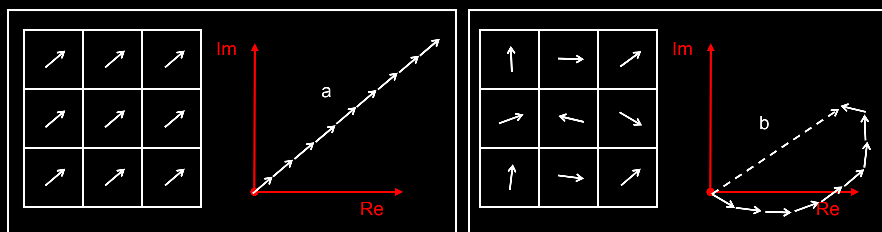
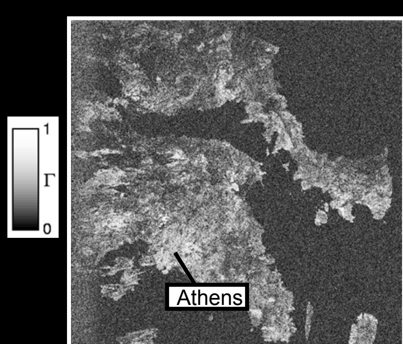


Coherence = b / a

Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$



Coherent surface types
- Bare Rock
- Buildings esp. towns/cities

- Grassland
- Agricultural fields
- Ice

Incoherent surface types
- Leafy Trees
- Water

Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

1. incoherence

- Changes in the ground cover cause a random phase shift for each pixel
- Large baselines

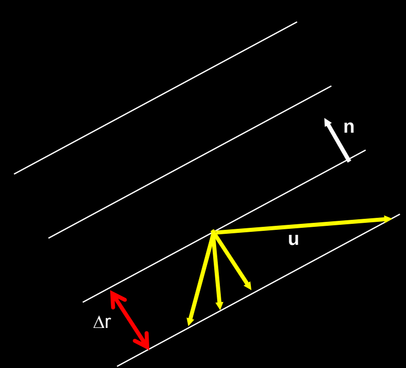
2. Unwrapping errrors

- Phase in interferograms is wrapped (each fringe is 2 π radians).
- Discontinuities or data gaps can cause phase unwrapping errors

Components of interferometric phase

$$\Delta\phi_{int} = \Delta\phi_{geom} + \Delta\phi_{topo} + \Delta\phi_{atm} + \Delta\phi_{noise} + \Delta\phi_{def}$$

INSAR ONLY MEASURES THE COMPONENT OF SURFACE DEFORMATION IN THE SATELLITE'S LINE OF SIGHT



$$\Delta r = - \mathbf{n} \cdot \mathbf{u}$$

where n is a unit vector pointing from the ground to the satellite

$$\Delta\phi_{def} = (4\pi / \lambda)\, \Delta r$$

i.e. 1 fringe = 28.3 mm l.o.s. deformation for ERS

Error Budget (1)
Single interferogram

$$\sigma_{def}^2 = \sigma_{gm}^2 + \sigma_{topo}^2 + \sigma_{atm}^2 + \sigma_{coh}^2 + \sigma_{sys}^2 + \sigma_{unw}^2$$

- Orbital errors ⇨ long-wavelength ramps.
- Envisat: ~0.3 mm/km (across-track) and 0.1 mm/km (along-track) [Wang, Wright and Biggs, GRL 2009].
- Can correct by processing long strips and tying to GPS **(see. Fringe presentations by Wang, Pagli and Hamlyn)**
- Should be negligible for future missions with onboard GPS receivers.

Error Budget (1)
Single interferogram

$$\sigma_{def}^2 = \sigma_{gm}^2 + \sigma_{topo}^2 + \sigma_{atm}^2 + \sigma_{coh}^2 + \sigma_{sys}^2 + \sigma_{unw}^2$$

$$\sigma_{topo} = \frac{\bar{r}_{slant} B_{\perp}}{\sin \theta_{inc}} \sigma_{DEM}$$

- SRTM error ~ 4 m absolute, of which 2.5 m is not spatially correlated [Rodriguez et al., PERS 2006]

B_{perp}	σ_{topo} (40° incidence)
150 m	1.1 mm
300 m	2.3 mm
1000 m	7.8 mm

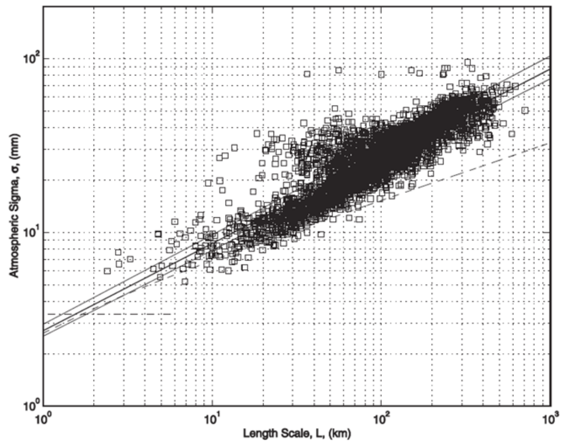
Error Budget (1)
Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

- Troposhere

Emardson et al., 2003:
$\sigma = cL^{\alpha}$ [c~2.5, α~0.5]
σ= 25 mm at 100 km

(assume no corrections)



Error Budget (1)
Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

- Ionosphere ($1/f^2$ dependence). Important at L-band, but not at C-band.
- Can correct with split band processing (e.g. 1200 and 1260 MHz) in future missions
- Ionospheric error on 100 km wavelength ~ 1mm after spatial averaging

Error Budget (1)
Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

- Coherence, γ
  - important at short wavelengths, but can be averaged through multilooking to < 1 mm for most ground cover types

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Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

- Coherence, γ
  - important at short wavelengths, but can be averaged through multilooking to < 1 mm for most ground cover types
- System (thermal) - modifies coherence
  - reduces effective coherence, but still insignificant after spatial averaging.

$$\sigma_{coh} = \left(\frac{\lambda}{4\pi}\right)\frac{1}{\sqrt{N_L}}\frac{\sqrt{1-\gamma^2}}{\gamma}$$

$$\gamma_c = \frac{\gamma}{1+SNR^{-1}}$$

Error Budget (1)
Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

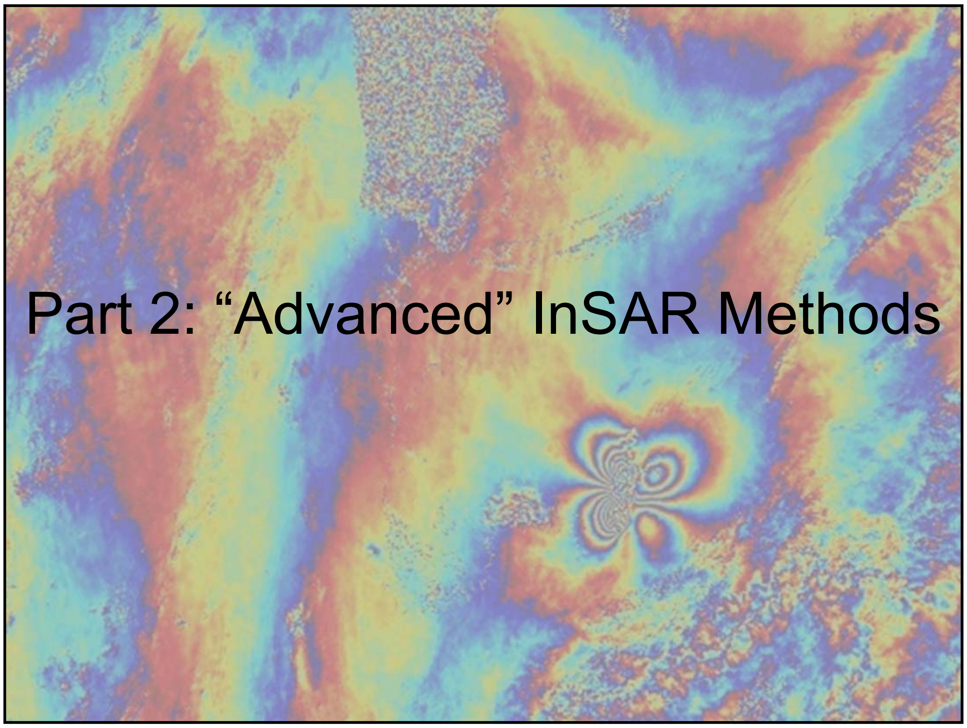
- Unwrapping errors difficult to quantify.
- Assume = 0 in this analysis (probably OK for L-band missions or missions with short revisits).

Error Budget (1)
Single interferogram

$$\sigma^2_{def} = \sigma^2_{gm} + \sigma^2_{topo} + \sigma^2_{atm} + \sigma^2_{coh} + \sigma^2_{sys} + \sigma^2_{unw}$$

Atmospheric (tropospheric) error dominates at 100 km length scales, at which single interferograms have error of ~25 mm.

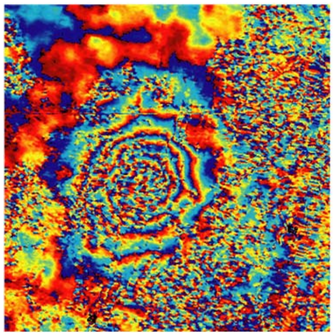
Part 2: “Advanced” InSAR Methods



Outline for Advanced Methods

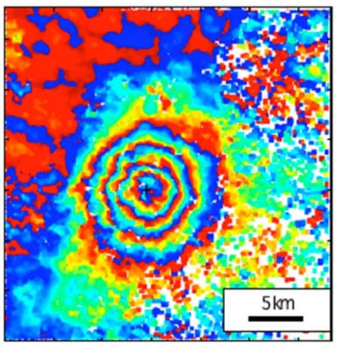
1. Combining interferograms
   - Stacking
   - Time series
   - SBAS/Permanent Scatterers
   - Error budget for Time Series Methods
2. Determining 3D displacements/velocities
   - Direct inversion
   - Combination with GPS
3. Atmospheric Corrections
   - Linear/Smooth Velocity Assumption
   - MERIS/MODIS
   - GPS
   - Weather Models

Stacking



Individual Interferogram

Typical atmospheric noise for individual interferogram ~ 1cm



Stack of 5 images

Stack: Add together 5 interferograms

Signal increases by a factor of 5

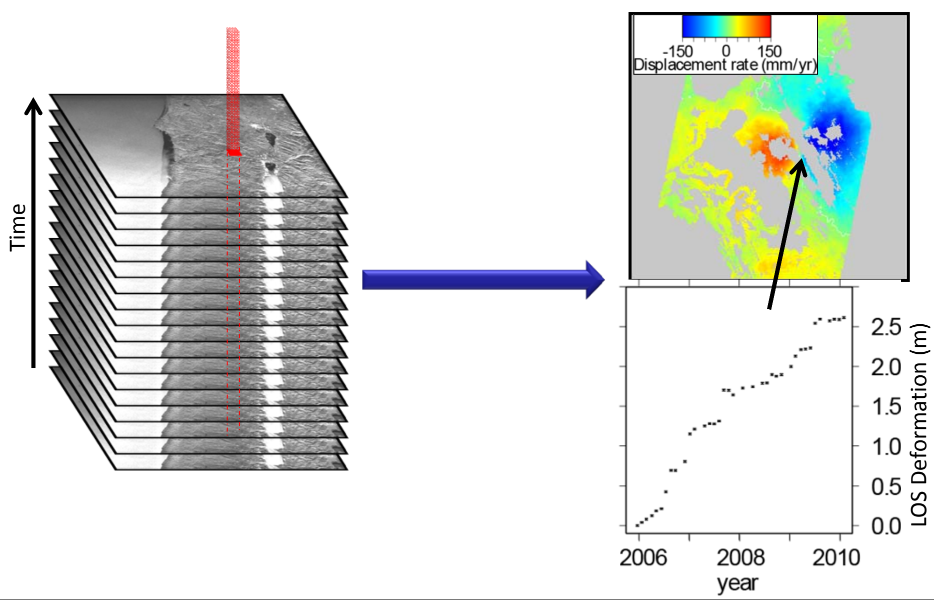
Noise increases by a factor of √5

Signal:Noise ratio increases by 5/√5 = √5 ~ 2.23

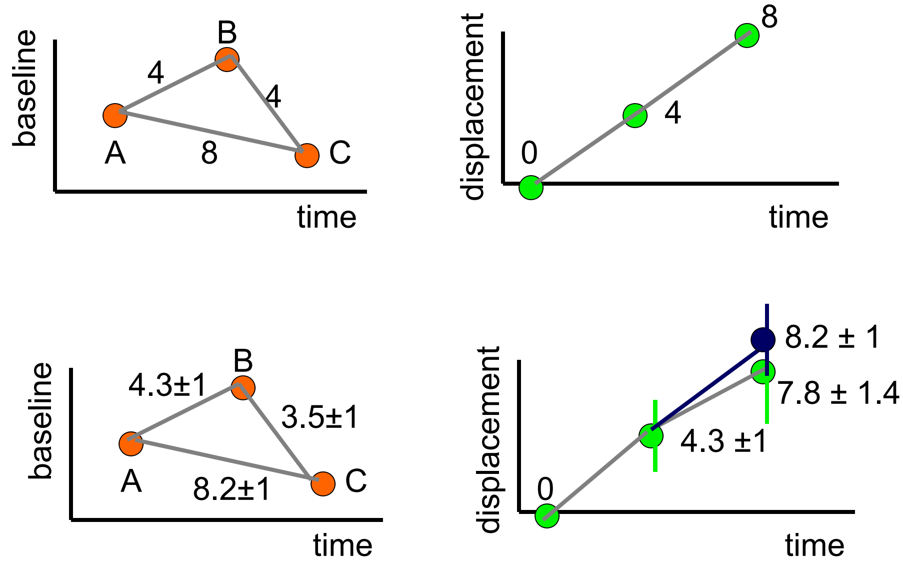
For continuous phenomena (e.g. interseismic strain) or discrete events (e.g earthquakes)

Biggs et al, 2009 (Geology)

All time series methods are essentially the same – rely on large stacks of imagery to separate signal from noise

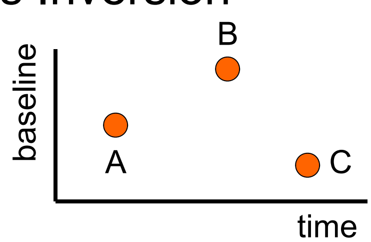


Time Series Example



Time Series Inversion

Acquisitions A,B,C



$i_{AB}=d_B-d_A \quad = v_{AB}t_{AB}$

$i_{BC}=d_C-d_B \quad = v_{BC}t_{BC}$

$i_{AC}=d_C-d_A \quad = (d_C-d_B) + (d_B-d_A) \quad = v_{AB}t_{AB} + v_{BC}t_{BC}$

$$\begin{bmatrix} t_{AB} & 0 \\ 0 & t_{BC} \\ \cancel{t_{AB}} & \cancel{t_{BC}} \end{bmatrix} \begin{bmatrix} v_{AB} \\ v_{BC} \end{bmatrix} = \begin{bmatrix} i_{AB} \\ i_{BC} \\ \cancel{i_{AC}} \end{bmatrix}$$

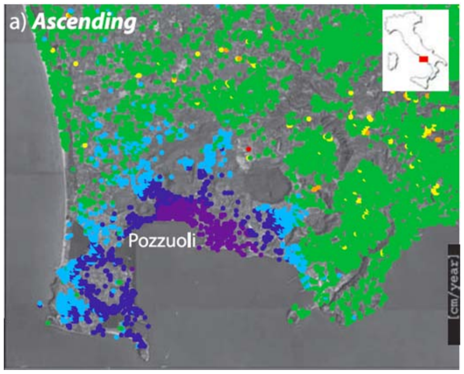
$\boxed{G_{INS}\, m = d_{INS}}$

To get correct answer with this method, weighting with covariances is essential

$\boxed{\Sigma^{-1}G_{INS}\, m = \Sigma^{-1}d_{INS}}$

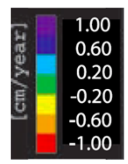
SBAS: Short BAseline Subset

Example: Campi Flegrei caldera (Italy).



30 ascending images
=> 180 interferograms

Max uplift of 2 cm/yr in Pozzuoli Harbour

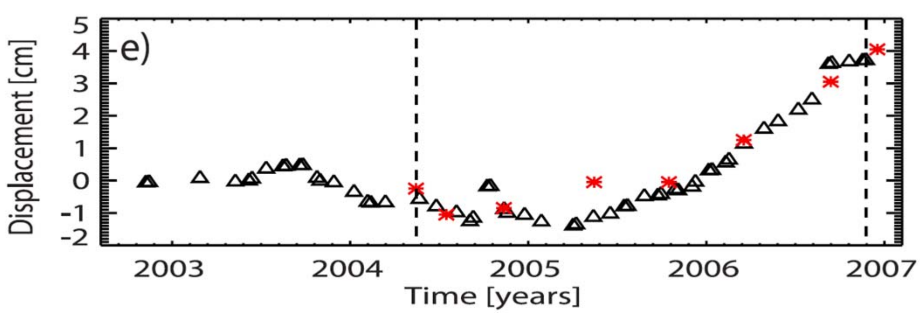


Modelled by an inflation rate of a magma chamber at a depth of 3.2 km with a volume change of 1.1 x 10^6 m^3/yr

(Trasatti et al, 2008; Casu et al, 2006)

SBAS

Pozzuoli Harbour time series:

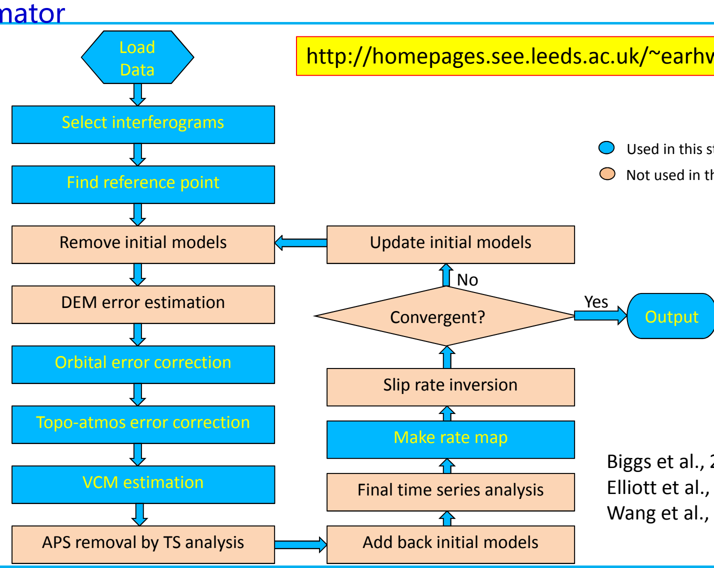


Stated accuracy: 1 mm/yr in rate. 5 mm in displacement.

Good match with levelling data (red).

PI-RATE: Poly-Interferogram Rate And Time-series Estimator

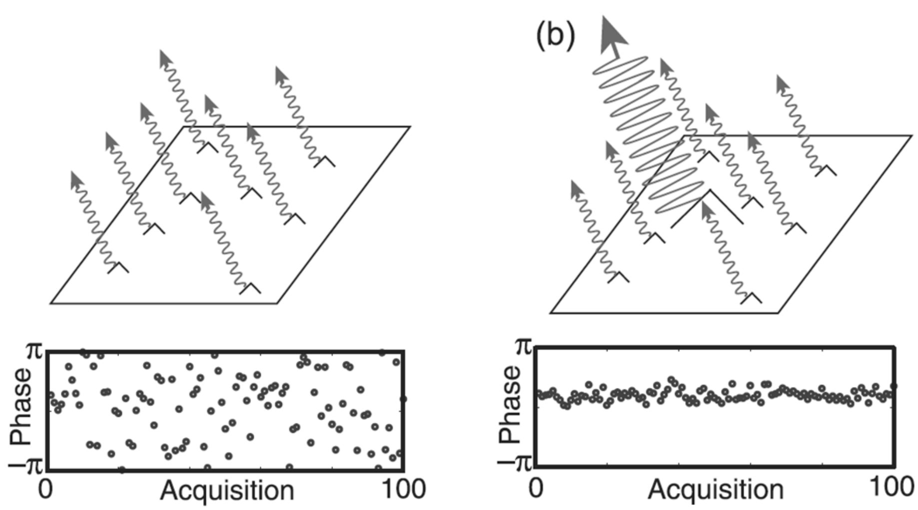
http://homepages.see.leeds.ac.uk/~earhw/pirate

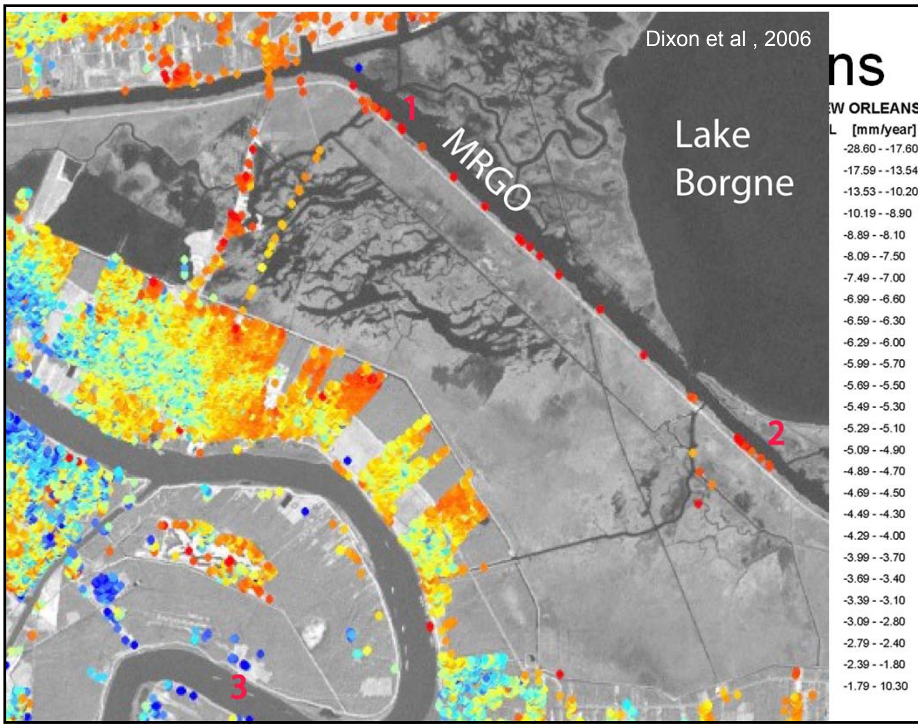


Used in this study
Not used in this study

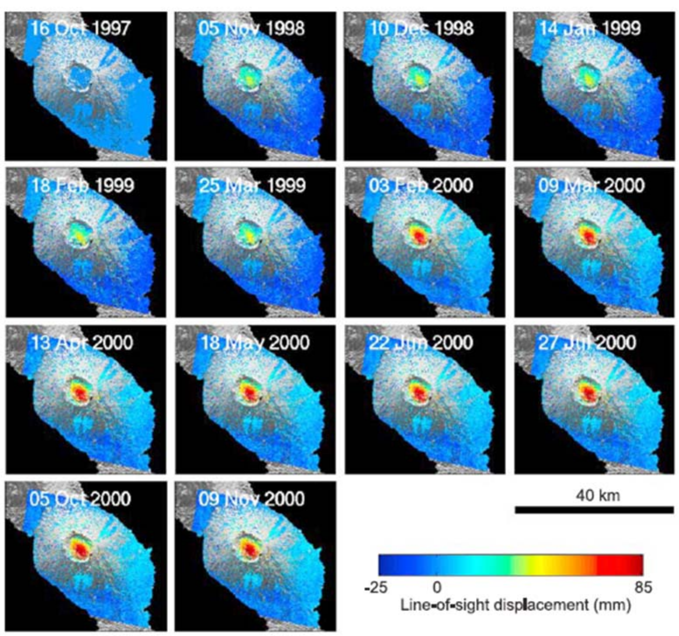
Biggs et al., 2007, GJI;
Elliott et al., 2008, GRL;
Wang et al., 2009, GRL.

PS InSAR:





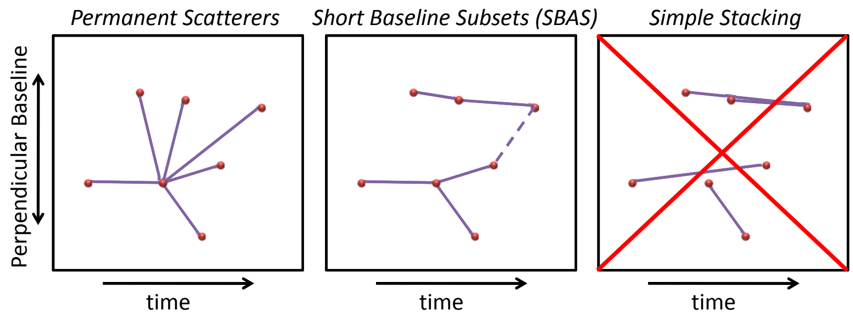
STAMPS: Volcan Alcedo, Galapagos



Hooper et al, 2007

Error Budget (2)

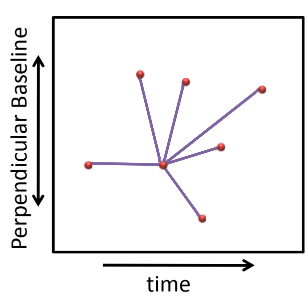
Optimum determination of Linear Deformation Rates



For the determination of linear deformation rates, optimum errors are determined through a connected network, since noise terms are associated with individual acquisitions not interferograms.

Error Budget (2)

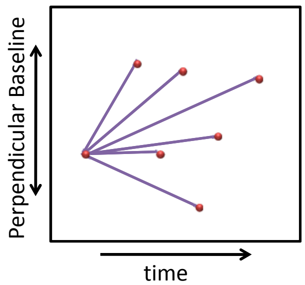
Optimum determination of Linear Deformation Rates



- Error on linear rate is independent of how network is connected (but of course short-baseline, short-time interferograms are best).

Error Budget (2)

Optimum determination of Linear Deformation Rates



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• To simplify mathematics, assume all connections to date d1...

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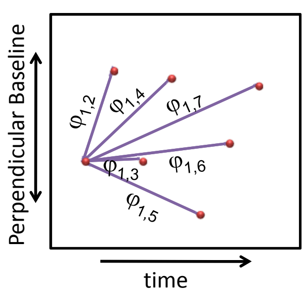

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• To simplify mathematics, assume all connections to date d1...
...and regular acquisition spacing, t_m

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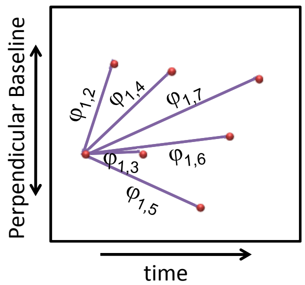


- Error on linear rate is independent of how network is connected (but of course short-baseline, short-time interferograms are best).
- To simplify mathematics, assume all connections to date d1...

...and regular acquisition spacing, t_r

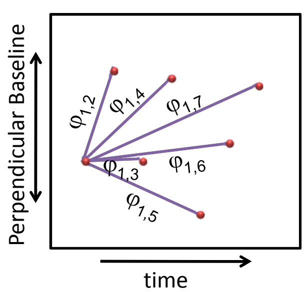
- We can determine the best-fit linear rate of phase change due to deformation, $\frac{d\varphi}{dt}$, using weighted least squares:

$$\boldsymbol{\Sigma}_{\mathbf{P}}^{-1}\mathbf{T}\frac{d\varphi}{dt} = \boldsymbol{\Sigma}_{\mathbf{P}}^{-1}\mathbf{P}$$

where **T** = $[t_r, 2t_r, \ldots Nt_r]^T$, **P** = $[\varphi_{1,2}, \varphi_{1,3}, \ldots \varphi_{1,N}]^T$, and $\boldsymbol{\Sigma}_{\mathbf{P}}^{-1}$ is the inverse of the variance-covariance matrix for the range change observations, **P**.

Error Budget (2)

Optimum determination of Linear Deformation Rates



- Using the correct VCM, $\boldsymbol{\Sigma}_{\mathbf{P}}$, is essential.
- In this particular network, all interferograms share a common acquisition (epoch 1).

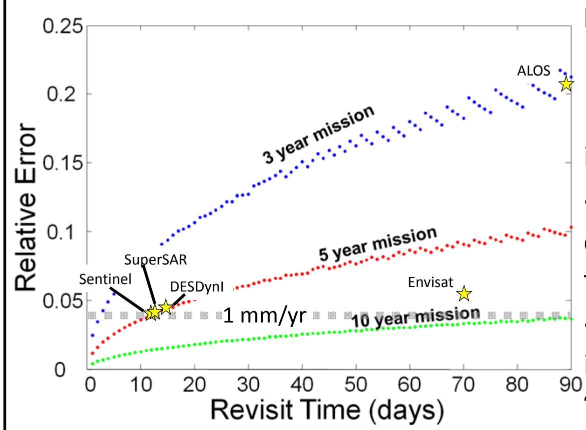
⇨ Cov $(\varphi_{1,i}, \varphi_{1,j}) = \sigma_1^2$ (the variance on epoch 1)

and Var $(\varphi_{1,i}) = \sigma_1^2 + \sigma_i^2$

$= 2\sigma^2$ (assuming noise is identical on all epochs)

Error Budget (2)

Optimum determination of Linear Deformation Rates



Error ∝ (*revisit time*)$^{0.5}$

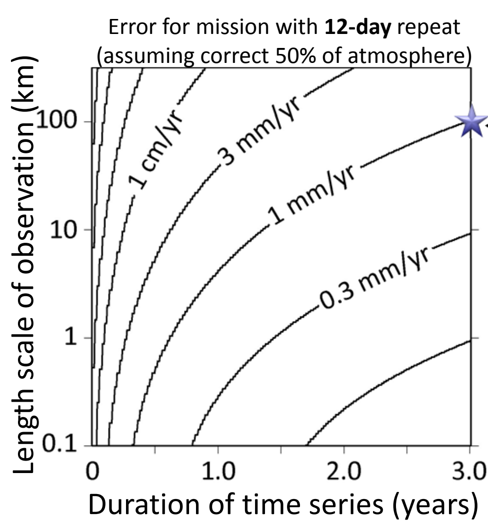
∝ (*mission length*)$^{-1.5}$

i.e.

- For a **fixed length mission**, cut revisit time by 4 to halve the linear rate error.
- For a **fixed revisit time**, increase mission length by ~60% to halve the linear rate error.

Error Budget (2)

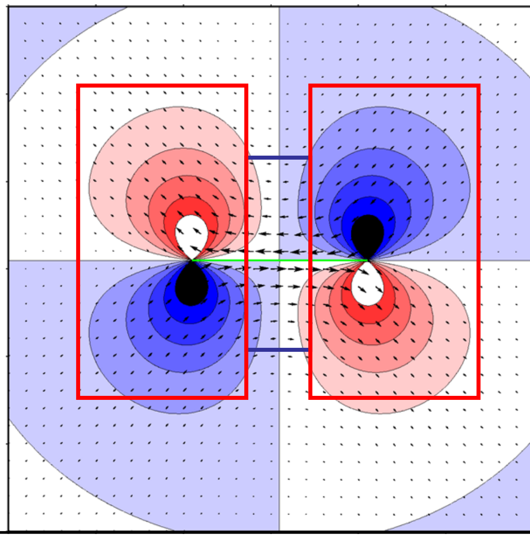
Optimum determination of Linear Deformation Rates



Reaching the target precision is tough!

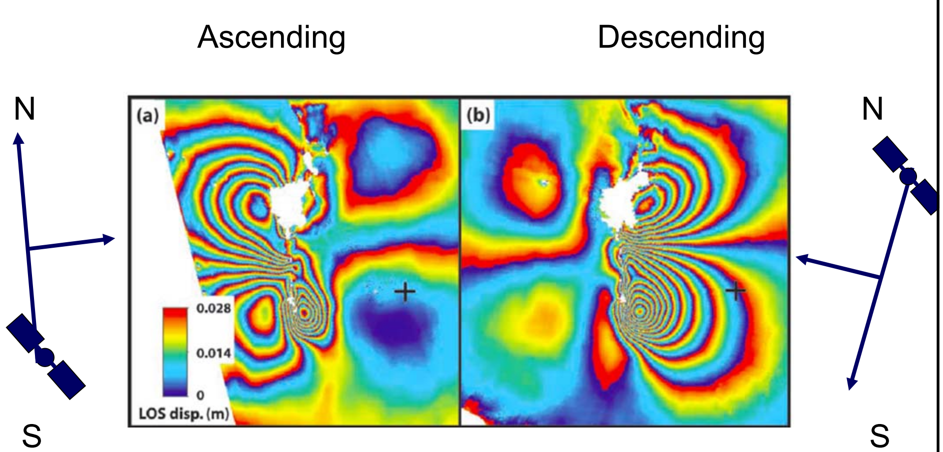
Everything so far has been for Line-of-sight deformation

Combining Viewing Geometries



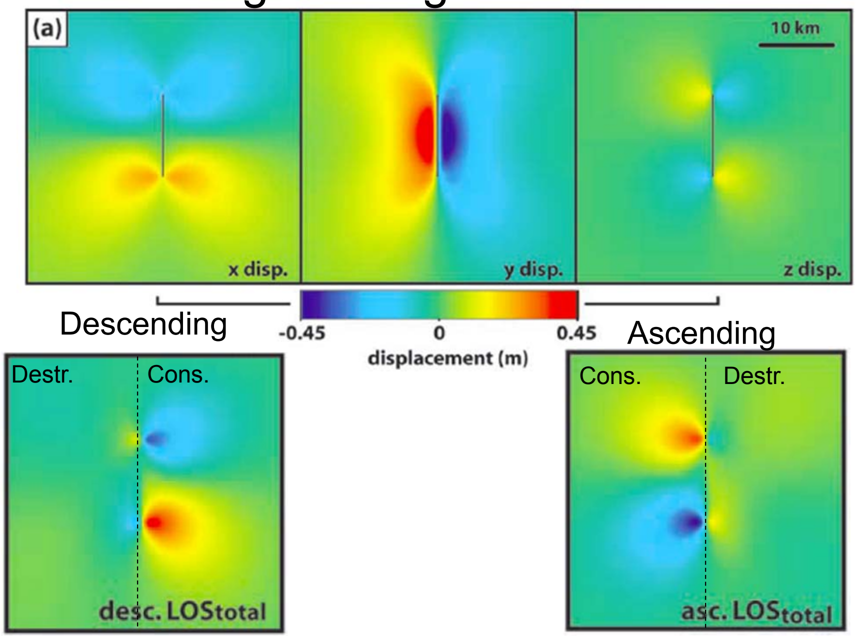
Surface Displacements of Strike Slip Faults

Combining Viewing Geometries

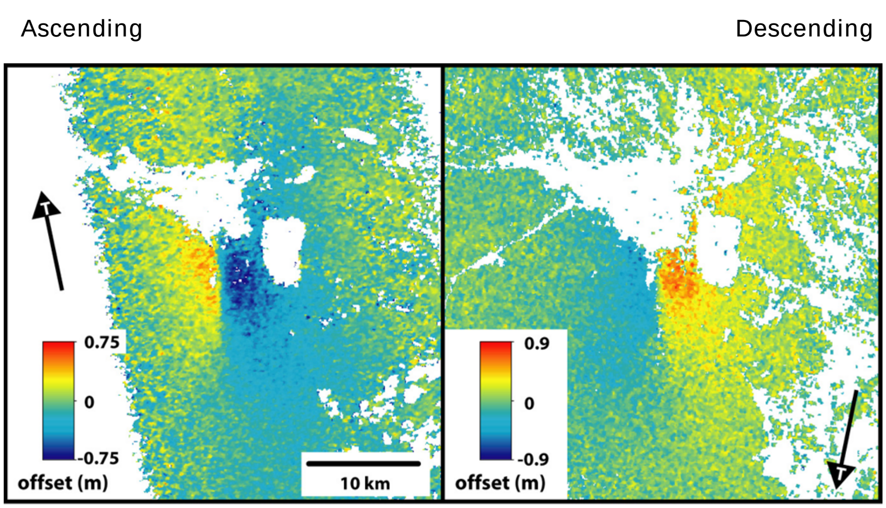


Bam, Earthquake, Iran, 2003. (Funning et al, 2006)

Combining Viewing Geometries

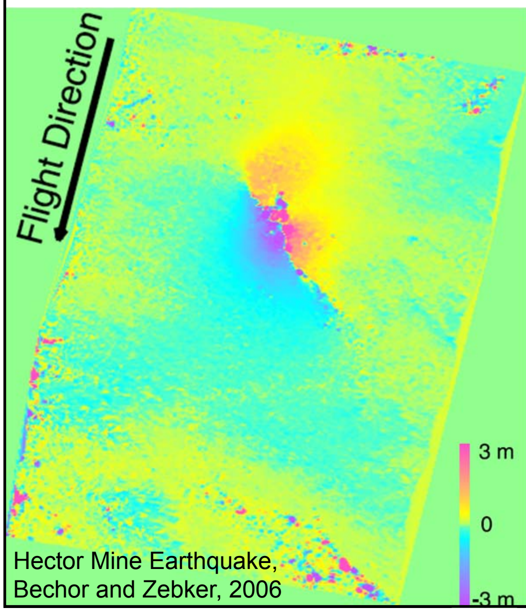


Azimuth offsets

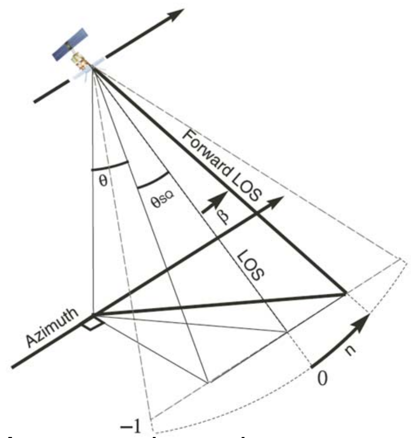


MAI: Split Beam Processing

Split beam into forward- and backward- looking sections to measure displacement in flight direction.



Hector Mine Earthquake, Bechor and Zebker, 2006



Accuracy depends on coherence and SNR. Up to 3 cm.

Determining 3D displacements

If the 3D displacement at a pixel is given by $\mathbf{u} = [u_x, u_y, u_z]$, then…

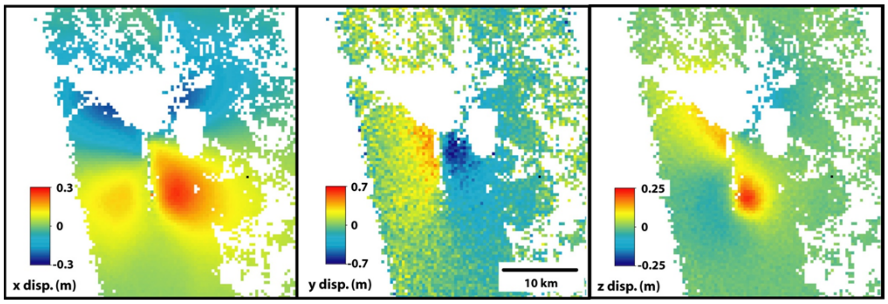
Ascending interferogram,	$d_1 = \mathbf{los_A} \bullet \mathbf{u}$
Descending interferogram,	$d_2 = \mathbf{los_D} \bullet \mathbf{u}$
Ascending az. offsets,	$d_3 = \mathbf{los_{AO}} \bullet \mathbf{u}$
Descending az. offsets,	$d_4 = \mathbf{los_{DO}} \bullet \mathbf{u}$

Which can be rewritten as a matrix equation, $\mathbf{d} = \mathbf{Lu}$, and solved for $\mathbf{u}$.

See e.g. Wright, T.J, B. Parsons, Z. Lu., Geophys Res. Lett. 30(18), p.1974, 2003

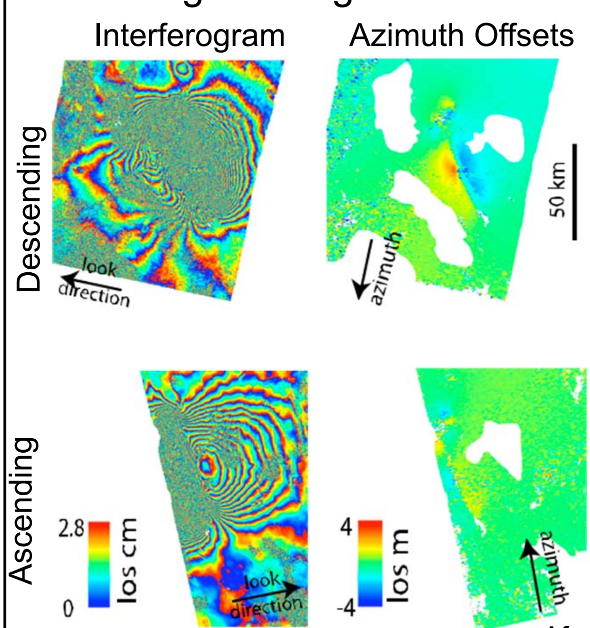
Bam earthquake 3D displacements

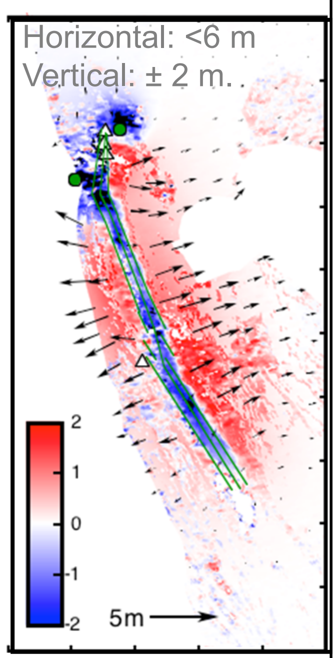
East North Up



σ = 0.01 m σ = 0.09 m σ = 0.01 m

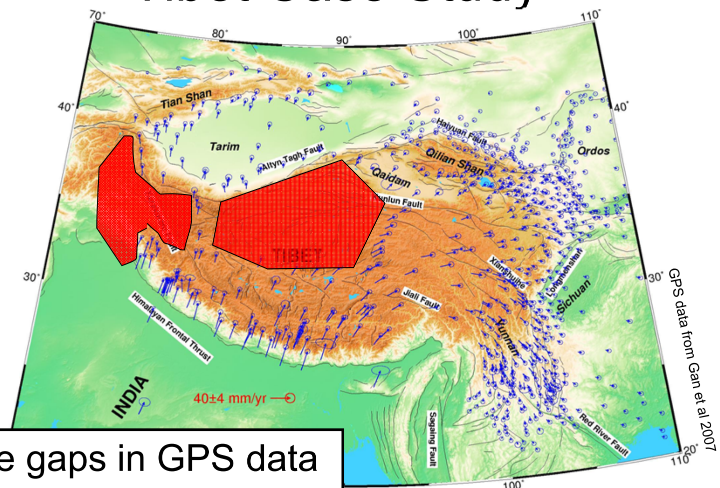
Combining Viewing Geometries





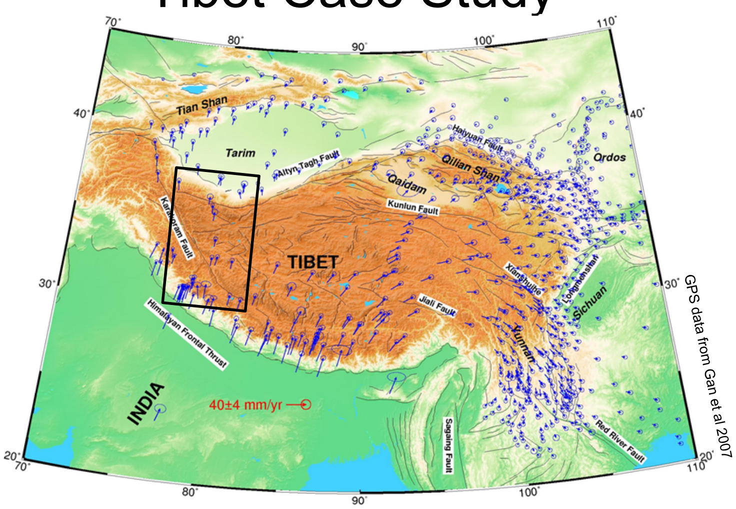
Afar Rift, Wright et al, Nature 2006

Tibet Case Study

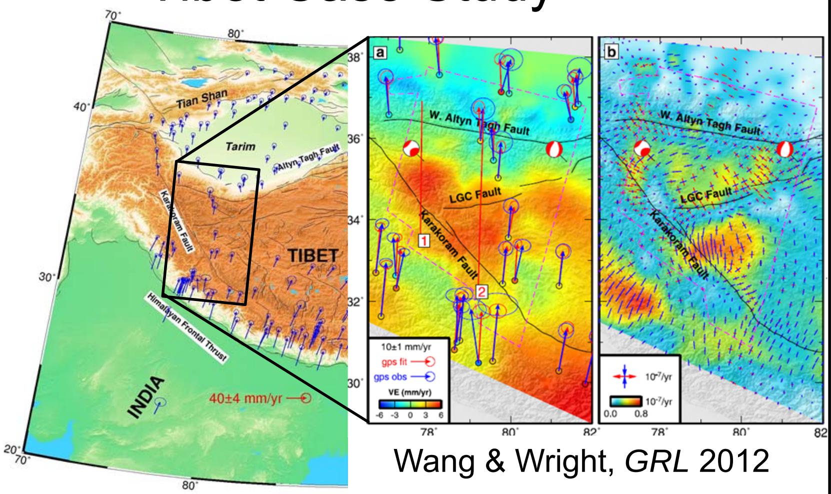


- Large gaps in GPS data
- Station spacing > 50 km

Tibet Case Study

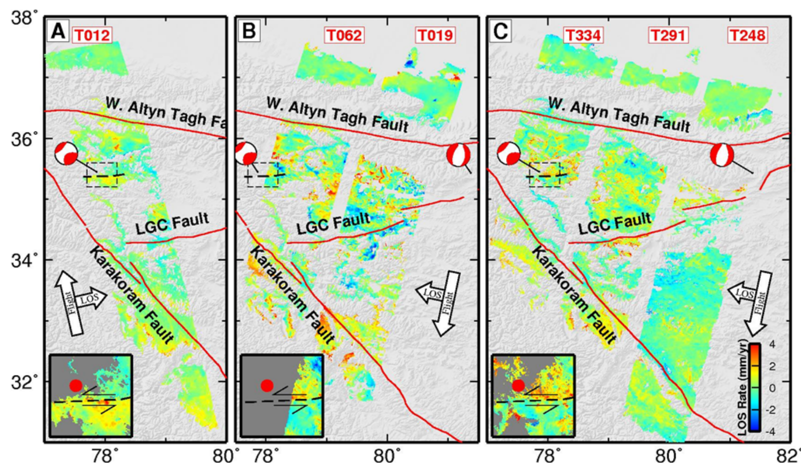


Tibet Case Study



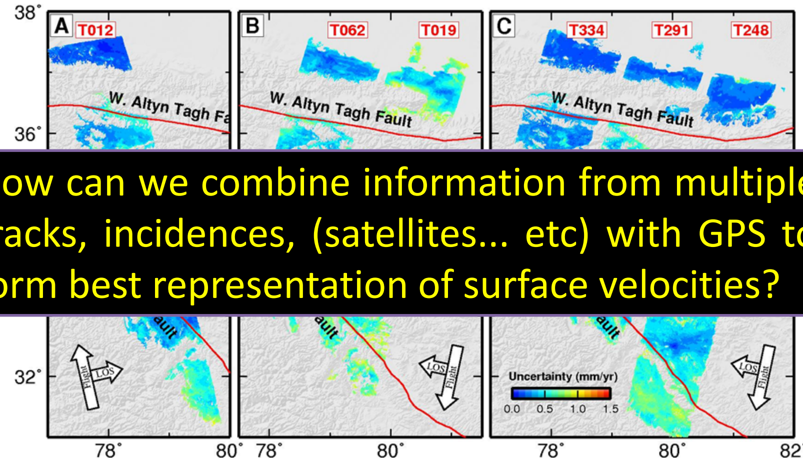
Wang & Wright, *GRL* 2012

InSAR Rate Maps from PI-RATE



RATE MAP = DEFORM + ORB + ATM + NOISE

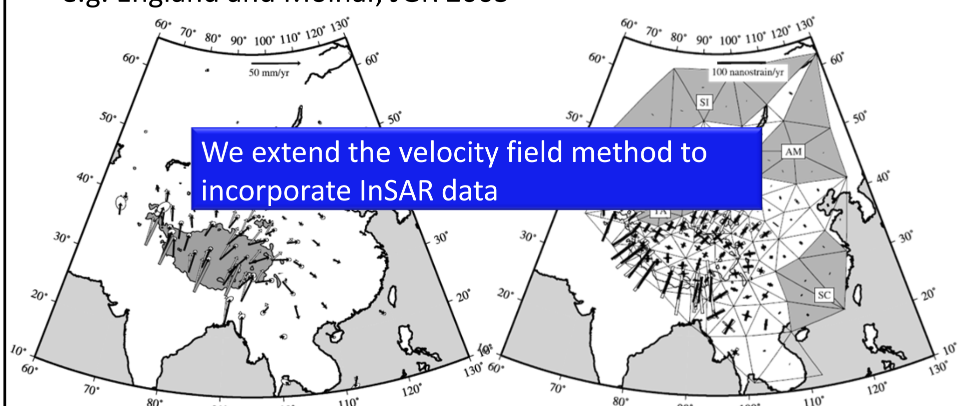
InSAR Error Maps from PI-RATE



How can we combine information from multiple tracks, incidences, (satellites... etc) with GPS to form best representation of surface velocities?

Velocity Field Method

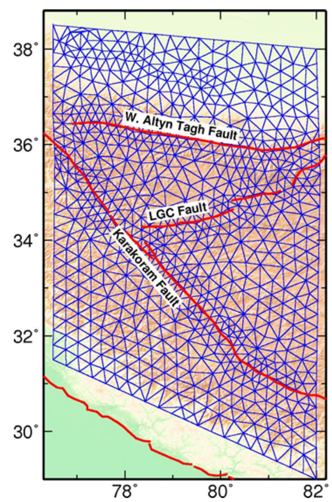
e.g. England and Molnar, JGR 2005

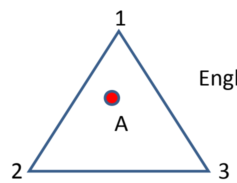


We extend the velocity field method to incorporate InSAR data

Velocities (left) and strain (right) from GPS, quaternary fault data and earthquake focal mechanisms

Velocity Field Method: Mesh and Interpolation





England & Molnar, 2005, JGR

[24] We divide the surface of the region of interest into spherical triangles and assume that within each triangle, the velocity varies linearly with latitude and longitude across the triangle. We may express the velocity in the interior of the triangle in terms of the velocities of its vertices:

$$\mathbf{U} = \sum_{m=1}^{3} N_m \mathbf{u}_m, \qquad (5)$$

where $\mathbf{u}_m$ is the velocity of vertex m and N_m are interpolation functions:

$$N_i = a_i + b_i\phi + c_i\theta, \qquad (6)$$

where ϕ is longitude and θ is latitude.

Wang and Wright, GRL 2012

Velocity Field Method: LS Solutions

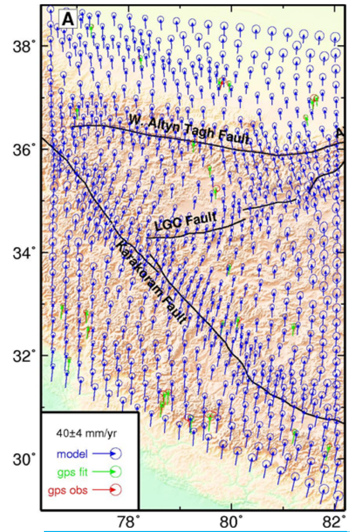
$$\begin{bmatrix} \mathbf{G}_{sar} & \mathbf{G}_{orb} & \mathbf{G}_{atm} \\ \mathbf{G}_{gps} & \mathbf{0} & \mathbf{0} \\ \kappa^2\nabla^2 & \mathbf{0} & \mathbf{0} \end{bmatrix} \begin{bmatrix} \mathbf{M}_{vel} \\ \mathbf{M}_{orb} \\ \mathbf{M}_{atm} \end{bmatrix} = \begin{bmatrix} \mathbf{d}_{sar} \\ \mathbf{d}_{gps} \\ \mathbf{0} \end{bmatrix}$$

Weighted LS solution: $\hat{\mathbf{M}} = (\mathbf{G}^{\mathrm{T}}\mathbf{W}\mathbf{G})^{-1}\mathbf{G}^{\mathrm{T}}\mathbf{W}\mathbf{d}$

Weighting by full data covariances

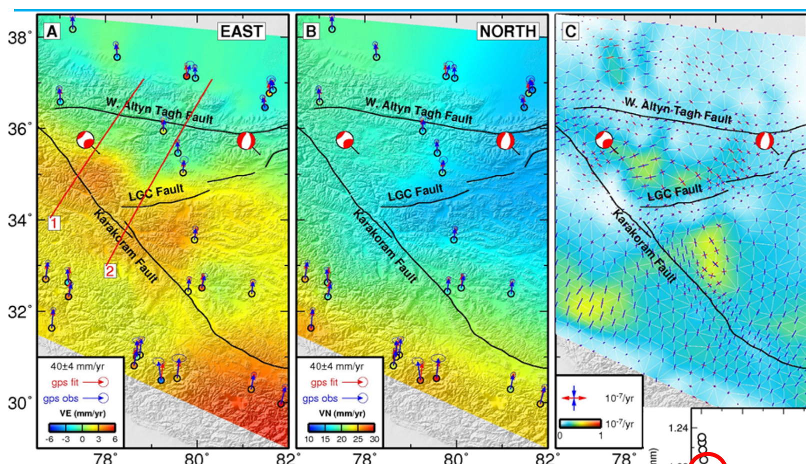
Wang and Wright, GRL 2012

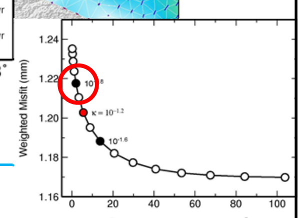
Velocity Field: From Vertices to Continuous



Wang and Wright, GRL 2012

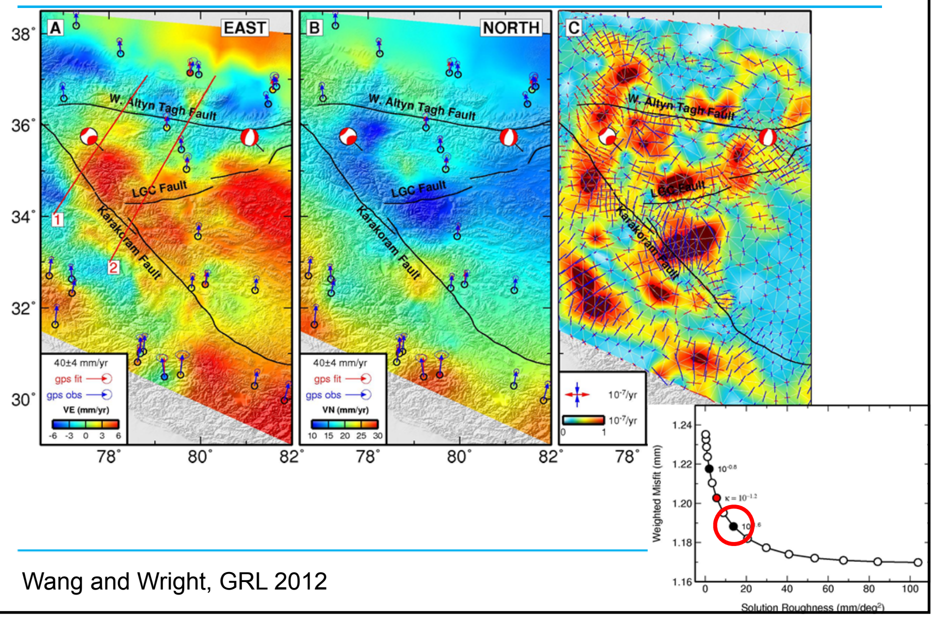
Laplacian Smoothing: Over-smoothed





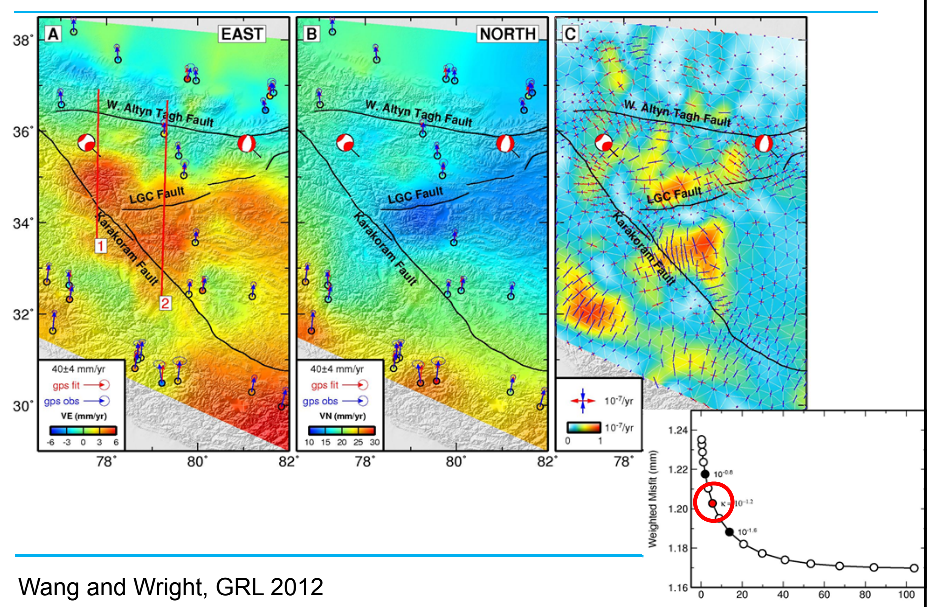
Wang and Wright, GRL 2012

Laplacian Smoothing: Little-smoothed



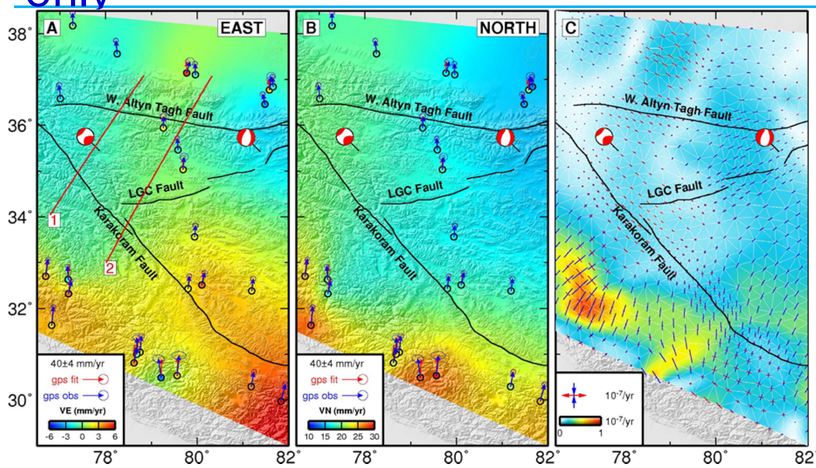
Wang and Wright, GRL 2012

Laplacian Smoothing: Best Solution



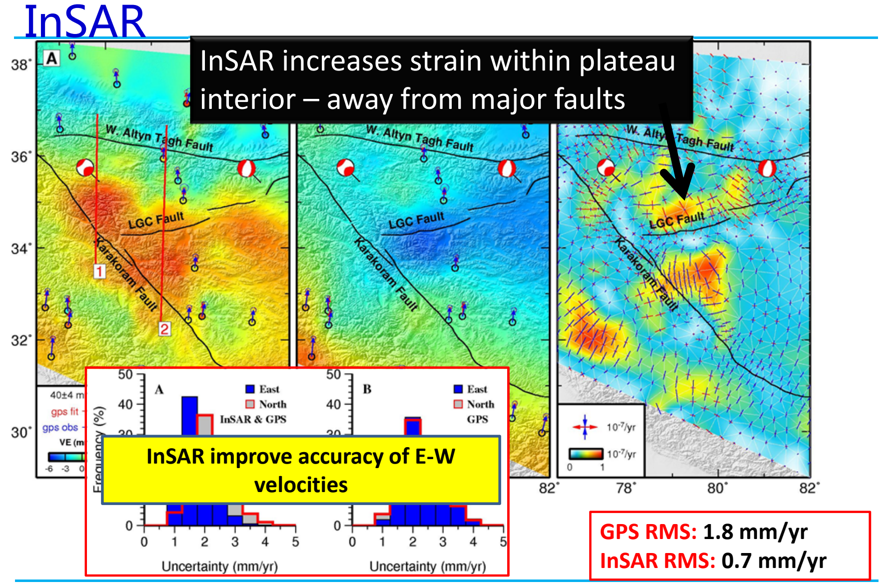
Wang and Wright, GRL 2012

Velocity & Strain Rate Field from GPS only



Wang and Wright, GRL 2012

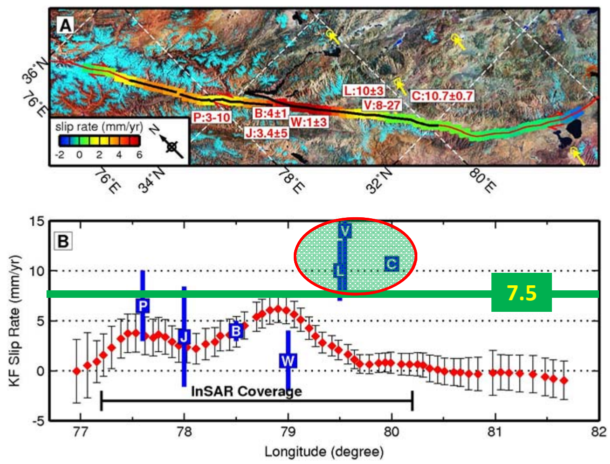
Velocity & Strain Rate Field from GPS & InSAR



Wang and Wright, GRL 2012

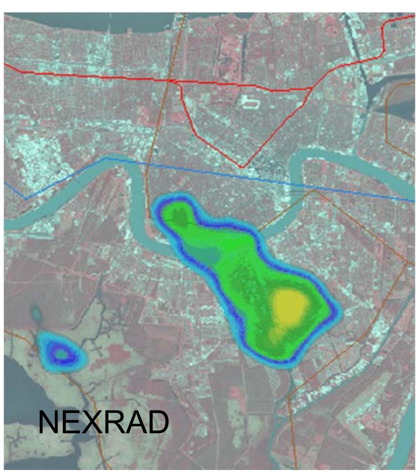
Slip Rates Along the Karakoram Fault

- ❑ Right-lateral slip along the entire fault
- ❑ Variable slip rate along the fault (0-6 mm/yr)
- ❑ Rule out present-day slip rates of >10 mm/yr
- ❑ No significant focused strain

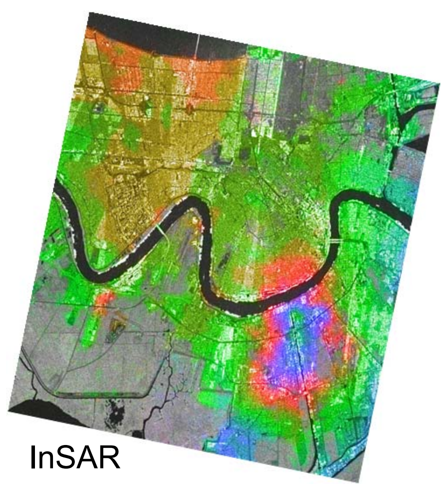


Limitation: Turbulent Atmosphere

Ground-based water vapour measurement

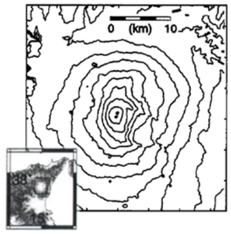


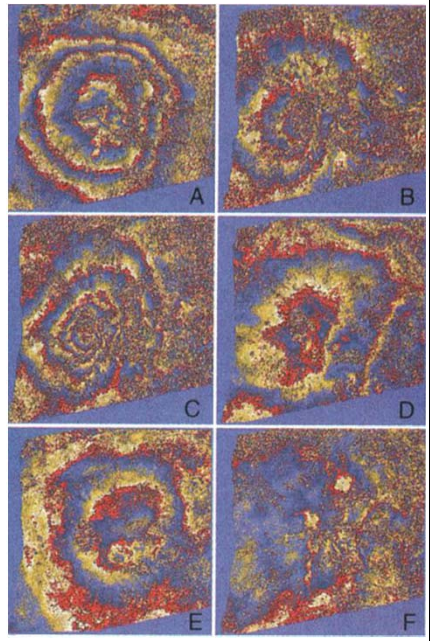
Interferogram

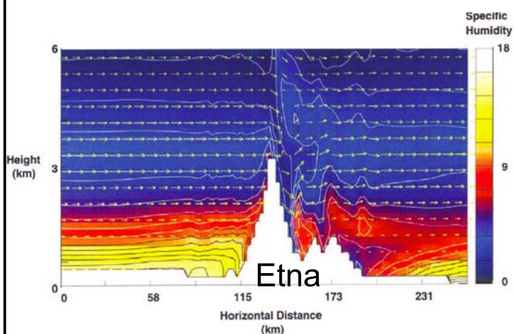


Limitation: Stratified Atmosphere

Mt Etna, Italy.



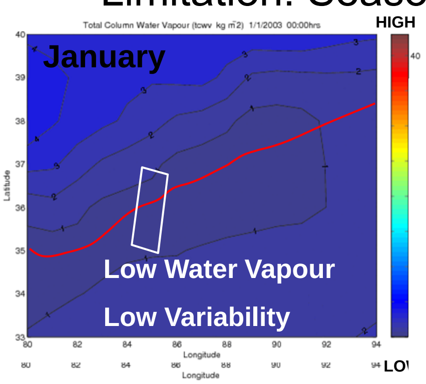


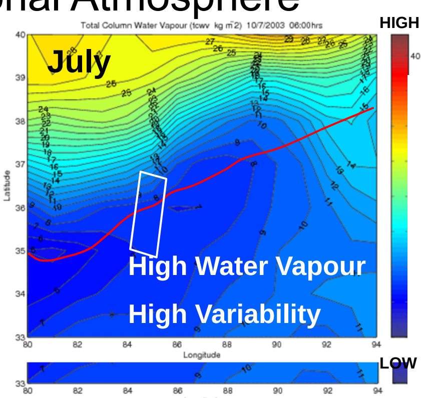


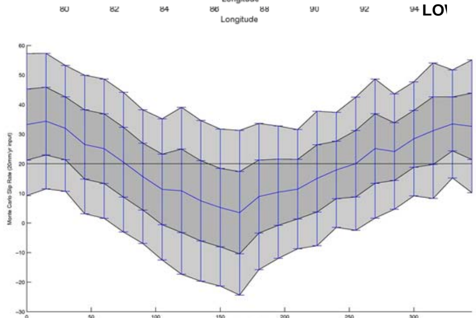
NH3D Model, Wadge 2002.

Limitation: Seasonal Atmosphere

Elliot, Fringe 07





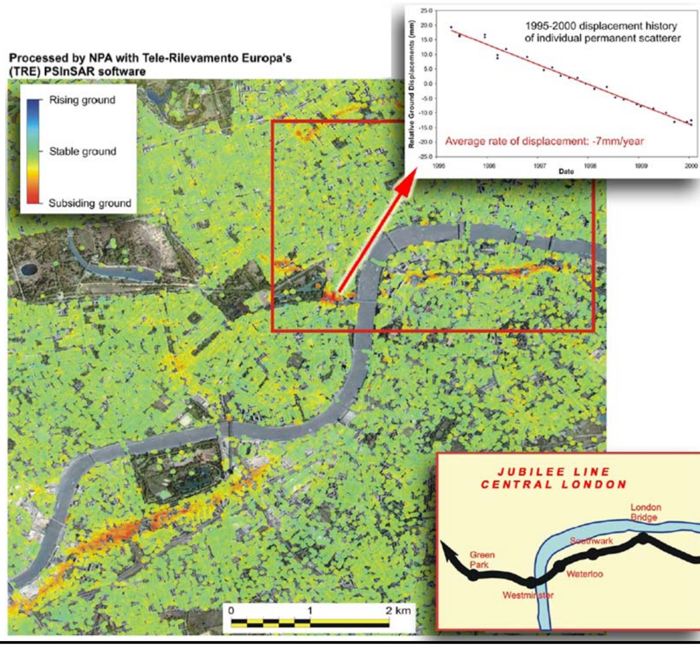


Synthetic Test of Rate Bias

Input Rate: 20 mm/yr

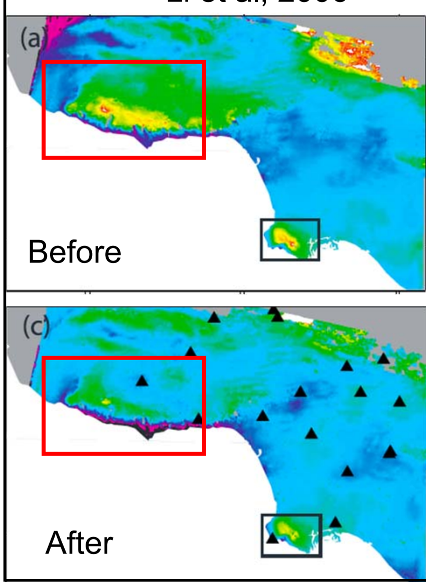
Recovered Rate: 5-35 mm/yr

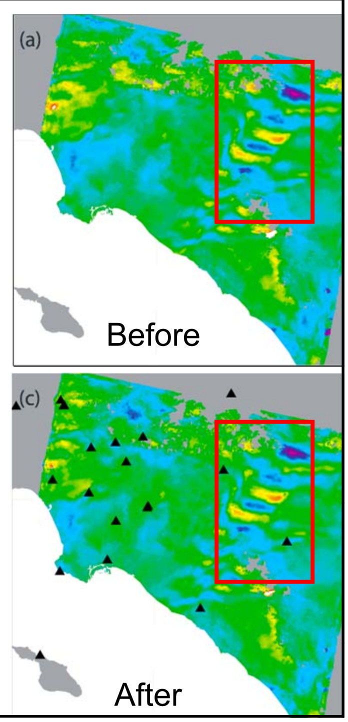
Corrections 1: Linear/Smooth Velocity Assumption



Correction 2: GPS

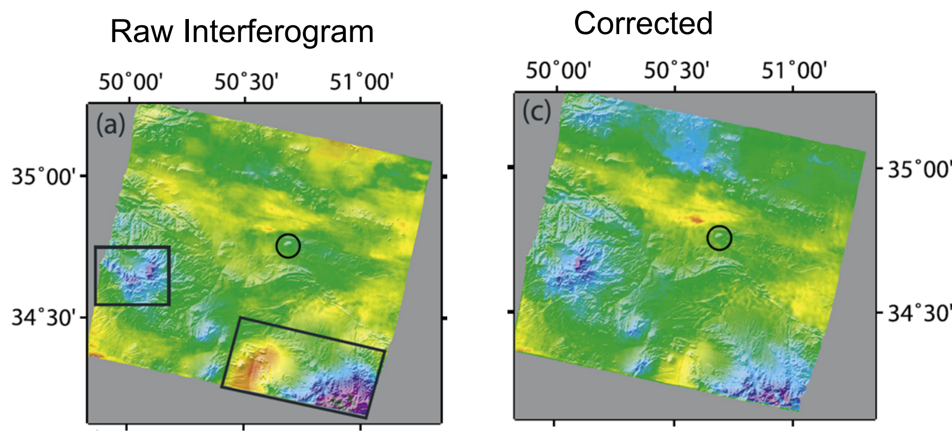
Requires dense GPS network
Li et al, 2006





Correction 3: MERIS (or MODIS)

Passive Optical/IR sensor on Envisat



Requires: descending orbit , daytime and cloud free conditions.

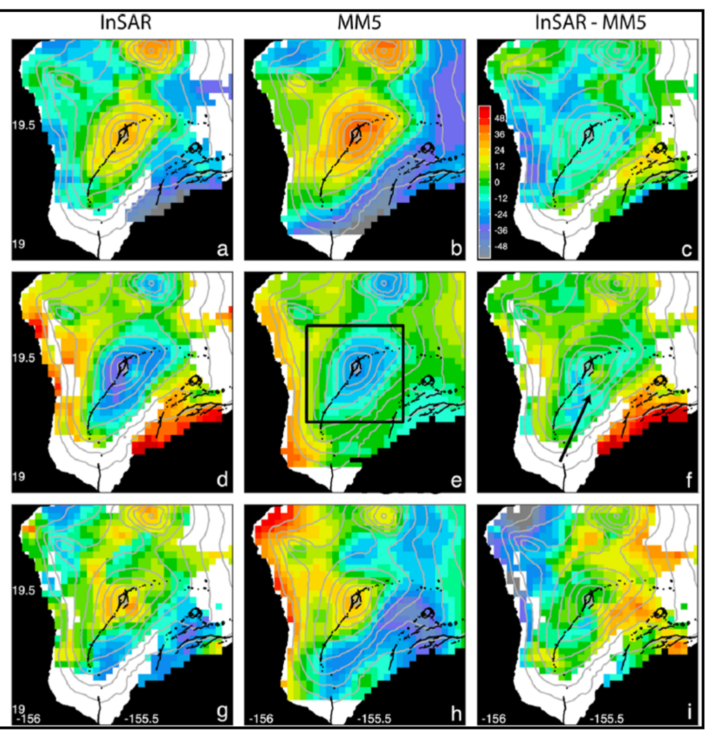
Li et al, 2006

Atmospheric Correction 4: Weather Model.

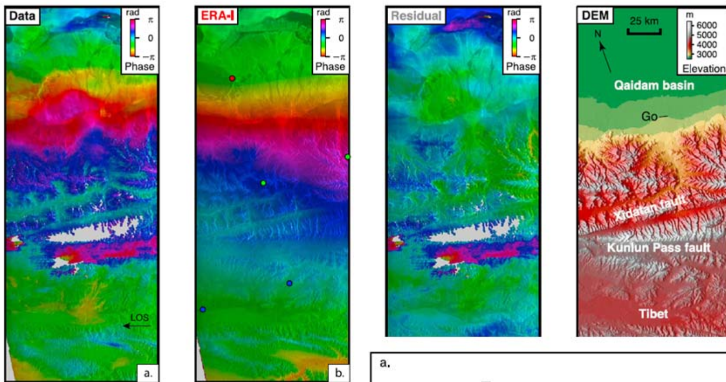
3 km resolution

Reduces long-wavelength (>30 km) effects but not smaller scale features.

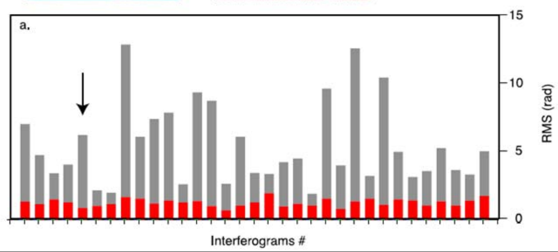
Foster et al, 2006



Atmospheric Correction 4: Weather Model (2)

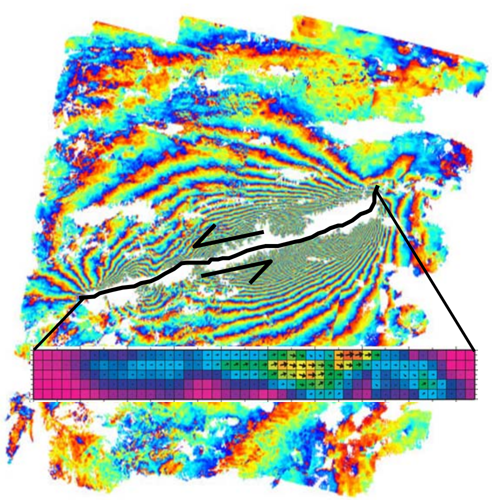


Jolivet et al., 2011



Earthquakes

1. Coseismic Deformation



Current Capability
• Map deformation fields for most damaging earthquakes.
• Identify responsible faults
• Estimate slip models.
• Assess impact on future hazard .

What could be done?
• Routine analysis of **ALL** damaging earthquakes, c.f. Harvard CMT.
• Real-time assessment of causative fault and likely damage area.
• Near-real time assessment of future hazard (aftershocks + triggered quakes).

Why are we not doing this already?
• Data.
• Method Development.
• Manpower.

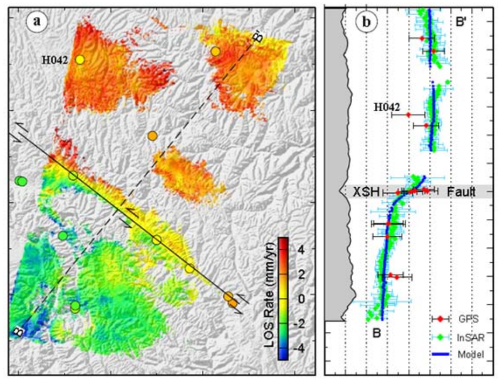
Earthquakes

2. Interseismic Strain

Current Capability
• Measure interseismic strain rates on suitable, targeted faults.
• Use these to constrain slip rate and hence assess future hazard.

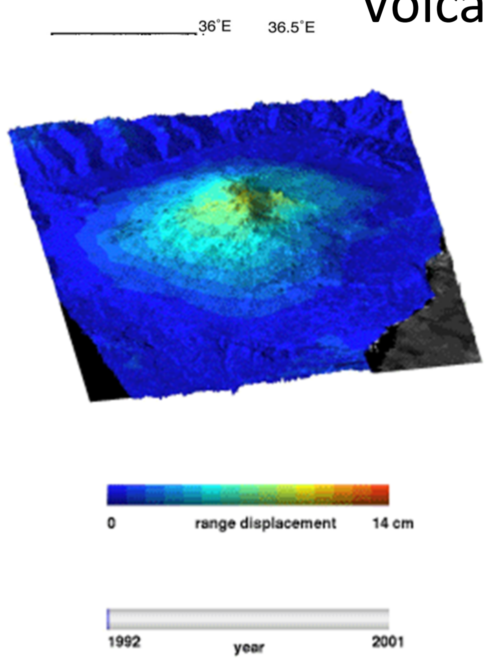
What could be done?
• Routine measurement of strain across whole regions.
• Assessment of slip rates and relative hazard of multiple faults (including unidentified faults).

Why are we not doing this already?
• Data.
• Method Development.
• Manpower.



Wang, Wright and Biggs., GRL 2009

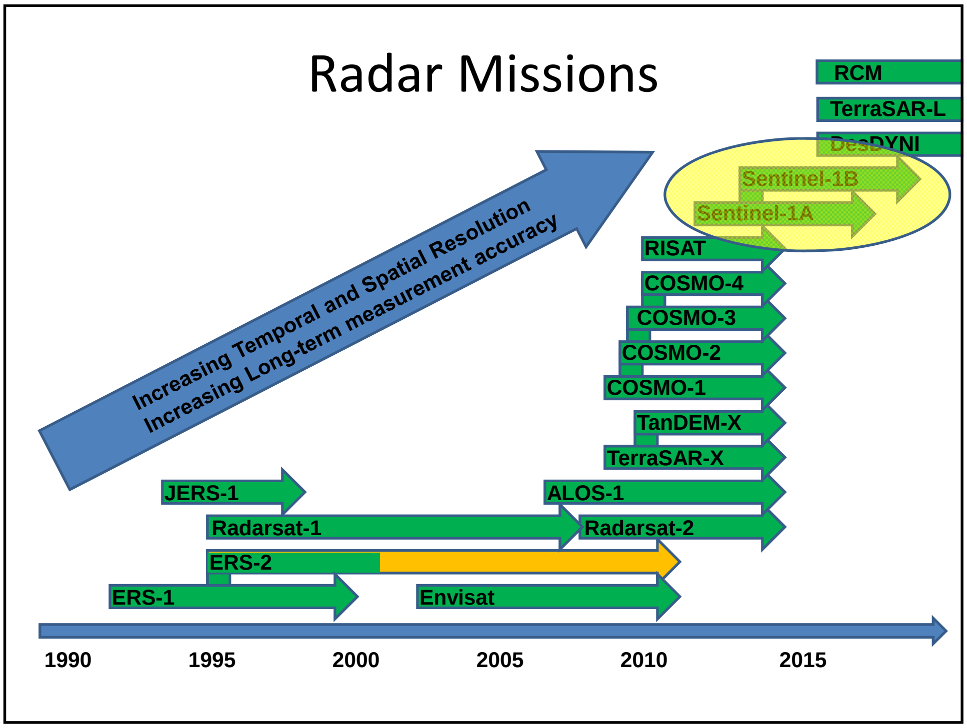
Volcanoes



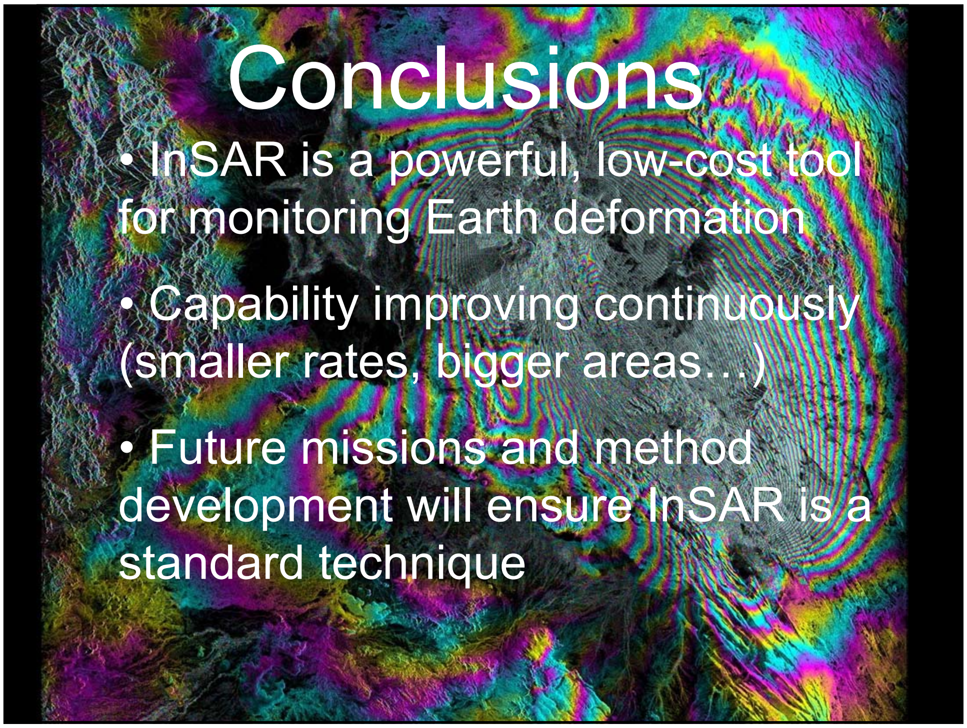
Current Capability
• Time-series analysis for suitable, targeted volcanoes .
• Snapshot regional surveys.
• Integration with other data sets.

What could be done?
• Routine monitoring of ALL volcanoes worldwide (or in a region).
• Target application of ground monitoring in countries where resources are limited.

Why are we not doing this already?
• Data.
• Method Development.
• Manpower.


Radar Missions
RCM
TerraSAR-L
DesDYNI
Sentinel-1B
Sentinel-1A
RISAT
COSMO-4
COSMO-3
COSMO-2
COSMO-1
TanDEM-X
TerraSAR-X
ALOS-1
Radarsat-2
JERS-1
Radarsat-1
ERS-2
ERS-1
Envisat
Increasing Temporal and Spatial Resolution
Increasing Long-term measurement accuracy
1990
1995
2000
2005
2010
2015


The Future
Sentinel-1 (ESA, GMES)
• “Operational” C-band InSAR
• 12 day repeat, 2 satellites ⇨ 3 day revisit
• Funded for 20 years, Launch early 2014
esa


Conclusions
• InSAR is a powerful, low-cost tool for monitoring Earth deformation
• Capability improving continuously (smaller rates, bigger areas…)
• Future missions and method development will ensure InSAR is a standard technique


SUPERVOLCANO
bbc.co.uk/supervolcano