

# Pandora Lunar Measurements





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General goal: Assess the possibilities of retrieving trace gas columns from lunar measurements by Pandora

In this WP (and beyond):

- Settle retrieval settings
- Test impact of lunar albedo correction
- Investigate potential limitations





#### All data shown in the following are from Rome, Pandora 117s1+s2



### 🕲 Lunar retrieval settings

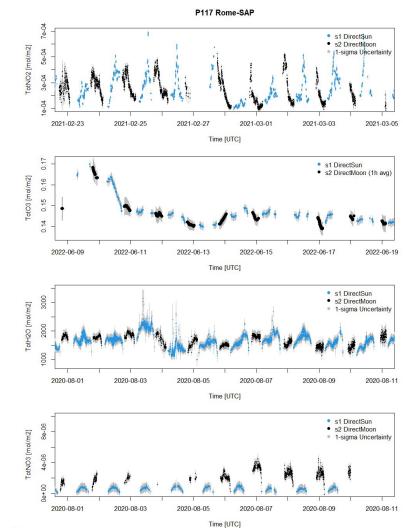
**NO2** (400-470 nm) separation between BL and strat NO2!

**03** (480-570 nm)

Lunar retrievals use solar reference spectra

H20 (490-510 nm)

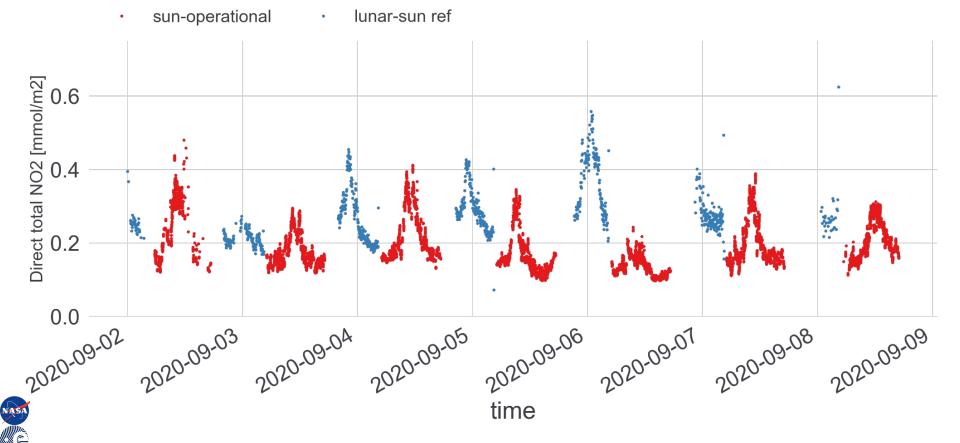
**NO3** (600-685 nm)



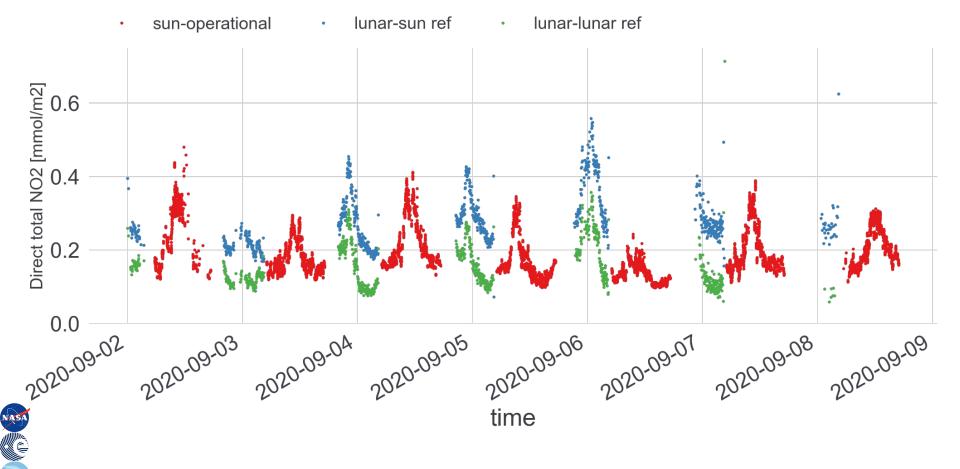
QA4EO workshop #4, Potsdam, 281



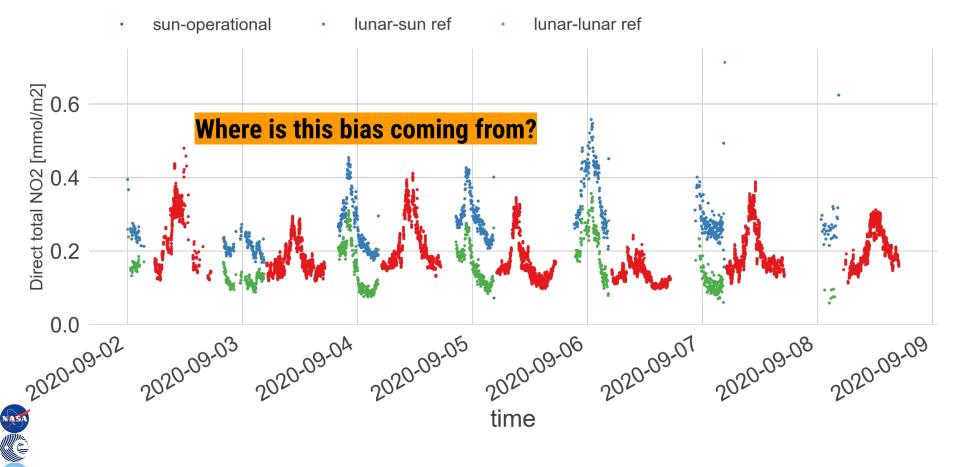
## 🐨 Impact of calibration source (solar or lunar spectrum)



### ( Impact of calibration source (solar or lunar spectrum)



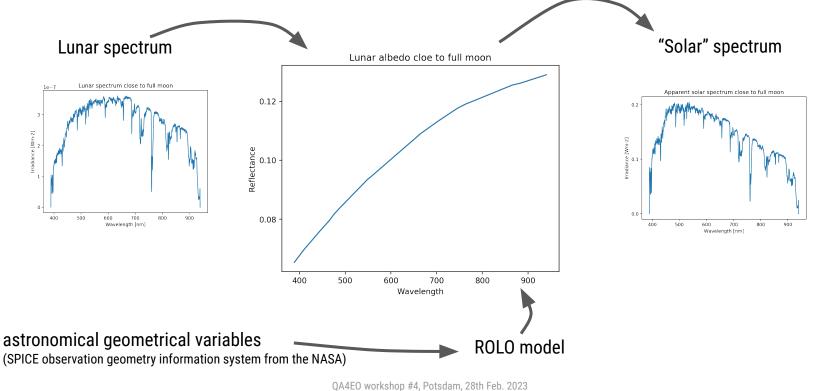
### ( Impact of calibration source (solar or lunar spectrum)



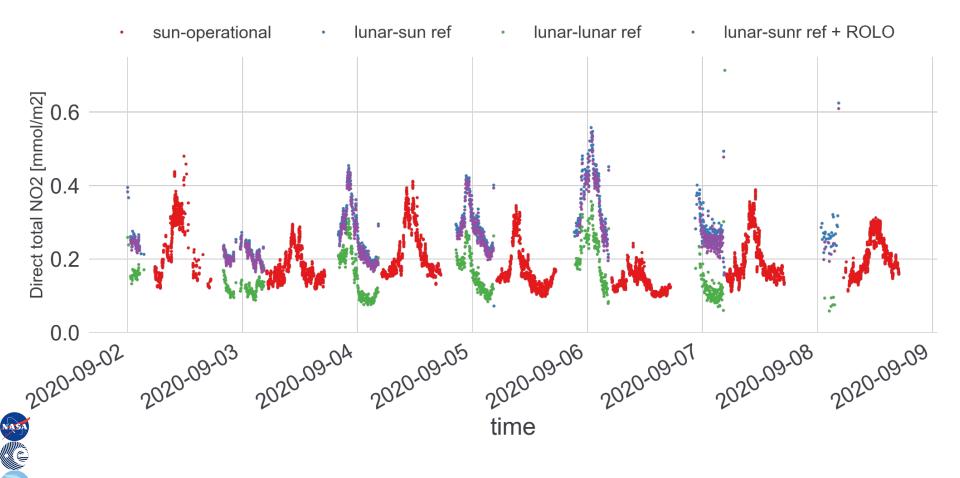
#### 🛞 Bias due to lunar spectral albedo?

Lunar spectral albedo correction applied using parameterization derived from RObotic Lunar Observatory (ROLO)

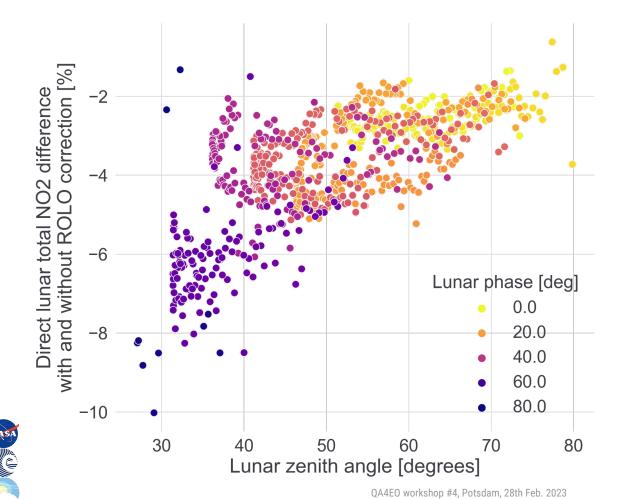
NAS



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Impact of correction higher for larger lunar phase, but generally small.

This confirms that smooth spectral features (line lunar albedo) are well captured by closure polynomials in the retrieval.

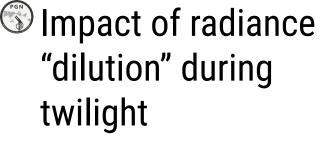
#### Not driver for bias !

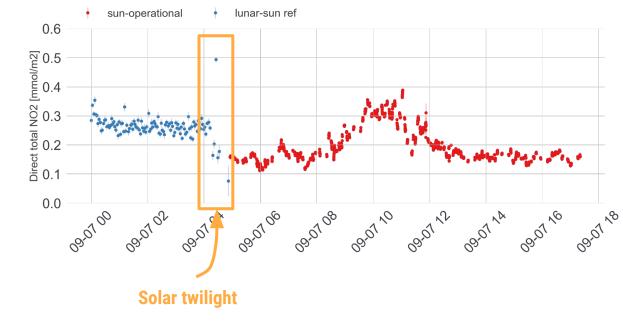
#### 🛞 Bias due to instrumental artifacts?

Since lunar spectral albedo correction does help much, **instrumental artifacts** seem to **drive the bias** !

The **source** of these artifacts has **not** been **identified yet**!

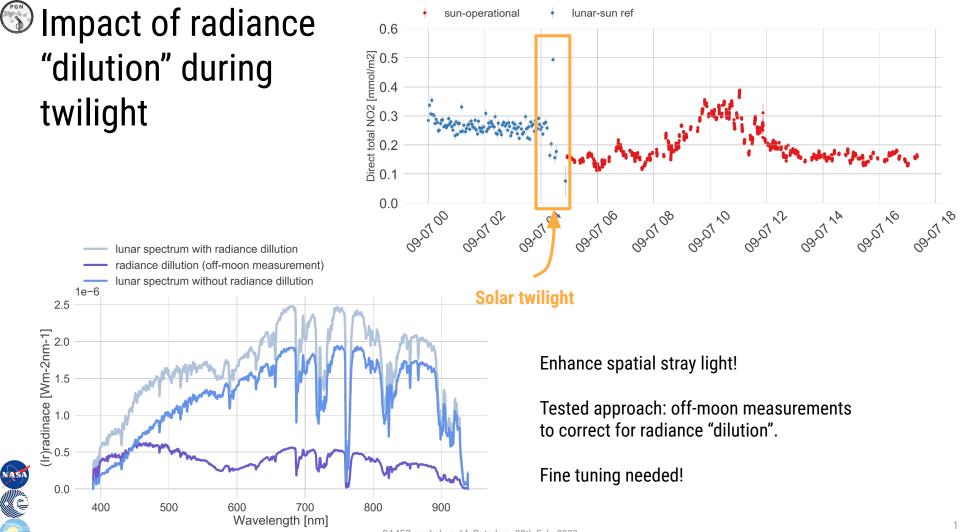




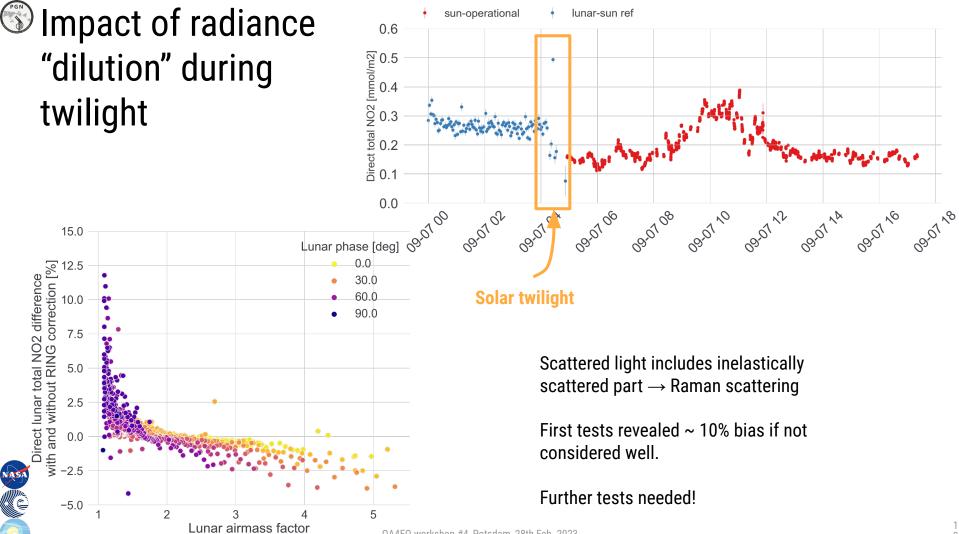




QA4EO workshop #4, Potsdam, 28th Feb. 2023



QA4E0 workshop #4, Potsdam, 28th Feb. 2023



QA4E0 workshop #4, Potsdam, 28th Feb. 2023

#### Suggested WP extensions

- The **source of the bias** (dependency on the calibration source) needs to be studied further. Is it a consequence of missing or improper instrument characterization? Note that this task would also impact other data products in the PGN in a positive way.
- **Radiance "dilution"** corrections during solar twilight will be developed. "Active" correction techniques, like radiance measurements recorded next to the moon and close in time, have been prototyped already, but would need further refinement. Feasibility studies towards "passive" correction techniques including modeling or aerosol proxies (like 0202 columns as utilized in MAX-DOAS) would be carried out.
- Investigate the improvement capability by considering the Raman effect under solar twilight conditions. Test in particular what gasses (respectively spectral regimes) are affected most and if molecular absorption (molecular Ring effect) needs to be considered for strong absorbers (e.g. H<sub>2</sub>O or O<sub>3</sub>).





## Thanks for your attention!



